

Using an Atom Interferometer to Infer Gravitational Entanglement Generation

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If gravitational perturbations are quantized into gravitons in analogy with the electromagnetic field and photons, the resulting graviton interactions should lead to an entangling interaction between massive objects. We suggest a test of this prediction. To do this, we introduce the concept of interactive quantum information sensing. This novel sensing protocol is tailored to provable verification of weak dynamical entanglement generation between a pair of systems. We show that this protocol is highly robust to typical thermal noise sources. Moreover, the sensitivity can be increased both using an initial thermal state and/or an initial phase of entangling via a nongravitational interaction. We outline a concrete implementation testing the ability of the gravitational field to generate entanglement between an atomic interferometer and a mechanical oscillator. Preliminary numerical estimates suggest that near-term devices could feasibly be used to perform the experiment.

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I. INTRODUCTION

If a particle is in a superposition of two locations, will its gravitational field also be in a superposition and can this field generate entanglement with another system? This foundational question [1,2] has received considerable attention [3–13]. Proposed experimental tests to detect entanglement due to gravity based on Bell tests (or, more generally, entanglement witnesses [14,15]) require performing measurements on both subsystems and are challenging in practice. As a result, there is still no direct experimental evidence as to whether gravitational interactions generate entanglement. Here, we propose a test that only requires observing a single subsystem [16–19]. We show that if an interaction (such as gravity) between two systems can cause both decoherence (collapse) and recoherence (revival) of a subsystem, then for restricted classes of systems the interaction is *necessarily* capable of generating entanglement. We propose a concrete implementation based on atom interferometry [20–23], in which an atom in a superposition of being in one of two interferometer arms interacts with a low-frequency mechanical

resonator [24,25] (see Fig. 1). The signal for entanglement generation is a collapse and revival of the atomic interference fringes due to the periodic motion of the resonator. The experiment does not require the preparation of a nonclassical state of the oscillator and can in fact be enhanced by placing the oscillator in a thermal state, which appears to make this experiment feasible with near-term devices.

The relation of such an experiment to the quantization of gravity is a subject of intense current study [26–29]. These experiments operate in a regime where the energy density (or, equivalently, the space-time curvature), is far below the Planck scale $\rho \ll m_{\text{Pl}}/\ell_{\text{Pl}}^3 \sim 10^{123}$ eV/cm³. Thus the nonlinearity of the gravitational interaction is very weak, and one can treat the metric $g_{\mu\nu}$ as a linear perturbation around flat space-time. In this limit, one can quantize the gravitational perturbations (“gravitons”) in exact analogy with quantum electrodynamics; graviton exchange generates a two-body Newton potential operator

$$V_N = -\frac{G_N m_1 m_2}{|\mathbf{x}_1 - \mathbf{x}_2|} \quad (1)$$

between a pair of masses, just as photons generate the Coulomb potential [30–35]. We review some standard demonstrations of this in Appendix A. In Eq. (1), $\mathbf{x}_{1,2}$ are the position operators on a pair of masses and thus this interaction can generate entanglement. However, there are dissenting opinions [4,36–38] about whether gravity should be quantized in this way and, indeed, one can produce models where classical gravitational interactions can

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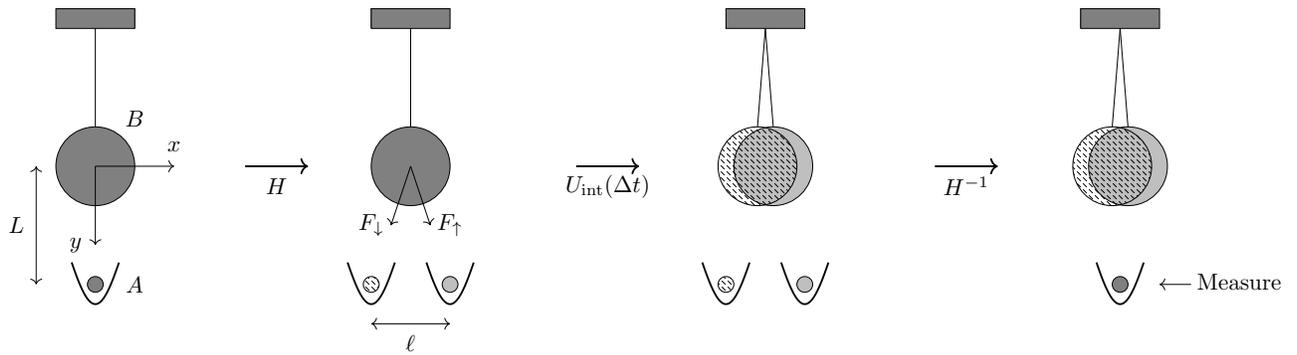


FIG. 1. Implementation of the basic protocol using an atom interferometer and a suspended pendulum (see Sec. VI). A trapped atom (labeled A) is prepared some distance L away from a mechanical resonator (B , here pictured as a pendulum). The atom is then put into a superposition of two different locations separated by ℓ , effecting a Hadamard gate H . This generates a state-dependent force between the atoms and resonator, leading to motion in opposite directions for some time Δt . Finally, the atom state is recombined using the inverse Hadamard gate and measured to check for decoherence caused by the atom-mechanical interaction. When the resonator undergoes a complete period of motion, its state no longer depends upon the atoms and coherence is recovered for the interferometer.

arise but without generating entanglement [39–42], providing substantial motivation to perform tests of Eq. (1).

The ability to test such a weak entanglement signal relies entirely on our central technical result, a novel sensing protocol that we refer to as *interactive quantum information sensing*. This is a detection scheme tailored specifically to the verification of weak dynamical entanglement generation. The traditional methods to detect entanglement in bipartite systems $H_A \otimes H_B$ use nonlocal measurements [14,15] and can be very difficult in practice with noisy systems and weak entanglement. However, in the past two decades, more sophisticated methods have been developed to address these types of problems [43,44]. We suggest here a new protocol that relies on time-dependent measurements on a single subsystem. Within standard quantum mechanics, system A will decohere—evolve from a pure to mixed state—if it becomes entangled with another system B that is not measured [16–19]. This loss of coherence can be observed via an interference measurement on A alone. Simple decoherence could be explained by entanglement but also by, for example, random classical noise [45]. However, if the same interaction can cause both decoherence and recoherence of A , in a manner controlled by B , then for certain classes of systems, we prove that the interaction is necessarily capable of generating entanglement between subsystems A and B . This protocol provides an indirect test of the quantum communication capabilities of the two systems and is a limited probe of the family of quantum channels associated with the interaction between the two systems. The interplay between the information-theoretic channel properties and the physical interaction provides our suggested nomenclature.

We outline the interactive sensing protocol in Secs. II and III. We find the remarkable result that using an initial state at high temperature can *increase* the sensitivity of the protocol, because it can increase the rate of entanglement

generation and lead to a thermally enhanced collapse-and-revival signal. In Sec. IV, we demonstrate that this conclusion is robust to typical sources of noise, essentially because the test does not involve producing large superpositions of the nonobserved subsystem. In Sec. V, we show how to further enhance the protocol using pre-entangled initial conditions. Finally, we outline an experimental realization with gravitational entanglement generation between an atom interferometer and a mechanical oscillator in Sec. VI, before concluding with a discussion of implications and loopholes in Sec. VII.

II. COLLAPSE-AND-REVIVAL DYNAMICS

To begin, we illustrate the basic idea of the collapse-and-revival dynamics with an example. The setup is similar to electron spin-echo envelope modulation [46,47] and the cavity QED experiments of Haroche *et al.* [48]. Consider a harmonic oscillator B coupled to a two-state system A through the Hamiltonian

$$H = \omega a^\dagger a + g(a + a^\dagger)\sigma_z. \quad (2)$$

In Sec. VI, we give an implementation of this Hamiltonian where the oscillator B is a mechanical resonator, the two-state system A corresponds to an atom located in one of two spatial locations, and $g \ll \omega$ is set by the atom-oscillator gravitational interaction [Eq. (1)], so that g is proportional to Newton’s constant G_N . The essential idea is to do an interferometry measurement on the two-state system A (the “control”) in the presence of system B (the “target”). The key is the dynamical response of the target system B to a superposition of A .

To understand the entanglement dynamics generated by Eq. (2), it is useful to note that the time-evolution operator

can be rewritten

$$U(t) = e^{-iHt} = D^\dagger(\sigma_z \lambda) e^{-i\omega a^\dagger a t} D(\sigma_z \lambda) \quad (3)$$

up to an overall phase, where

$$D(\alpha) \equiv \exp\{\alpha a^\dagger - \alpha^* a\} \quad (4)$$

is the usual displacement operator [49]. Here and throughout, we use the dimensionless quantity

$$\lambda \equiv \frac{g}{\omega}. \quad (5)$$

This is the length, measured in units of the zero-point length x_0 , through which the oscillator equilibrium is displaced under the force from the atom. This ratio sets the scale of all observables considered in this paper.

Observation of the collapse-and-revival can be done with a typical interferometric measurement. Consider starting the full system in its decoupled ground state $|0\rangle_A \otimes |0\rangle_B$. The interferometry experiment then proceeds by performing a Hadamard gate (or any other beam-splitter operation) on the two-state system A , $|0\rangle \rightarrow (|0\rangle + |1\rangle)/\sqrt{2}$, evolving the joint system for some time t , performing the inverse Hadamard gate to recombine the two-level system, and then measuring its population. Mathematically, this proceeds as follows:

$$\begin{aligned} |\psi\rangle &= |0\rangle_A \otimes |0\rangle_B \\ &\xrightarrow{H} \frac{|0\rangle_A + |1\rangle_A}{\sqrt{2}} \otimes |0\rangle_B \\ &\xrightarrow{U_{\text{int}}} \frac{|0\rangle_A |\delta\rangle_B + |1\rangle_A |-\delta\rangle_B}{\sqrt{2}} \\ &\xrightarrow{H^\dagger} |0\rangle_A \frac{|\delta\rangle_B + |-\delta\rangle_B}{2} + |1\rangle_A \frac{|\delta\rangle_B - |-\delta\rangle_B}{2}. \end{aligned} \quad (6)$$

Here, the conditionally evolved states of the oscillator are simply coherent states

$$|\pm\delta\rangle_B = D[\pm\lambda(e^{-i\omega t} - 1)]|0\rangle. \quad (7)$$

If we now measure the two-state system A , we find for example that the probability of being in the $|0\rangle$ state is

$$P_A(0) = \frac{1}{2} + \frac{1}{2} \text{Re} \langle \delta | -\delta \rangle_B = \frac{1}{2} \left(1 + e^{-8\lambda^2 \sin^2(\omega t/2)} \right). \quad (8)$$

We see that the interference term is reduced, with a period set by the oscillator frequency ω . In particular, at half period** we have a maximum reduction of the phase contrast and after a full period the contrast is completely restored, as in Fig. 2.

Before moving on, we mention for later use an alternative calculation of the same effect. Consider the Pauli lowering operator $\sigma_- = (\sigma_x - i\sigma_y)/2$ on the two-level system. The expectation value $\langle \sigma_-(t) \rangle$ tracks the loss of phase contrast; we refer to the absolute value as the interferometric visibility $V = |\langle \sigma_- \rangle|$. Using the time-evolution operator given in Eq. (3), we have

$$\begin{aligned} \sigma_-(t) &= U^\dagger(t) \sigma_- U(t) \\ &= D^\dagger(-\lambda) e^{i\omega a^\dagger a t} D(-\lambda) \sigma_- D^\dagger(\lambda) e^{-i\omega a^\dagger a t} D(\lambda) \\ &= \sigma_- D[2\lambda(1 - e^{i\omega t})]. \end{aligned} \quad (9)$$

This is easy to show by working with explicit components in the σ_z basis, where $\sigma_- = |1\rangle\langle 0|$. With an oscillator

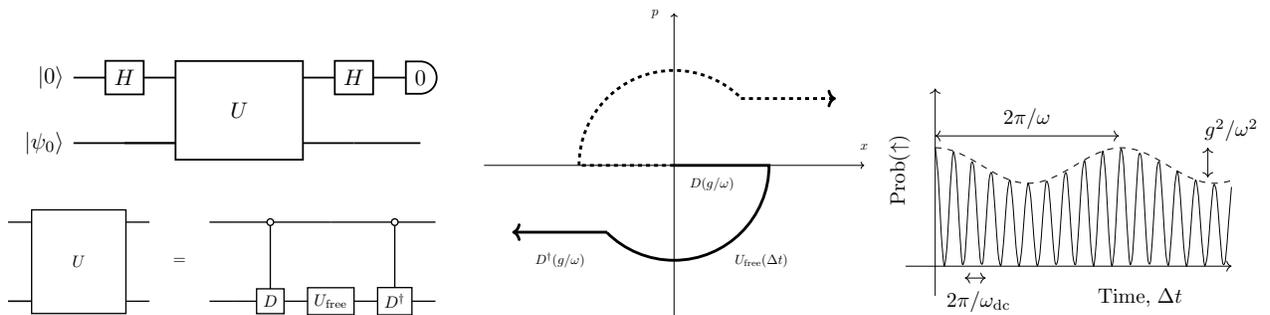


FIG. 2. Equivalent-circuit (left) and phase-space (center) descriptions of the experiment and schematic interferometric data output (right). In the circuit, the top line represents the atom and the bottom line the resonator. The large box represents joint evolution of the trapped atom and the resonator, which can be decomposed into conditional displacement of the resonator, followed by free evolution and an inverse displacement operator. This sequence can be visualized in the phase space of the oscillator, where the solid and dashed lines represent the two oscillator evolutions conditioned on the two possible atomic locations. Interferometric measurement of the atom population will show rapid fringes with frequency ω_{dc} due to any stray dc accelerations (e.g., due to electric fields, Earth's gravity, or off-center location of the resonator or atom), modulated by an overall reduction and then increase due to the atom-resonator entanglement. Resonator motion over a full period leads to nominal full recovery of the fringes.

initially in the ground state, this gives

$$\langle \sigma_-(\pi/\omega) \rangle = \langle \sigma_-(0) \rangle e^{-8\lambda^2}, \quad \langle \sigma_-(2\pi/\omega) \rangle = \langle \sigma_-(0) \rangle. \quad (10)$$

Here, we see again the loss of phase contrast at half period followed by the revival at a full period.

Up to this point, we assume that the oscillator is initialized in its ground state $|0\rangle$. In a realistic implementation—particularly one where the oscillator is a massive mechanical object—the oscillator instead starts in a mixed state, such as a thermal state, due to its coupling to an environment. Although one may be concerned that this would destroy the revival of coherence in the atom, it turns out that not only does the revival persist but, in fact, the relative contrast between decoherence and revival is *enhanced* so long as the thermalization time scale remains very long. That the revival persists is a consequence of the harmonic potential: after a full period, the state of the oscillator must return to its initial condition.

To see this, consider first the oscillator initialized to an arbitrary coherent state $|\alpha\rangle$. Using Eq. (9), we have

$$\begin{aligned} \langle \alpha | \sigma_-(t) | \alpha \rangle &= e^{-2\lambda[\alpha^*(1-e^{i\omega t}) - \alpha(1-e^{-i\omega t})]} \\ &\times e^{-8\lambda^2 \sin^2(\omega t/2)} \langle \sigma_-(0) \rangle. \end{aligned} \quad (11)$$

We see the complete revival after a full period, while at half period we now pick up a phase involving the initial oscillator momentum $p_\alpha = \alpha + \alpha^*$. To obtain the thermal-state result, one can now average over the coherent states [i.e., use the oscillator density matrix $\rho_{\text{th}} = \int d^2\alpha e^{-|\alpha|^2/\bar{n}}/(\pi\bar{n}) |\alpha\rangle\langle\alpha|$, with \bar{n} the thermal-phonon occupancy]. The result for the qubit visibility is

$$V_{\text{th}}(t) = \exp[-8\lambda^2(2\bar{n} + 1)\sin^2(\omega t/2)]. \quad (12)$$

In particular, we have $V_{\text{th}}(2\pi/\omega) = 1$, showing a full revival of the qubit coherence after a full oscillator period. On the other hand, at half period, we have $V_{\text{th}}(\pi/\omega) = \exp[-8\lambda^2(2\bar{n} + 1)]$, an enhancement to the loss of visibility by a factor of \bar{n} . Thus, starting with a thermal state increases the contrast between the “dip” of coherence halfway through oscillation and the recovery at full oscillation. The experiment is *easier* with a hot oscillator.

III. REVIVAL VERIFIES ENTANGLEMENT GENERATION

As this example clearly shows, entanglement generation between two systems A and B can cause periodic collapse and revival of the wave function of A . The crucial question is then: does observation of this collapse and revival *necessarily* require entanglement generation between A and B ? Our central result says that the answer is yes, under

some particular assumptions. We characterize this with a theorem:

Theorem 1. *Let L be a channel on $H_A \otimes H_B$, where H_A is a two-state system and H_B is arbitrary. Assume that:*

- (a) *The channel L generates time evolution, in a manner consistent with time-translation invariance, thus obeying a semigroup composition law $L_{t \rightarrow t'} = L_{t \rightarrow t''} L_{t'' \rightarrow t'}$ for all $t \leq t' \leq t''$.*
- (b) *The two-level subsystem H_A has its populations preserved under the time evolution, $\sigma_z(t) = \sigma_z(0)$.*
- (c) *L is a separable channel [50]: all of its Krauss operators are simple products. In particular, this means that any initial separable (nonentangled) state evolves to a separable state: $\rho(t) = L_t[\rho(0)]$ is separable for all separable initial states $\rho(0)$.*

Then the visibility $V(t) = |\langle \sigma_-(t) \rangle|$ is a monotonic function of time.

Here, we model the time evolution of the A - B system as a quantum channel L , a map on density matrices $\rho(t) = L_t[\rho(0)]$. For example, within standard quantum mechanics, the unitary evolution of the universe (A , B and their environment C , including the experimentalist) generates such a channel for the reduced A - B evolution. Suppose that we can experimentally convince ourselves that time-translation invariance in the form (a) and population condition (b) hold. Then the theorem says that if L cannot generate entanglement (c), then the only possible evolution for the qubit A is to have its interferometric visibility decay monotonically. Thus if we observe nonmonotonic visibility such as the oscillatory signal described above, we can conclude that the channel must be capable of generating entanglement.

We note that nonentangling channels still allow for nontrivial interactions. For example, semiclassical gravity $G_{\mu\nu} = 8\pi \langle T_{\mu\nu} \rangle$ (appropriately completed by a modified version of the Schrödinger equation) is of this form [5]. On the other hand, the graviton model produces an entangling channel.

We now give a proof of this theorem. By assumption (a), there exists a generator \mathcal{L} of L_t of Lindblad form [51,52]:

$$\dot{\rho} = \mathcal{L}\rho = -i[H, \rho] - \sum_j \gamma_j \left[E_j^\dagger E_j \rho + \rho E_j^\dagger E_j - 2E_j \rho E_j^\dagger \right]. \quad (13)$$

These Lindblad operators E_j are highly constrained by the separability assumption, because they cannot be used to generate A - B entanglement. To make this precise, we write the channel in its Krauss representation $L[\rho] = \sum_{j \geq 0} L_j \rho L_j^\dagger$. Expanding for small times and comparing to

Eq. (13), one finds that the Krauss operators L_j take the form, to lowest order in dt ,

$$L_0 = 1 - iHdt + Kdt, \quad L_j = E_j \sqrt{dt}, \quad K = -\frac{1}{2} \sum_{j>0} E_j^\dagger E_j. \quad (14)$$

See, for example, Chap. 3 of Ref. [53]. Now we invoke the separability criterion (c), which says that the Krauss operators for $j > 0$ take the form of simple product operators, i.e., $E_j = A_j \otimes B_j$ [50]. Furthermore, the separability of L_0 to order dt means that $L_0 = (1 + A_0 dt) \otimes (1 + B_0 dt)$ for some A_0, B_0 and this can only be satisfied if both H and $E_j^\dagger E_j$ can be written as sums of operators acting either on \mathcal{H}_A or \mathcal{H}_B . This in turn requires that for each $j > 0$, either $A_j^\dagger A_j = 1_A$ or $B_j^\dagger B_j = 1_B$. Finally, we impose the requirement (b) that the atom populations are invariant. This means that $\dot{\sigma}_z = 0$. The only possible nontrivial interaction term that satisfies these requirements is $E_z = \sigma_z \otimes B$, with B any operator on H_B .

We are then left with the very simple form of the Lindblad generator:

$$\mathcal{L}\rho = -\gamma [B^\dagger B \rho + \rho B^\dagger B - 2B\sigma_z \rho \sigma_z B^\dagger] + \mathcal{L}_A + \mathcal{L}_B. \quad (15)$$

Here, $\mathcal{L}_{A(B)}$ are Lindblad operators (including Hamiltonians) acting only on $\mathcal{H}_{A(B)}$ and $\mathcal{L}_A(\sigma_z) = 0$. With this result for the structure of the channel, we can compute the time derivative of the interferometric visibility $V(t) = |\langle \sigma_-(t) \rangle|$. Since $[H, \sigma_z] = 0$, the most general qubit Hamiltonian is a sum of σ_z and the identity. We thus have, in the Heisenberg picture,

$$\left\langle \frac{d\sigma_-}{dt} \right\rangle = -i \langle [H, \sigma_-] \rangle + \gamma \left[\langle E_z^\dagger \sigma_- E_z \rangle - \frac{1}{2} \langle \{E_z^\dagger E_z, \sigma_-\} \rangle \right] = 2(-i\omega_0 - \gamma) \langle \sigma_- \rangle, \quad (16)$$

where the oscillatory term is generated by the qubit Hamiltonian. If we take the absolute value to compute the visibility $V = |\langle \sigma_- \rangle|$, this removes the oscillating phase and we have

$$\frac{dV}{dt} = -2\gamma V, \quad (17)$$

so it is monotonically decreasing, as we set out to prove.

IV. EFFECTS OF NOISE DURING EVOLUTION

The sensing protocol is subject to errors caused by random noise during the time evolution. In a typical realization, the dominant sources of this continuous noise consist of thermal load on the oscillator and dephasing in the atomic system (e.g., from background fields and

gas weakly measuring the atomic position [18,54]). These sources of noise can be modeled by a Lindblad evolution of the form

$$\dot{\rho} = -i[H, \rho] - \sum_i \frac{1}{2} \{L_i^\dagger L_i, \rho\} - L_i \rho L_i^\dagger, \quad (18)$$

where the error operators are $L_i \in \{\sqrt{\bar{n}}\gamma_m a^\dagger, \sqrt{(\bar{n}+1)}\gamma_m a, \sqrt{\gamma_a}\sigma_z\}$. The decay rates of the oscillator and atom are γ_m, γ_a , respectively, and \bar{n} is the thermal-phonon occupancy. This description should be accurate for times similar to or shorter than the damping time $1/\gamma_m$, and assuming only small changes over time in the mechanical frequency.

It is possible to analytically solve for the atomic visibility [Eq. (9)] in the presence of this noise, using an explicit Ohmic heating model where the bath is taken to be an infinite set of bosonic modes linearly coupled to the mechanical system. The same displacement-operator picture as used in Eq. (3) generalizes to this linear bath (see Appendix B). One finds that the visibility at half- and full-period evolution is given by

$$V(\pi/\omega) = \exp[-\pi\gamma_a/\omega] \exp[-8\lambda^2(2\bar{n}+1)]$$

$$V(2\pi/\omega) = \exp[-2\pi\gamma_a/\omega] \exp[-8\lambda^2(2\bar{n}+1)/Q]. \quad (19)$$

Here, we assume that the mechanical damping factor $Q = \omega/\gamma_m \gg 1$.

This recovers the previous result for the visibility given in Eq. (12), up to an overall exponential damping from the atomic dephasing and small correction from mechanical heating. If we neglect atomic dephasing, the visibility at half period is exactly the same as given by Eq. (12), while at full period, for $Q \gg 1 \gg \bar{n}\lambda^2$, we have $V(2\pi/\omega) \approx 1$, i.e., we have full recovery up to a correction at order $1/Q$. Thus, with a sufficiently high- Q oscillator and with atomic coherence times longer than the mechanical period $\gamma_a \lesssim \omega$, damping does not pose a substantial barrier to the experiment.

Before moving on, we consider the effects of decoherence from another inevitable source: black-body radiation of the oscillator. Here, we are discussing position superpositions of the oscillator at distances of about λx_0 . With the sorts of experimental parameters we suggest later, this will be a length many orders of magnitude smaller than a typical black-body photon wavelength (or the de Broglie wavelength of an ambient gas molecule). Thus these interactions are incapable of efficiently decohering the oscillator, because they are too long wavelength to efficiently measure the position of the oscillator [54].

V. PROTOCOL LINEAR IN THE WEAK COUPLING

Our basic observable [Eq. (8)] is quadratic in the ratio $\lambda = g/\omega$, which for a weak coupling is a small dimensionless number. Here, we suggest a “boosted” method in which linear sensitivity can be achieved by first preparing an *entangled* state of the atom-oscillator system (as demonstrated, for example, in Ref. [55,56]).

Let $\lambda' = g'/\omega$, where g is the coupling of interest (e.g., gravity) and g' is some other coupling. Consider performing a π gate with the coupling $V_{\text{int}} = (g + g')\sigma_z x$. This produces an initial entanglement set by displacement operators $D[\pm(\lambda + \lambda')]$, as in Eq. (9). Turning off the non-gravitational g' coupling then leads to only a partial revival of the atomic signal at later times $t > \pi/\omega$. This leads to the visibility, for $t > \pi/\omega$,

$$V_b(t) = \exp \left[-8(2\bar{n} + 1) \left(\lambda^2 + 2\lambda\lambda' \sin^2 \frac{\omega t}{2} + \lambda'^2 \sin^2 \frac{\omega t}{2} \right) \right]. \quad (20)$$

A detailed calculation is given in Appendix C. For times $0 < t < \pi/\omega$, the visibility is given by the previous result [Eq. (12)] but with $\lambda \rightarrow \lambda + \lambda'$.

The observable in which we are interested is the difference in visibility at half period and full period:

$$\begin{aligned} \Delta V_b &= V_b(2\pi/\omega) - V_b(\pi/\omega) \\ &\approx \exp[-8(2\bar{n} + 1)\lambda^2] [1 - 16(2\bar{n} + 1)\lambda'\lambda + O(\lambda^2)], \end{aligned} \quad (21)$$

assuming that $\lambda \ll \lambda'$. We see again that using an initially “hot” resonator increases the relative visibility. However, here the observable is *linear* in the weak gravitational coupling λ . We note that if \bar{n} or λ' are too large, the signal will be destroyed by the overall prefactor $e^{-8(2\bar{n}+1)\lambda^2}$. The optimal solution is to tune the nongravitational coupling

to satisfy $\lambda'_{\text{opt}} = 1/\sqrt{8(2\bar{n} + 1)}$, in which case the prefactor is order one and the relative visibility is given roughly by $\Delta V_b \approx \sqrt{8(2\bar{n} + 1)}\lambda$. Use of this boosted protocol substantially improves the viability of an experiment with a weak coupling g . We note that this protocol does not violate our assumptions about time-translation invariance in Theorem 1: once the extra g' coupling is turned off, the entire system proceeds in a time-independent fashion.

VI. EXPERIMENTAL IMPLEMENTATION WITH ATOM INTERFEROMETRY

We now show how to apply our sensing protocol to a test of quantum gravity. The idea is to realize the qubit in the Hamiltonian [Eq. (2)] as an optical-lattice atom interferometer [23] with a hold time τ and splitting ℓ between the matter wave packets. The majority of the interferometer time sees the atoms trapped in one of two different potential wells created by the lattice. The atom position thus becomes a two-state system with σ_z eigenvalues corresponding to the two locations. The mechanical oscillator has a mass M and a fundamental frequency ω . Expanding the Newtonian atom-oscillator potential [Eq. (1)], we then have the total Hamiltonian

$$H = \omega a^\dagger a - g\sigma_z(a + a^\dagger). \quad (22)$$

Here, a, a^\dagger are oscillator operators, so the second term represents the position-position coupling. The coupling strength is

$$g = \kappa \frac{G_N m M \ell x_0}{\hbar R^3}, \quad (23)$$

where $x_0 = \sqrt{\hbar/2M\omega}$ is the ground-state oscillator uncertainty, R parametrizes the distance between the oscillator and atom, and κ is a dimensionless number of order one that depends on the specific oscillator-mass geometry (see Fig. 3).

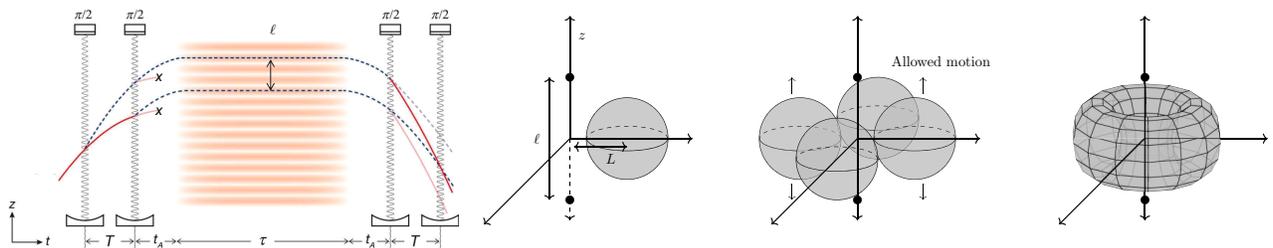


FIG. 3. Left: the experimental realization of the atomic system as a lattice interferometer. The lines marked “x” denote populations that do not interfere. Right: some example implementations with one or more mechanical masses connected rigidly. The small black dots represent the atom. In each case, the mechanical system is restricted to oscillate along the z axis. More masses enable a stronger gravitational coupling. A natural limiting case would be to use a toroidal mass. In the example with a single sphere, we have $R = \sqrt{L^2 + (\ell/2)^2}$ and $\kappa = 1$.

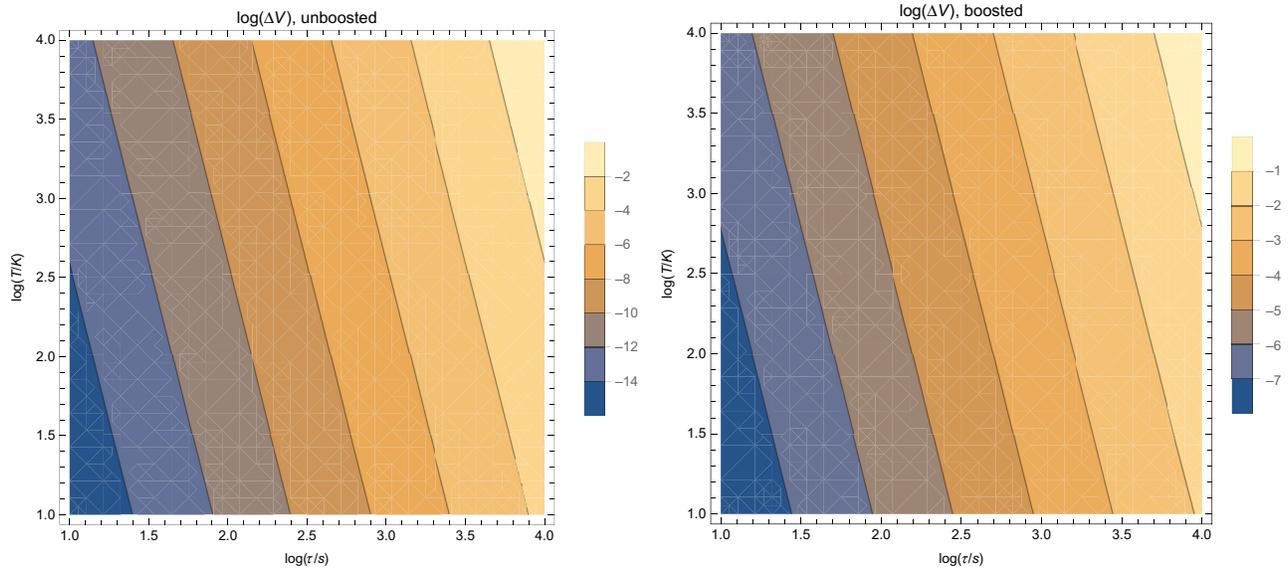


FIG. 4. The logarithm \log_{10} of the visibility change ΔV as a function of the logarithms of the hold time τ in seconds and the temperature T in kelvin for the unboosted scheme (left) and the boosted scheme (right). The plots assume $\ell = 1$ mm, $\rho = 20$ g/cm³, and $m = m_{\text{Cs}}$.

The information sensing protocol requires generation of an initial state $|0\rangle + |1\rangle$. This can be generated, for example, by a pair of Raman pulses separated by a free evolution time [23], by spin-dependent kicks [63], by optical lattice techniques [64], or by rapidly splitting a single-well potential to the double well. Measuring in the σ_z basis at the end of the protocol corresponds to closing the atom interferometer and counting the atoms in the two output ports. To implement the “boosted” protocol of Sec. V, we can use a number of nongravitational interactions to generate the initial entanglement. For example, a hyperfine or Rydberg atomic state could be magnetically or optically coupled to the oscillator. Entanglement of this type has recently been demonstrated experimentally [55,56].

Let us consider how we can obtain a visibility change that is large enough to be measured. In order to observe at least one full cycle of decay and revival, we choose $\omega = 2\pi/\tau$ where τ is the atom hold time. In this case, the

visibility change is given by

$$\Delta V = \frac{\pi G_N^2 m^2 \rho}{3\sqrt{2}\ell\omega^3 \hbar} (8 + \bar{n}) \xrightarrow{k_B T/\omega \rightarrow \infty} \frac{\pi}{3\sqrt{2}} K^2, \quad (24)$$

$$\Delta V_b = 2^{1/4} G_N m \sqrt{\frac{\pi \rho}{3\ell\omega^3 \hbar}} (8 + \bar{n}) \xrightarrow{k_B T/\omega \rightarrow \infty} \frac{2^{1/4}}{\sqrt{3}} K$$

in the unboosted and boosted scheme, respectively, where

$$K^2 = \frac{G_N^2 m^2 \rho k_B T}{\ell \omega^4 \hbar^2}$$

$$\approx 1.04 \times 10^{-14} \left(\frac{T}{300 \text{ K}} \right) \left(\frac{\ell}{1 \text{ mm}} \right)^{-1} \left(\frac{\tau}{10 \text{ s}} \right)^4. \quad (25)$$

Here, we take a solid density $\rho = 20$ g/cm³, cesium atoms $m = m_{\text{Cs}} = 133$ amu, use the four-sphere configuration

TABLE I. Some systematic effects and other perturbations expected in a realistic implementation.

Technical challenge	Examples	Possible strategies
Nongravitational interactions	Van der Waals, stray fields, scattered laser light	Superconducting shielding, place atoms in waveguide [57]
Mean field shift	Parasitic atom-atom interactions leading to inhomogeneous dephasing [58]	Spin-echo techniques [59] (see also Appendix E), fermionic atoms (e.g., ¹⁷¹ Yb or ¹⁷³ Yb) [60,61]
Exponential decay of signal	Atomic dephasing	Interleaved differential measurement, e.g., by toggling the mass between near and far positions [62]
Deviations from harmonicity	Time-dependent oscillator frequency, anharmonic perturbations	Keep effective temperature below nonlinear thresholds; change materials, mounting, or frequency

(Fig. 3) for definiteness, and maximize the coupling g for a given splitting ℓ by choosing a sphere radius of $R_s = \ell/(\sqrt{8})$. Longer atomic interrogation times τ are preferable. This would require a correspondingly low-frequency oscillator, e.g., a megahertz-scale torsional pendulum. While 20 s has been experimentally realized [23], 100 s may be a reasonable expectation for the future. Using a small matter-wave splitting ℓ is desirable, but subject to mechanical constraints. Choosing, for example, $\ell = 1$ mm, $L = 1/\sqrt{2}$ mm and $R_s = 0.35$ mm would leave about 0.15 mm free space between the spheres. For $\tau = 100$ s and $T = 300$ K, we obtain $\Delta V \sim 10^{-10}$; but for the boosted scheme, it will be as large as $\Delta V = 7 \times 10^{-6}$ (see Fig. 4). At the standard quantum limit, this can be detected with $5 - \sigma$ significance by running the experiment with approximately 5×10^{11} atoms (for details on noise scaling with many atoms, see Appendix D). Assuming that the experiment has 10^7 atoms per run and that each run takes 2 min, this will be possible in a total run time of two months.

Remarkably, this suggests that the experiment may be feasible in the near future. A number of systematic effects and technical issues need to be understood. We postpone detailed discussion to future work, but flag some likely issues and ways to handle them in Table I.

VII. IMPLICATIONS, LOOPHOLES, AND CONCLUSIONS

Our interactive information sensing protocol is a novel strategy for verification of dynamical entanglement generation. While a standard Bell-type test requires measurements on both parts of a bipartite system, our protocol can verify entanglement generation with only single-body measurements. Crucially, the test verifies the ability of an interaction channel to generate entanglement, without needing to directly verify the entanglement of the final

state. However, it is important to note that this test is subject to loopholes. Some are analogous to those in standard Bell tests and others are particular to our proposal. We suggest a few of these in Table II.

In our view, the most important loophole stems from our time-translation invariance assumption, which we use to write the atom-oscillator dynamics in Lindblad form [Eq. (13)]. The non-Markovian time dependence introduced by an experimentalist or Maxwell's demon could, in principle, reproduce the observed collapse-and-revival dynamics. One way to improve the situation would be to reformulate the theorem to include some level of non-Markovianity; for example, a bath-relaxation time scale. A more robust option would be to prove experimentally that it is simply the Markovian thermalizing channel acting on the mechanical system. Methods for this include precision quantum thermometry [65], which can support the hypothesis of detailed balance. In any case, extending the results here beyond the strictly Markovian assumption is a crucial next step.

The central technical advances suggested here are the interactive sensing protocol and the use of atoms as a sensor. The key advantage of the periodic collapse-and-revival protocol is that it enables a huge enhancement with a thermal state of the mechanical system; understanding if this can be extended beyond the specific context here would be very interesting. While using trapped atoms is perhaps counterintuitive since it decreases the strength of the signal (the Newton potential), we emphasize that the extremely long coherence lifetime and ability to generate spatially well-separated superpositions of the atoms lead to similar parametric scaling of the overall signal strength.

We show how the interactive sensing protocol can be used to test the ability of the gravitational field to communicate quantum information. If the answer is yes, this would constitute the first direct evidence that the

TABLE II. Some loopholes and pathologies in our proposed test.

Loophole or pathology	Typical sources	Problematic behavior allowed	Possible solutions
Nongravitational interactions between atom and oscillator	Casimir forces, van der Waals interactions	Can generate entanglement (reproduce the full desired signal), can generate extra noise	Vary parameters (masses and distance) to check proper scaling with $V = G_N m_1 m_2 / r$ law
Stationarity assumption on bath (and/or experimentalist) violated	Explicit time dependence introduced by experimentalist (e.g., spin-echo protocol); low-frequency noise (e.g., gravity gradients, seismic noise)	Violates assumption of theorem in Sec. III. In principle, could mimic collapse and revival	Adjust theorem to allow for bath relaxation time scale; experimentally verify Markovian nature of oscillator noise
Nonlocality	Time of interaction for experiment is much longer than light-crossing time $T_{\text{int}} \gg T_{\text{com}}$	Allows for nonlocal hidden-variable model explaining the entanglement (same as Bell test)	Long-baseline version?

gravitational field itself is a quantum mechanical degree of freedom [13,26–29]. On the contrary, if the answer is negative, the existence of the graviton is ruled out [13]. The simple estimates of Sec. VI suggest that this experiment is feasible with realistic devices, even in the presence of noise. We will present a more detailed proposal and analysis of systematic effects in a future paper.

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APPENDIX A: NEWTONIAN ENTANGLEMENT FROM GRAVITON EXCHANGE

For completeness, we review here some standard arguments about the perturbative quantization of gravity and its relation to entanglement generation via the Newton potential [Eq. (1)]. Our goal is to explain the standard logic by which one treats small fluctuations of the metric as a quantum field and uses this to make predictions in nonrelativistic systems. We do not mean to say that this derivation somehow proves that this is the correct model of low-energy quantum gravity—on the contrary, determining if this is the correct set of predictions is a central objective of the experiment proposed in this paper.

By far the most common and efficient method to compare a field-theoretical description to the nonrelativistic setting relevant to these experiments is to do a “matching” calculation. For example, one can compute scattering amplitudes in the field theory, compare these to the same amplitude computed in a potential scattering model, and thus obtain the effective nonrelativistic potential. Since the scattering states form a complete basis for the Hilbert space (other than bound states), if these two calculations agree for all scattering states, we can conclude that the two descriptions are equivalent quantum-mechanically in the regime in which the calculations match.

For example, consider the electrodynamics: photons A_μ coupled to a Dirac fermion ψ (e.g., the electron, with charge e). For a textbook treatment of what follows, see Sec. 4.8 of Ref. [66]. The Lagrangian is

$$\mathcal{L} = \frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \bar{\psi}\gamma^\mu(\partial_\mu - ieA_\mu)\psi, \quad (\text{A1})$$

where $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ the electromagnetic field tensor and γ^μ are the usual Dirac matrices. With this Lagrangian, we can perturbatively compute the elastic $e^-e^- \rightarrow e^-e^-$ scattering amplitude to lowest order in the charge using

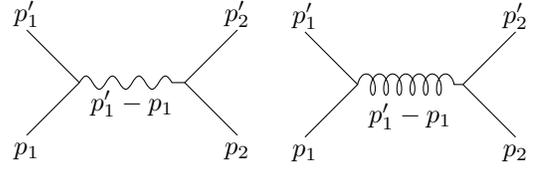


FIG. 5. Feynman diagrams for single-photon and -graviton exchange, respectively.

standard methods. See the Feynman diagram of Fig. 5. One finds [67]

$$\langle \mathbf{p}'_1 \mathbf{p}'_2 | T | \mathbf{p}_1 \mathbf{p}_2 \rangle = 4\pi e^2 \frac{N_{\text{EM}}}{-t - i\epsilon}, \quad (\text{A2})$$

where $t = -(p'_1 - p_1)^2$ is the Lorentz-invariant four-momentum transfer and the numerator is

$$N_{\text{EM}} = \bar{u}(\mathbf{p}'_1)\gamma^\mu u(\mathbf{p}_1)\bar{u}(\mathbf{p}'_2)\gamma_\mu u(\mathbf{p}_2). \quad (\text{A3})$$

Here, u, \bar{u} are Dirac spinors. In the nonrelativistic limit where the rest masses dominate over the spatial momenta, the numerator reduces to $N_{\text{EM}} \rightarrow 1$ and $t \rightarrow (\mathbf{p}'_1 - \mathbf{p}_1)^2$. In the center-of-mass frame, this means that the amplitude for the relative momentum $\mathbf{p} = (\mathbf{p}_1 - \mathbf{p}_2)/2$ to transition to \mathbf{p}' is given by

$$\langle \mathbf{p}' | T | \mathbf{p} \rangle = \frac{4\pi e^2}{(\mathbf{p}' - \mathbf{p})^2}, \quad (\text{A4})$$

where we use conservation of the total momentum to write $\mathbf{p}'_1 - \mathbf{p}_1 = \mathbf{p}' - \mathbf{p}$.

We can then compare this to the amplitude computed nonrelativistically with a Hamiltonian description

$$H = H_1 + H_2 + V(r), \quad (\text{A5})$$

including a central potential $V(r) = V(|\mathbf{x}_1 - \mathbf{x}_2|)$. In the center-of-mass frame, the first Born approximation gives

$$\langle \mathbf{p}' | T | \mathbf{p} \rangle = \tilde{V}(\mathbf{p}' - \mathbf{p}), \quad (\text{A6})$$

where \tilde{V} is the Fourier transform of the potential. Comparing these expressions, we see that the effective potential is

$$\tilde{V}(\mathbf{q}) = \frac{4\pi e^2}{\mathbf{q}^2}, \quad (\text{A7})$$

or, in position space,

$$V(\mathbf{x}) = \int \frac{d^3\mathbf{q}}{(2\pi)^3} e^{i\mathbf{q}\cdot\mathbf{x}} \tilde{V}(\mathbf{q}) = \frac{e^2}{|\mathbf{x}|}, \quad (\text{A8})$$

recovering the usual Coulomb force.

We emphasize that this line of argument is at the level of transition *amplitudes* $\langle \psi_f | U(t_f, t_i) | \psi_i \rangle$, where U means time evolution and we take $t_i \rightarrow -\infty, t_f \rightarrow +\infty$. Thus, the potential [Eq. (A8)] is the potential operator in the full meaning of the term—it is a two-body operator and can entangle particles, and so forth. Indeed, entanglement generated by this Coulomb potential is exactly what underlies, for example, quantum information processing with chains of trapped ions in Coulomb crystals.

The same type of matching calculation can be performed in gravity. The additional complication is that we need to work perturbatively around a background space-time and only quantize the perturbations. For an experiment well localized in space-time like ours, we can expand around a locally flat space-time metric $\eta_{\mu\nu}$, by the equivalence principle [68]. We write

$$g_{\mu\nu} = \eta_{\mu\nu} + \frac{h_{\mu\nu}}{m_{\text{Pl}}}, \quad (\text{A9})$$

where we scale out a factor of the Planck mass to give $h_{\mu\nu}$ dimensions of mass (i.e., the same dimensions as a canonical bosonic field in four dimensions). In this expansion, the Einstein-Hilbert Lagrangian becomes

$$\mathcal{L} = \frac{1}{2} \partial_\alpha h_{\mu\nu} \partial^\alpha h^{\mu\nu} + \dots + \frac{1}{m_{\text{Pl}}} h_{\mu\nu} T^{\mu\nu}, \quad (\text{A10})$$

where the dots represent other similar second-derivative kinetic terms for the perturbations $h_{\mu\nu}$ and $T_{\mu\nu}$ is the matter stress tensor. These metric perturbations $h_{\mu\nu}$ can be quantized through the exact same procedure as the electromechanical potential A_μ [30–33]. There is a key difference—namely that the interaction is nonrenormalizable—which means that we have an effective quantum field theory [34,35,69–71], which can only make reliable predictions at energy densities well below the Planck scale [72]. However, we are well within this limit in the kind of experiment envisaged here, as discussed in Sec. I.

Proceeding accordingly, the scattering of a pair of masses (here, modeled as single-particle excitations of a massive spinless field) via gravitons is given by [34,35]

$$\langle \mathbf{p}'_1 \mathbf{p}'_2 | T | \mathbf{p}_1 \mathbf{p}_2 \rangle = \frac{4\pi}{m_{\text{Pl}}^2} \frac{N_{\text{grav}}}{-t - i\epsilon}. \quad (\text{A11})$$

The numerator is more complicated due to the tensorial nature of the interaction,

$$N_{\text{grav}} = 2(p_1 \cdot p'_1)(p_2 \cdot p'_2) + 2(p_1 \cdot p'_2)(p'_1 \cdot p_2) + 8(p_1 \cdot p'_1 + m^2)(p_2 \cdot p'_2 + m^2), \quad (\text{A12})$$

but reduces in the nonrelativistic limit to the simple value $N_{\text{grav}} \rightarrow m^2$. Recognizing that $m_{\text{Pl}}^2 = 1/G_N$ in terms of the Newton constant, we can compare this again to the

Born approximation [Eq. (A6)] and determine the effective potential

$$\tilde{V}(\mathbf{q}) = \frac{4\pi G_N m^2}{\mathbf{q}^2}, \quad (\text{A13})$$

which again is just $V(r) = G_N m^2/r$ in real space. In this way, we see that “graviton exchange” leads to the Newton potential operator [Eq. (1)] in the nonrelativistic limit.

Finally, we note that can one directly obtain a Hamiltonian operator for the field theory and read off the nonrelativistic potential directly, without resorting to scattering or other matching calculations. In contrast to the gauge-invariant scattering amplitude approach, this is complicated by the gauge symmetries of the model (in both the electrodynamics and gravity cases). To see how this works, consider the electrodynamical Lagrangian [Eq. (A1)]. To perform the transformation from the Lagrangian to Hamiltonian we have to fix a gauge—say, Coulomb gauge $\partial_i A^i = 0$. This gauge leads to a second-class Dirac constraint $\partial_i F^{i0} = -J^0$, so that $\nabla^2 A^0 = -J^0$, i.e., the A^0 part of the potential is nondynamical and simply fixed by the current

$$A^0(\mathbf{x}, t) = - \int d^3\mathbf{y} \frac{J^0(\mathbf{y}, t)}{|\mathbf{x} - \mathbf{y}|}. \quad (\text{A14})$$

Performing the Legendre transformation to obtain the Hamiltonian, one then finds a coupling

$$\begin{aligned} H_{\text{Coul}} &= - \int d^3\mathbf{x} A^0(\mathbf{x}) J^0(\mathbf{x}) \\ &= \int d^3\mathbf{x} d^3\mathbf{y} \frac{J^0(\mathbf{x}) J^0(\mathbf{y})}{|\mathbf{x} - \mathbf{y}|}, \end{aligned} \quad (\text{A15})$$

which is just the usual nonrelativistic Coulomb interaction, since J^0 is the charge density. Again, everything here is at the level of operators. In the case of gravity, the exactly analogous calculation can be performed and one finds an instantaneous Newton interaction $H = m_{\text{Pl}}^{-2} \int d^3\mathbf{x} d^3\mathbf{y} T^{00}(\mathbf{x}) T^{00}(\mathbf{y}) / |\mathbf{x} - \mathbf{y}|$. For a thorough treatment of this kind of Hamiltonian approach in the case of electrodynamics, see Chap. 8 of Ref. [73] and for the calculation in perturbative gravity, see, e.g., Ref. [74].

The fact that the field component responsible for the Coulomb-Newton interaction is nondynamical (e.g., A^0 in the above example) has led some authors to argue that observing Newtonian entanglement would tell us nothing about the quantization of the “physical” (i.e., dynamical) degrees of freedom of the gravitational field; for a prototypical expression of this view, see Ref. [75]. There are, however, strong arguments against this [26–29], which essentially say that there is no consistent way to have both an entangling Newton interaction and nonquantized metric fluctuations. We anticipate substantial further debate on this topic and will present a detailed discussion in a separate paper [76].

APPENDIX B: DETAILED CALCULATION OF OSCILLATOR NOISE

To obtain a quantitative estimate of the effects of thermal loading on the oscillator, let us assume that we can safely neglect atomic dephasing $\gamma_a t \ll 1$ for the experimental time scale of interest. To develop an exact result, we use an input-operator method, in which we include an explicit heat bath for the oscillator. The Lindblad system [Eq. (18)] can be derived through this method by tracing out the oscillator bath. For a lucid review of this technique, see Ref. [77].

Let H_0 denote the Hamiltonian for the oscillator *and its bath*. The total time-evolution operator is then

$$U(t) = e^{-i(H_0 + g\sigma_z)t} = e^{-iH_0 t} \hat{T} e^{-ig\sigma_z \int_0^t x_I(t') dt'}. \quad (\text{B1})$$

Here, $x_I(t)$ is the oscillator position operator in the interaction picture and \hat{T} is the time-ordering operator. For the case of a linear bath, such as that assumed in quantum optics or in quantum Brownian motion, we can explicitly find x_I . Writing $x = (a + a^\dagger)/\sqrt{2}$, we have

$$a_I(t) = \exp[-i(\omega + \gamma_m/2)t] a(0) + \sqrt{\gamma_m} \int_0^t \exp[-i(\omega + \gamma_m/2)(t - t')] a_{\text{in}}(t') dt', \quad (\text{B2})$$

where $a_{\text{in}}(t)$ is the vacuum-noise-fluctuation operator, satisfying $[a_{\text{in}}(t), a_{\text{in}}^\dagger(t')] = \delta(t - t')$. Using the linearity of this expression and the Baker-Campbell-Hausdorff relation, we then have that

$$\hat{T} \exp\left(-ig\sigma_z \int_0^\tau x_I(t) dt\right) = \exp\left(-ig\sigma_z \int_0^\tau x_I(t) dt\right) \exp[-ig^2 C(t)], \quad (\text{B3})$$

where $C(t)$ is a real time-dependent number, arising from the noncommuting elements of $x_I(t)$.

Having dispensed with the time ordering, we can now explicitly perform the time integral (including a change of integration order in the a_{in} term). Dropping the $e^{-ig^2 C(t)}$ phase, which cancels out of our observable, we find that the time evolution reduces to a simple product of displacement operators, one for the oscillator and one for each mode $a_{\text{in}}(t')$ for $0 \leq t' \leq t$; that is,

$$U(t) = e^{-iH_0 t} D_a[\sigma_z \alpha(t)] \prod_{0 \leq t' \leq t} D_{a_{\text{in}}(t')}[\sigma_z \alpha_{\text{in}}(t')], \quad (\text{B4})$$

where

$$\alpha(t) = \frac{ig}{i\omega - \gamma_m/2} (1 - e^{(i\omega - \gamma_m/2)t}), \quad (\text{B5})$$

$$\alpha_{\text{in}}(t') = \frac{ig}{i\omega - \gamma_m/2} (1 - e^{(i\omega - \gamma_m/2)(t-t')}).$$

Finally, we can evaluate our visibility $\sigma_-(t) = U^\dagger(t) \sigma_- U(t)$, assuming an initial thermal state for the oscillator and each bath mode and the $|+\rangle$ state for the atom. Using the same results for coherent states as above, one finds

$$\langle \sigma_- \rangle = \langle D_a[2\alpha(t)] \rangle \prod_{t'} \langle D_{a_{\text{in}}(t')}[\alpha_{\text{in}}(t')] \rangle = \exp[-8\lambda^2(2\bar{n} + 1)f(t)], \quad (\text{B6})$$

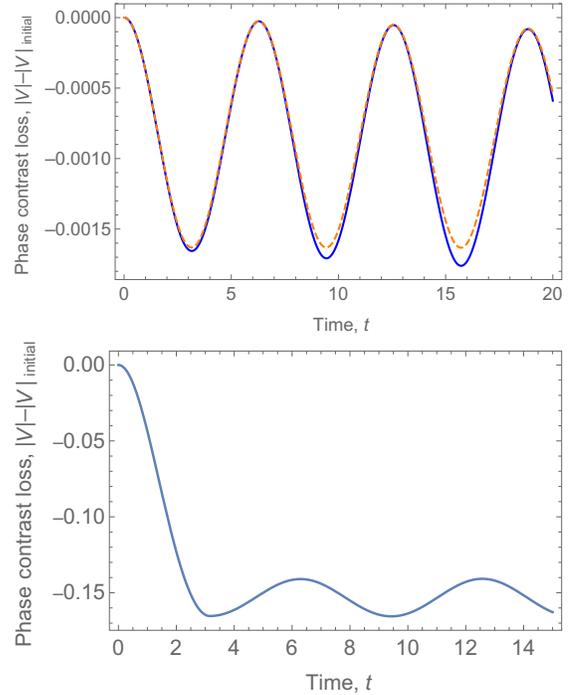


FIG. 6. Examples of the signal of interest, the phase contrast $V = |\langle \sigma_-(t) \rangle|$, compared with its initial value $V(0) = 1/2$. Left: direct simulation of the Lindblad evolution [Eq. (18)] (in blue)—we see good agreement with our analytic solution including noise [Eq. (B6)] (in orange, dashed). Normalizing all units to the oscillator frequency $\omega = 1$, here we use values $g = 10^{-2}$ for the gravitational coupling, $\gamma_m = 5 \times 10^{-3}$ for the mechanical damping, and $T = 2$ for the temperature. (Numerical simulation with a much higher $T \gg \omega$ as discussed in the paper is infeasible due to restrictions on the oscillator Hilbert-space dimension.) Bottom: the same parameters as the top figure, but with an initial π pulse using a nongravitational coupling $g' = 10^{-1}$. The difference between the first collapse and revival is now much larger than in the unboosted protocol, as predicted in Eq. (21).

with

$$f(t) = \frac{\omega^2/4}{\omega^2 + \gamma_m^2/4} \left(2 - 2 \cos(\omega t) e^{-\gamma_m t/2} + \gamma_m t - \frac{8\gamma_m}{\omega} \sin(\omega t) e^{-\gamma_m t/2} + O(1/Q^2) \right), \quad (\text{B7})$$

where $Q = \omega/\gamma_m$ is assumed to be much larger than one. In particular, at full and half period this gives

$$\begin{aligned} V(\pi/\omega) &= \exp[-8\lambda^2(2\bar{n} + 1)], \\ V(2\pi/\omega) &= \exp[-8\lambda^2(2\bar{n} + 1)/Q]. \end{aligned} \quad (\text{B8})$$

Here, we assume that the mechanical damping factor $Q = \omega/\gamma_m \gg 1$. Reinserting the exponential damping factor for atomic dephasing then reproduces the results in Eq. (19).

In Fig. 6, we compare this analytical model with a numerical simulation of the Lindblad equation [Eq. (13)], showing excellent agreement.

APPENDIX C: DETAILED CALCULATION OF THE BOOSTED PROTOCOL

Here, we give the full computation of the visibility in our entanglement-enhanced “boosted” protocol of Sec. V. The total evolution is a product of two unitaries, one for the first half period under the coupling $g + g'$ and the second under only g . We write these as

$$\begin{aligned} U_{g+g'} &= D^\dagger[(\lambda + \lambda')\sigma_z] e^{-i\omega n} D[(\lambda + \lambda')\sigma_z], \\ U_g(t) &= D^\dagger(\lambda\sigma_z) e^{-i\omega n(t-\pi/\omega)} D(\lambda\sigma_z). \end{aligned} \quad (\text{C1})$$

With this notation, the visibility of the atom, given some initial coherent state $|\alpha\rangle$ for the oscillator, is given by (defining $\tilde{\lambda} = \lambda + \lambda'$ for brevity)

$$\begin{aligned} V_{b,\alpha}(t) &= \langle \alpha | U_{g+g'}^\dagger U_g(t) \sigma_- U_g(t) U_{g+g'}(t) | \alpha \rangle \\ &= \langle \alpha | D^\dagger(-\tilde{\lambda}) e^{i\omega n/\pi} D(-\tilde{\lambda}) U_g^\dagger(t) \\ &\quad \times \sigma_- U_g(t) D^\dagger(\tilde{\lambda}) e^{-i\omega n/\pi} D(\tilde{\lambda}) | \alpha \rangle \\ &= \langle \alpha | D(2\tilde{\lambda}) e^{i\omega n/\pi} U_g^\dagger(t) \sigma_- U_g(t) e^{-i\omega n/\pi} D(2\tilde{\lambda}) | \alpha \rangle \\ &= \langle \alpha | D \left[2\tilde{\lambda} - \lambda(1 + e^{i\omega t}) \right] D \left[2\tilde{\lambda} - \lambda(1 + e^{i\omega t}) \right] | \alpha \rangle \\ &= \langle 0 | D(-\alpha) D \left[2\tilde{\lambda} - \lambda(1 + e^{i\omega t}) \right] \\ &\quad \times D \left[2\tilde{\lambda} - \lambda(1 + e^{i\omega t}) \right] D(\alpha) | 0 \rangle \\ &= e^\phi \langle \alpha - 2\tilde{\lambda} + \lambda(1 + e^{i\omega t}) | \alpha + 2\tilde{\lambda} - \lambda(1 + e^{i\omega t}) \rangle \\ &= e^{2\phi} e^{-|4\tilde{\lambda}^2 - 2\lambda(1 + e^{i\omega t})|^2/2}. \end{aligned} \quad (\text{C2})$$

To go from the second to the third line, we insert a pair of identity operators $1 = e^{-i\omega n/\pi} e^{i\omega n/\pi}$ and use

$e^{i\omega n/\pi} D(\tilde{\lambda}) e^{-i\omega n/\pi} = D(-\tilde{\lambda})$. From the third line to fourth line, we use the same trick and the more general time evolution $e^{i\omega n} D(\tilde{\lambda}) e^{-i\omega n} = D(\tilde{\lambda} e^{i\omega t})$. In the last few lines, the “phase” is

$$\phi = \alpha^* [2\tilde{\lambda} - \lambda(1 + e^{i\omega t})]/2 + \text{c.c.} \quad (\text{C3})$$

Note that we obtain two factors of this: one in the fifth line, from the braiding relation $D(\alpha)D(\beta) = e^{(\alpha\beta^* - \alpha^*\beta)/2}$, and then another in the subsequent line, from the inner product $\langle \beta | \alpha \rangle = e^{-|\beta - \alpha|^2/2} e^{(\alpha\beta^* - \alpha^*\beta)/2}$. Note also that the second exponential does not depend on the coherent state parameter α . Thus we only need to average this phase term over the Glauber representation, which gives

$$\begin{aligned} \int \frac{d^2\alpha}{\pi\bar{n}} e^{-|\alpha|^2/\bar{n}} e^{2\phi} &= \exp \left[8\lambda n (\lambda + 2\lambda') \cos(t\omega) \right. \\ &\quad \left. - 8n (\lambda^2 + 2\lambda\lambda' + 2\lambda'^2) \right], \end{aligned} \quad (\text{C4})$$

where we use the explicit coefficient $\tilde{\lambda} = \lambda + \lambda'$. Doing the same with the second term in Eq. (C2) and simplifying the terms, we finally obtain

$$\begin{aligned} V_b(t) &= \int \frac{d^2\alpha}{\pi\bar{n}} V_{b,\alpha}(t) \\ &= \exp \left[-8(2\bar{n} + 1) \left(\lambda^2 + 2\lambda\lambda' \sin^2 \frac{\omega t}{2} \right. \right. \\ &\quad \left. \left. + \lambda^2 \sin^2 \frac{\omega t}{2} \right) \right], \end{aligned} \quad (\text{C5})$$

as quoted in Eq. (20). Note that the limit $\lambda' \rightarrow 0$ reproduces the basic unboosted protocol. We show the form of this visibility evolution in Fig. 6.

APPENDIX D: USING MANY ATOMS

The sensitivity of the protocol can be substantially improved by moving from using a single atom to using a collection of atoms, as is typical in an atom interferometer [20–22]. For simplicity, we take g the same for all the atoms, though that is not necessary in practice. In that limit, we can define $J_z = \sum_j \sigma_z^j$ and $J_- = \sum_j \sigma_-^j$ as the collective variables that will enter.

Consider the extension of Eq. (9) to the case of N atoms prepared in the initial state $|++\dots\rangle$. The “observable” of interest is $\langle J_-(t) \rangle$. This is easiest to calculate term by term for each atom. The total time-evolution operator, following the same logic as in Eq. (3), is

$$U(t) = \exp(-i\beta J_z^2) D^\dagger(\lambda J_z) e^{i\omega a^\dagger a t} D(\lambda J_z), \quad (\text{D1})$$

with

$$\beta = \frac{g^2 t}{\omega}. \quad (\text{D2})$$

This total J_z^2 term is peculiar to the case of $N > 1$ atoms; for $N = 1$, it is just an overall phase that we drop in Eq. (3). Here, however, it is a nontrivial operator, physically representing the ponderomotive squeezing of the spins due to the gravitational coupling with the oscillator. At time $t > 0$, we have, for each $i = 1, \dots, N$, the operator evolution

$$\sigma_-^i(t) = U^\dagger(t)\sigma_-^i U(t). \quad (\text{D3})$$

If we define $\tilde{J}_z^i = J_z - \sigma_z^i$, we can see that we can write the J_z^2 contribution to Eq. (D3) as

$$\exp(i\beta J_z^2)\sigma_-^i \exp(-i\beta J_z^2) = \exp(-2i\beta\tilde{J}_z^i)\sigma_-^i. \quad (\text{D4})$$

This prefactor then commutes with the rest of the operators in Eq. (D3). Using this and the same basic logic as in Eq. (9), we find that all the $i \neq j$ spins just give a phase proportional to \tilde{J}_z :

$$\sigma_-^i(t) = \sigma_-^i e^{-2i\beta\tilde{J}_z} D(-\lambda) D(2\lambda e^{i\omega t}) D(-\lambda). \quad (\text{D5})$$

Acting on the initial state $|0, + + + \dots\rangle$ with the oscillator prepared in $|0\rangle$ and each atom in the $|+\rangle$ state—that is, with N unentangled atoms—we obtain

$$\langle \sigma_-^i(t) \rangle = \cos^{N-1}(2\beta) \langle \sigma_-(t) \rangle_1, \quad (\text{D6})$$

where the term $\langle \sigma_-(t) \rangle_1$ denotes the answer with a single spin, as in Eqs. (9) and (10). For $N \gg 1$ and $\beta \ll 1$ (recall $\beta = g^2 t / \omega$, so this condition is certainly satisfied for us), we can Taylor expand the cosine and match it to an exponential for convenience, i.e., $\cos^{N-1}(2\beta) \approx e^{-2N\beta^2}$. Thus since $\langle J_- \rangle$ has N terms of the form given in Eq. (D6), we finally obtain

$$\langle J_-(t) \rangle = N e^{-2Ng^4 t^2 / \omega^2} \langle \sigma_-(t) \rangle_1. \quad (\text{D7})$$

For example, at a half period and a full period, we then have

$$\begin{aligned} \langle J_-(\pi/\omega) \rangle &= N e^{-2\pi^2 N \lambda^4} e^{-8\lambda^2}, \\ \langle J_-(2\pi/\omega) \rangle &= N e^{-4\pi^2 N \lambda^4}. \end{aligned} \quad (\text{D8})$$

Note that this noisy phase is independent of the initial state of the oscillator; so, for example, we obtain the same answer if the oscillator begins in a thermal state. We see the basic N enhancement to the signal here as the prefactor. The phase noise scales like $N\lambda^4$. For our particular implementation with parameters such as those quoted in Eq. (25), we have $\lambda \sim 10^{-13}$, so for $N \sim 10^{10}$ atoms these phase-noise exponentials are completely negligible. The overall N factor here represents the usual \sqrt{N} statistical enhancement in the signal-to-noise ratio, assuming uncorrelated atom errors.

The same calculation extends directly to the entanglement-enhanced, g -linear protocol [Eq. (20)]. This is clear by the algebraic structure of the argument given above. Explicitly, we now have two time-evolution operators of the form given in Eq. (D1), one with a coupling $g + g'$ from $t = -\pi/\omega$ to $t = 0$, followed by another with only the gravitational g coupling from $t = 0$ onward. In an obvious notation, we can write

$$\sigma_-^i(t) = U_g^\dagger(t) U_{g+g'}^\dagger \sigma_-^i U_{g+g'} U_g(t). \quad (\text{D9})$$

In these U operators, we have the same phase-noise terms, namely $e^{i\beta J_z^2}$ in the U_g and another factor $e^{i\beta' J_z^2}$, with $\beta' = -(g + g')^2 \pi / \omega$, from the $U_{g+g'}$ factor. These depend only on the J_z operator and thus commute with the other terms (displacement operators and free-oscillator evolution) in U_g and $U_{g+g'}$. Thus we obtain an expression

$$\begin{aligned} \sigma_-^i(t) &\sim \exp[i(\beta + \beta') J_z^2] \sigma_-^i \exp[-i(\beta + \beta') J_z^2] \\ &= \exp[-2i(\beta + \beta') \tilde{J}_z] \sigma_-^i, \end{aligned} \quad (\text{D10})$$

just as in Eq. (D4), except now with β replaced by

$$\beta + \beta' = \frac{g^2 t}{\omega} - \frac{(g + g')^2 \pi}{\omega}. \quad (\text{D11})$$

In particular, all the terms other than these phase-noise exponentials contribute as given by Eq. (20). The overall signal is still increased linearly in N , as in Eq. (D8), times a negligible contribution from this ponderomotive squeezing noise.

APPENDIX E: SPIN-ECHO VERSION FOR FASTER PHYSICAL OSCILLATORS

The physical oscillator frequency is crucially important to the size of the observable effect. The interferometric contrast scales as a power of $\lambda = g/\omega$, so a low-frequency oscillator is ideal. However, in practice, the use of a very-low-frequency (subhertz) oscillator would present substantial technical problems. This can be alleviated by using a high-frequency oscillator and a spin-echo-like sequence to mimic the effect of a low-frequency oscillator. Specifically, after every π/ω half period, we swap the two atom locations, i.e., perform a σ_x operation (see Fig. 7). This produces the evolution

$$\begin{aligned} U &= \sigma_x e^{-iH\tau/2} \sigma_x e^{-iH\tau/2} \\ &= e^{-i(H_p - V)\tau/2} e^{-i(H_p + V)\tau/2} \\ &= D(\lambda\sigma_z) e^{-iH_p\tau/2} D(-\lambda\sigma_z) D(-\lambda\sigma_z) e^{-iH_p\tau/2} D(\lambda\sigma_z) \\ &= D(\lambda\sigma_z) D(2\lambda\sigma_z) D(\lambda\sigma_z) = D(4\lambda\sigma_z), \end{aligned} \quad (\text{E1})$$

where we use $\tau = 2\pi/\omega$ for the final line. With N_π iterations of this, we produce the total evolution $U(N_\pi) =$

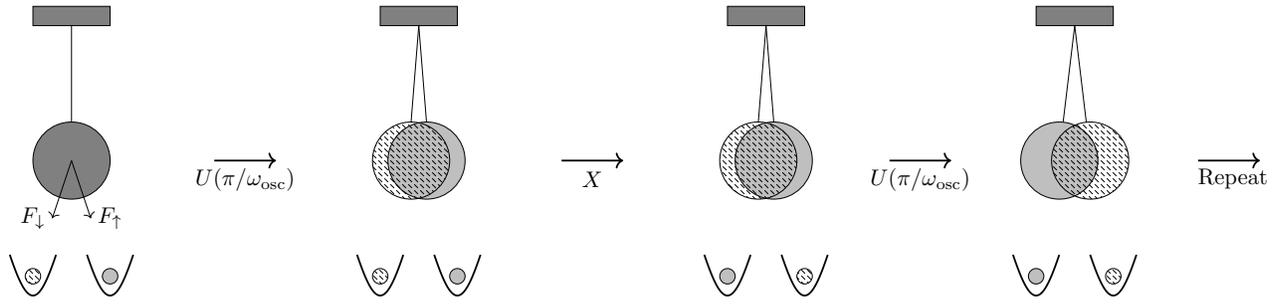


FIG. 7. The spin-echo variant of the basic protocol. After a half period of evolution, the two pathways through oscillator phase space are maximally distant. The atomic positions are then flipped (X gate), followed by another half period of evolution. This procedure can be repeated arbitrarily, leading to a net amplification of the basic protocol. The bottom-right figure shows the resulting conditioned paths of the oscillator through phase space, through one iteration.

$D(4N_\pi\lambda\sigma_z)$. Performing N_π iterations, followed by a σ_x operation, followed by N_π further iterations, we recover the revival:

$$\begin{aligned} U_{\text{spin-echo}} &= \sigma_x D(4N_\pi\lambda\sigma_z) \sigma_x D(4N_\pi\lambda\sigma_z) \\ &= D(-4N_\pi\lambda\sigma_z) D(4N_\pi\lambda\sigma_z) \\ &= \mathbb{I}. \end{aligned} \quad (\text{E2})$$

In this spin-echo-style variant, the wave-function overlap after a total time $t = N_\pi\tau$, i.e., to the halfway point, is given by

$$O = \exp\left[-32 \frac{N_\pi^2 g^2}{\omega^2}\right]. \quad (\text{E3})$$

Thus we have an effect scaling like $g^2/\omega_{\text{eff}}^2$, with the effective frequency $\omega_{\text{eff}} = 2\pi/t = \omega/N_\pi$. This means that we can have an effectively slow oscillator (which is beneficial for the signal strength) while using a faster physical oscillator (beneficial for noise reasons), at the cost of having to perform some σ_x operations on the two-state system.

We note that application of this spin-echo protocol would violate the Markov assumption used to prove the theorem in Sec. III. Adjustment of that proof to accommodate this specific type of non-Markovian control would be necessary to draw the same conclusion, namely, that the gravitational entanglement in this protocol is necessarily due to the mass-atom entangling interaction.

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