# Locality and heating in periodically driven, power-law-interacting systems 

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#### Abstract

We study the heating time in periodically driven $D$-dimensional systems with interactions that decay with the distance $r$ as a power law $1 / r^{\alpha}$. Using linear-response theory, we show that the heating time is exponentially long as a function of the drive frequency for $\alpha>D$. For systems that may not obey linear-response theory, we use a more general Magnus-like expansion to show the existence of quasiconserved observables, which imply exponentially long heating time, for $\alpha>2 D$. We also generalize a number of recent state-of-the-art LiebRobinson bounds for power-law systems from two-body interactions to $k$-body interactions and thereby obtain a longer heating time than previously established in the literature. Additionally, we conjecture that the gap between the results from the linear-response theory and the Magnus-like expansion does not have physical implications, but is, rather, due to the lack of tight Lieb-Robinson bounds for power-law interactions. We show that the gap vanishes in the presence of a hypothetical, tight bound.


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## I. INTRODUCTION

Periodically driven systems can host interesting nonequilibrium physics, such as Floquet topological insulators [1], time crystals [2,3], and anomalous Floquet phases [4]. However, most driven systems eventually heat up to equilibrium, infinite-temperature states, erasing the interesting features in the process.

The timescale before heating becomes appreciable in periodically driven systems is known as the heating time, and it generally exhibits a nontrivial dependence on the frequency of the drive, $\omega$. Previous works [5-8] established that finiterange interacting systems under rapid, local [9], periodic drives could not heat up until after a time $t_{*}=e^{O(\omega)}$ that is exponentially long in the drive frequency $\omega$. This slow heating rate stems at least in part from the locality of the interactions, which constrains the probability that distant particles collectively absorb an energy quantum $\hbar \omega$.

This result also applies to systems with long-range interactions that decay with the distance $r$, e.g., as a power law $1 / r^{\alpha}$. Such systems are of great interest, as they can be implemented in a wide variety of experiments, such as trapped ions [10,11], Rydberg atoms [12], ultracold atoms and molecules [13,14], nitrogen-vacancy centers [15], and superconducting circuits [16]. On the theoretical side, for spin systems with disordered, sign-changing power-law couplings, Ref. [17] demonstrated the exponentially suppressed heating rate when $\alpha>D / 2$, where $D$ is the dimensionality of the system. Furthermore, Ref. [8] proved an exponential heating time $t_{*}=e^{O(\omega)}$ for general power-law interactions with $\alpha>2 D$. In contrast, for $D<\alpha<2 D$, Ref. [8] only obtained a linear heating time
$t_{*}=O(\omega)$, while numerical evidence [18] suggests that the heating time is still exponential within this regime of $\alpha$.

In this paper, we study the heating time in periodically driven, power-law-interacting systems with $\alpha>D$ from two different perspectives. Within linear-response theory, we show that such systems only heat up after some time exponentially large as a function of the drive frequency. This result mirrors the statement established for finite-range interactions in Ref. [5] and extends Ref. [17] to systems without disorder (though at the expense of a smaller range of valid $\alpha$ ). The result also matches the numerical evidence in Ref. [18]. For generic periodically driven, power-law-interacting systems that may not obey the linear-response theory-such as those under a strong drive-we generalize Ref. [6] and construct an effective time-independent Hamiltonian with power-law interactions. This Hamiltonian closely describes the dynamics of the driven system up to time $t_{*}$, where $t_{*}$ is exponentially large as a function of the drive frequency. We thereby show that the system cannot heat up until at least after this timescale.

We note that, although our generalization of Ref. [6] is different from Ref. [8], it is similar in spirit to their arguments. While Ref. [8] mainly focused on finite-range interactions, their construction of the effective Hamiltonian by truncating the Magnus series would also apply to power-law systems. However, our approach here also provides insights into the structure of the effective Hamiltonian. In particular, we show that the effective Hamiltonian is also power-law with the same exponent $\alpha$ as the driven Hamiltonian. Furthermore, we prove a stronger, albeit still exponential in $\omega$, bound on the heating time than one would get from the argument in Ref. [8]. This improvement relies on the use of state-of-the-art

Lieb-Robinson (LR) bounds [19,20], which we develop for this purpose. In particular, through a different technique, we generalize the bound in Ref. [20] from two-body to manybody interactions.

Similarly to Ref. [8], our construction requires $\alpha>2 D$, in contrast to the numerical evidence in Ref. [18] and to the wider range of validity $\alpha>D$ found in the linear-response theory. Because both Ref. [8] and this paper crucially rely on Lieb-Robinson bounds to prove that the heating time is at least exponential in $\omega$, we conjecture that the aforementioned gap stems from the lack of a tight Lieb-Robinson bound for $\alpha>D$, and we show the gap would vanish if such a tight bound were to exist. While the linear-response theory also utilizes Lieb-Robinson bounds, it has weaker assumptions and, therefore, does not require a tighter bound to achieve the desired result of exponentially suppressed heating for $\alpha>D$.

The remainder of the paper is organized as follows. In Sec. II, we provide definitions and describe the systems of interest. In Sec. III, we review various Lieb-Robinson bounds for power-law systems and extend two of them-including one with the tightest light cone known to date-to $k$-body interactions. In Sec. IV, we prove that in the linear-response regime the heating time is at least exponential in $\omega$ for all $\alpha>$ $D$. In Sec. V, we provide a more general analysis using the Magnus-like expansion and existing Lieb-Robinson bounds to prove exponentially long heating times for $\alpha>2 D$. We also conjecture a tight Lieb-Robinson bound that would extend this range of validity to $\alpha>D$. Finally, we summarize and discuss potential improvements in Sec. VI.

## II. SETUP AND DEFINITIONS

We consider a system of $N$ spins in a $D$-dimensional square lattice [21]. The system evolves under a periodic, timedependent Hamiltonian $H(t)$ with period $T$, i.e., $H(t+T)=$ $H(t)$. While the following analysis works for any $H(t)$ that is a sum of finite-body interactions, we assume that $H(t)$ consists of only two-body interactions for simplicity. Without loss of generality, we can write $H(t)=H_{0}+V(t)$ as the sum of a time-independent part $H_{0}$ and a time-dependent part $V(t)$ such that $\frac{1}{T} \int_{0}^{T} V(t)=0$. We further assume that $H_{0}$ and $V(t)$ are both power-law Hamiltonians with an exponent $\alpha$.

Definition 1. A Hamiltonian $H$ on a lattice $\Lambda$ is powerlaw with an exponent $\alpha$ and a local energy scale $\eta$ if we can write $H=\sum_{X} h_{X}$, where $h_{X}$ are Hamiltonians supported on subsets $X \subset \Lambda$, such that for any two distinct sites $i, j \in \Lambda$

$$
\begin{equation*}
\sum_{X \ni i, j}\left\|h_{X}\right\| \leqslant \frac{\eta}{\operatorname{dist}(i, j)^{\alpha}} \tag{1}
\end{equation*}
$$

and the norm $\left\|h_{\{i\}}\right\| \leqslant \eta$ for all $i \in \Lambda$, where $\|\cdot\|$ denotes the operator norm and $\operatorname{dist}(i, j)$ denotes the distance between sites $i, j$. In addition, we call $\sup _{X}|X|$ the local support size, where $|X|$ is the number of sites in $X$, and define $\|H\|_{l}=$ $\sup _{i} \sum_{X \ni i}\left\|h_{X}\right\|$ to be the local norm of $H$.

In the following discussion, we assume $\eta=1$, which sets the timescale for the dynamics of the system. In addition, we will occasionally write $H$ instead of $H(t)$ for brevity.

## III. LIEB-ROBINSON BOUNDS FOR MANY-BODY POWER-LAW INTERACTIONS

Before discussing the linear-response theory and the Magnus-like expansion, it is helpful to review the existing LR bounds for power-law interactions. We will also generalize several of them from two-body to arbitrary $k$-body interactions for all $k \geqslant 2$. In particular, we discuss the relations between the bounds in Refs. [22,23], which imply logarithmic light cones for all $\alpha>D$, and the bounds in Refs. [19,20,24], which imply algebraic light cones for $\alpha>2 D$.

## A. Lieb-Robinson bounds for $\boldsymbol{\alpha}>\boldsymbol{D}$

First, we discuss the bounds in Refs. [22,23], which are valid for all $\alpha>D$. Recall that LR bounds are upper bounds on the norm of the commutator $[A(t), B]$, where $A, B$ are two operators supported on some subsets $X, Y$ of the lattice, and $A(t)$ is the time-evolved version of $A$ under a time-dependent Hamiltonian $H$. The minimum distance between a site in $X$ and a site in $Y$ is $r=\operatorname{dist}(X, Y)>0$. Since the sets $X, Y$ are disjoint, $[A(0), B]=0$ initially. As time grows, the operator $A(t)$ may spread to $Y$, making the commutator nontrivial.

The first LR bound for power-law interactions was proven in Ref. [22] by Hastings and Koma (HK):

$$
\begin{equation*}
\mathcal{C}(t, r) \equiv\|[A(t), B]\| \leqslant C\|A\|\|B\||X||Y| \frac{e^{v t}}{r^{\alpha}} \tag{2}
\end{equation*}
$$

where $r=\operatorname{dist}(X, Y), v \propto \eta$ is a constant that may depend on $\alpha$, and $C$ is a constant independent of the system. We shall also use the same $C$ to denote different inconsequential prefactors. Setting the commutator norm to a constant yields the light cone $t \gtrsim \log r$, which means it takes time at least proportional to $\log r$ for the commutator to reach a given constant value.

Technically, we can already use the HK bound in our later analysis of the heating time because it applies to $k$-body interactions for all $k$. However, this bound is loose for large $\alpha$ for two reasons: (i) the velocity $v \propto 2^{\alpha}$ diverges for $\alpha \rightarrow \infty$ and (ii) the light cone is logarithmic for all $\alpha$, which is unphysical since larger values of $\alpha$ correspond to shorter-range interactions and, therefore, we expect slower spreading of correlations. In particular, we expect the light cone to become linear for large enough $\alpha$, given that the interactions are finite range at $\alpha=\infty$.

Gong et al. [23] resolved the first challenge and derived a bound for two-body interactions:

$$
\begin{equation*}
\mathcal{C}(t, r) \leqslant C\|A\|\|B\||X \| Y|\left(\frac{e^{v t}}{[(1-\mu) r]^{\alpha}}+e^{v t-\mu r}\right) \tag{3}
\end{equation*}
$$

where $\mu \in(0,1)$ is an arbitrary constant. The light cone implied by this bound is still logarithmic, but the velocity $v$ is finite for all $\alpha$. Although the bound in Ref. [23] was derived only for two-body interactions, their proof applies to arbitrary $k$-body interactions, where $k$ is a finite integer (see Appendix A for a proof).

## B. Lieb-Robinson bounds for $\alpha>2 D$

In this section, we discuss the LR bounds for power-law interactions with $\alpha>2 D$. While the bounds in Refs. [22,23] work for $\alpha>D$, they all have logarithmic light cones. For
$\alpha>2 D$, it is possible to derive tighter algebraic light cones. The first such bound was proven by Foss-Feig et al. [24] for two-body interactions (and generalized by Refs. [19,25] to $k$-body interactions for all $k \geqslant 2$ ). A recent bound by Tran et al. [20], however, gives a tighter algebraic light cone. Here, we provide the generalization of that bound to $k$-body interactions:
$\mathcal{C}(t, r) \leqslant C\|A\|\|B\|\left(r_{0}+r\right)^{D-1}\left(\frac{1}{(1-\mu)^{\alpha}} \frac{t^{\alpha-D}}{r^{\alpha-D-1}}+t e^{-\frac{\xi r}{\tau}}\right)$,
where $r_{0}$ is the radius of the smallest ball that contains $X$ and $\mu, \xi \in(0,1)$ are arbitrary constants. The second term decays exponentially with $r / t$ and becomes negligible compared to the first term when $r \gg t$. Note that, other than its dependence on $r_{0}$, this bound is independent of the size of $X, Y$ and is valid for $\alpha>2 D$.

Before we present the proof of Eq. (4), we summarize the key steps of the proof.
(1) First, divide $[0, t]$ into $M$ equal time intervals and define $t_{0}, t_{1}, \ldots, t_{M}$ such that $t_{0}=0$ and $t_{j+1}-t_{j}=\tau=t / M$. We denote by $U_{t_{i}, t_{j}}$ the evolution unitary of the system from time $t_{i}$ to $t_{j}$.
(2) Setting $U_{j}=U_{t_{M-j}, t_{M-j+1}}$ for brevity, we can decompose the evolution of $A$ into $M$ time steps:

$$
\begin{equation*}
A(t)=U_{M}^{\dagger} U_{M-1}^{\dagger} \ldots U_{1}^{\dagger} A U_{1} \ldots U_{M-1} U_{M} \tag{5}
\end{equation*}
$$

(3) We then use a truncation technique (explicitly described below) to approximate $U_{1}^{\dagger} A U_{1}$ by some operator $A_{1}$ such that

$$
\begin{equation*}
\left\|U_{1}^{\dagger} A U_{1}-A_{1}\right\|=\varepsilon_{1} \tag{6}
\end{equation*}
$$

and $A_{1}$ is supported on a ball of size at most $\ell$ larger than the size of the support of $A$.
(4) Repeat the above approximation for the other time slices, i.e., find $A_{2}, \ldots, A_{M}$ such that

$$
\begin{align*}
& \left\|U_{2}^{\dagger} A_{1} U_{2}-A_{2}\right\|=\varepsilon_{2}  \tag{7}\\
& \left\|U_{3}^{\dagger} A_{2} U_{3}-A_{3}\right\|=\varepsilon_{3}  \tag{8}\\
& \ldots  \tag{9}\\
& \left\|U_{M}^{\dagger} A_{M-1} U_{M}-A_{M}\right\|=\varepsilon_{M}
\end{align*}
$$

By the end of this process, we have approximated $A(t)$ by an operator $A_{M}$ the support of which is at most $M \ell$ larger than the support of $A$.
(5) By choosing $M \ell$ just smaller than $r$, the support of $A_{M}$ does not overlap with the support of $B$. Therefore, $\left[A_{M}, B\right]=$ 0 , and $\mathcal{C}(t, r)$ is at most the total error of the approximation, i.e.,

$$
\begin{equation*}
\varepsilon=\varepsilon_{1}+\cdots+\varepsilon_{M} \tag{10}
\end{equation*}
$$

The total error $\varepsilon$, and hence the bound, depends on the truncation technique used in step 3. In Ref. [20], the authors used a technique inspired by digital quantum simulation, which works for $\alpha>2 D$. However, in addition to truncating the evolution unitary, the technique in Ref. [20] also truncates the Hamiltonian. The large error from this truncation makes it difficult to further improve the bound. Here, we use a different, simpler technique to generalize the bound in Ref. [20]
to $k$-body interactions for all $k \geqslant 2$. Our technique does not require truncating the Hamiltonian, eliminating a hurdle for future improvements on the bound [26].

Let us start without any assumption on the interactions of the system. We only assume that there already exists a bound on the commutator norm for the system:

$$
\begin{equation*}
\mathcal{C}(t, r) \leqslant f(t, r) \phi(X)\|A\|\|B\|, \tag{11}
\end{equation*}
$$

for some function $f$ that increases with $t$ and decreases with $r$, where $\phi(X)$ is the boundary area of $X$.

To truncate $U_{1}^{\dagger} A U_{1}$, we simply trace out the part of $U_{1}^{\dagger} A U_{1}$ that lies outside a ball of radius $\ell$ around the support of $A$ [27]:

$$
\begin{align*}
A_{1} & \equiv \frac{1}{\operatorname{Tr}\left(\mathbb{I}_{\left.\mathcal{B}_{\ell}(A)^{c}\right)}\right.} \operatorname{Tr}_{\mathcal{B}_{\ell}(A)^{c}}\left(U_{1}^{\dagger} A U_{1}\right) \otimes \mathbb{I}_{\mathcal{B}_{\ell}(A)^{c}}  \tag{12}\\
& =\int_{\mathcal{B}_{\ell}(A)^{c}} d \mu(W) W\left(U_{1}^{\dagger} A U_{1}\right) W^{\dagger} \tag{13}
\end{align*}
$$

where $\mathcal{B}_{\ell}(A)$ is a ball of radius $\ell+r_{0}$ centered on $A$ and $X^{c}$ denotes the complement of the set $X$. In Eq. (13), we rewrite the trace over $\mathcal{B}_{\ell}(A)^{c}$ as an integral over the unitaries $W$ supported on $\mathcal{B}_{\ell}(A)^{c}$ and $\mu(W)$ denotes the Haar measure for the unitaries. Effectively, $A_{1}$ is the part of $A$ that lies inside the ball $\mathcal{B}_{\ell}(A)$. The error from approximating $U_{1}^{\dagger} A U_{1}$ with $A$ is

$$
\begin{align*}
\varepsilon_{1} & =\left\|U_{1}^{\dagger} A U_{1}-A_{1}\right\| \\
& =\left\|U_{1}^{\dagger} A U_{1}-\int_{\mathcal{B}_{\ell}(A)^{c}} d \mu(W) W\left(U_{1}^{\dagger} A U_{1}\right) W^{\dagger}\right\| \\
& =\left\|\int_{\mathcal{B}_{\ell}(A)^{c}} d \mu(W)\left[U_{1}^{\dagger} A U_{1}-W\left(U_{1}^{\dagger} A U_{1}\right) W^{\dagger}\right]\right\| \\
& \leqslant \int_{\mathcal{B}_{\ell}(A)^{c}} d \mu(W)\left\|\left[U_{1}^{\dagger} A U_{1}, W\right]\right\| . \tag{14}
\end{align*}
$$

Note that $W$ is a unitary the support of which is at least a distance $\ell$ from the support of $A$. Therefore, using the LR bound in Eq. (11), we have

$$
\begin{align*}
\varepsilon_{1} & =\left\|U_{1}^{\dagger} A U_{1}-A_{1}\right\| \leqslant \int_{\mathcal{B}_{\ell}(A)^{c}} d \mu(W)\|A\| \phi(X) f(\tau, \ell) \\
& =\|A\| \phi(X) f(\tau, \ell) \tag{15}
\end{align*}
$$

where $\tau$ is the time interval of each time slice. In addition, it is clear from the definition of $A_{1}$ in Eq. (13) that $\left\|A_{1}\right\| \leqslant\|A\|$. Therefore, the error of the approximation in the $j$ th time slice is at most

$$
\begin{equation*}
\varepsilon_{j} \leqslant\|A\| \phi\left(X_{j-1}\right) f(\tau, \ell) \tag{16}
\end{equation*}
$$

where $X_{j}$ is the support of $A_{j}$. Thus, the new bound is

$$
\begin{align*}
\mathcal{C}(t, r) & \leqslant 2\|B\| \varepsilon \leqslant 2 M\|A\|\|B\| \phi_{\max } f(\tau, \ell)  \tag{17}\\
& =2\|A\|\|B\| \frac{t}{\tau} \phi_{\max } f(\tau, \ell) \tag{18}
\end{align*}
$$

where $\phi_{\max }=\max _{j} \phi\left(X_{j}\right)$ and $M$ has been replaced by $t / \tau$. Note that the above bound is valid for all choices of $t, \ell$, as
long as

$$
\begin{gather*}
M=\frac{t}{\tau}<\frac{r}{\ell}  \tag{19}\\
\ell \geqslant 1  \tag{20}\\
\tau \leqslant t \tag{21}
\end{gather*}
$$

The first condition ensures that the operator after the last time slice $A_{M}$ is still outside the support of $B$, while the last two are practical constraints.

Equation (19) is equivalent to $\ell<r \tau / t$. Because $f(\tau, \ell)$ is a decreasing function of $\ell$, the bound Eq. (18) would be the tightest if we chose $\ell=\xi r \tau / t$ for some $\xi$ less than, but very close to, 1 . The bound Eq. (18) becomes

$$
\begin{equation*}
\mathcal{C}(t, r) \leqslant 2\|A\|\|B\| \phi_{\max } f\left(\tau, \frac{\xi r \tau}{t}\right) \frac{t}{\tau} \tag{22}
\end{equation*}
$$

Note that the only free parameter left is $\tau$, which is constrained by [see Eqs. (19)-(21)]

$$
\begin{equation*}
t \geqslant \tau>\frac{t}{r} \tag{23}
\end{equation*}
$$

We are now ready to generalize the bound in Ref. [20] to many-body interactions. Plugging the $k$-body generalization of Eq. (3) [see Eq. (A13) in Appendix A] into Eq. (22), we have

$$
\begin{aligned}
\mathcal{C}(t, r) \leqslant & C\|A\|\|B\| \phi_{\max } \frac{t}{\tau} \\
& \times\left(\frac{1}{(1-\mu)^{\alpha}} \frac{e^{v \tau}}{\left(\frac{\xi r \tau}{t}\right)^{\alpha-D-1}}+e^{v \tau-\frac{\xi r \tau}{t}}\right) \\
\leqslant & C\|A\|\|B\|\left(r_{0}+r\right)^{D-1} \frac{t}{\tau} \\
& \times\left[\frac{1}{(1-\mu)^{\alpha}} \frac{e^{\nu \tau}}{\tau^{\alpha-D-1}}\left(\frac{t}{r}\right)^{\alpha-D-1}+e^{v \tau-\frac{\xi r \tau}{t}}\right]
\end{aligned}
$$

where we have assumed without loss of generality that $X$ is a ball of radius $r_{0}$ and replaced $\phi_{\max } \propto\left(r_{0}+r\right)^{D-1}$. Taking $\tau=1$ to be a constant, we obtain a bound that is valid for all $\alpha>D+1$ :
$\mathcal{C}(t, r) \leqslant C\|A\|\|B\|\left(r_{0}+r\right)^{D-1}\left(\frac{1}{(1-\mu)^{\alpha}} \frac{t^{\alpha-D}}{r^{\alpha-D-1}}+t e^{-\frac{t r}{t}}\right)$.
In particular, if $r_{0}$ is a constant, we can simplify (in the limit of large $t, r$ ) to

$$
\begin{equation*}
\mathcal{C}(t, r) \leqslant C\|A\|\|B\|\left(\frac{1}{(1-\mu)^{\alpha}} \frac{t^{\alpha-D}}{r^{\alpha-2 D}}+t r^{D-1} e^{-\frac{\xi r}{t}}\right) \tag{24}
\end{equation*}
$$

Note that although the bound is, in principle, valid for $\alpha>$ $D+1$, it is only useful for $\alpha>2 D$.

## IV. LINEAR-RESPONSE THEORY

In this section, we present the derivation of an exponentially suppressed heating rate for periodically driven, powerlaw Hamiltonians under the assumptions of linear-response theory. We will assume that the drive $V(t)$ is harmonic and local. That is, we can write $V(t)=g \cos (\omega t) O$, for some small constant $g$ and some time-independent operator $O=\sum_{i} O_{i}$
composed of local operators $O_{i}$. For simplicity, we assume each $O_{i}$ is supported on a single site $i$ (but our results also hold when $O_{i}$ is supported on a finite number of sites around $i$ ). We also assume the system is initially in a thermal state $\rho_{\beta}$ of $H_{0}$ with a temperature $\beta^{-1}$. Within the linear-response theory, the energy absorption rate is proportional to the dissipative (imaginary) part of the response function $\sigma(\omega)=\sum_{i, j} \sigma_{i j}(\omega)$ [5], where

$$
\begin{equation*}
\sigma_{i j}(\omega)=\frac{1}{2} \int_{-\infty}^{\infty} d t e^{i \omega t}\left\langle\left[O_{i}(t), O_{j}(0)\right]\right\rangle_{\beta} \tag{25}
\end{equation*}
$$

$\langle O\rangle_{\beta} \equiv \operatorname{Tr}\left(\rho_{\beta} O\right)$ denotes the thermal average of $O$, and $O(t)=e^{i H_{0} t} O e^{-i H_{0} t}$ is the time-evolved version of $O$ under $H_{0}$.

The authors of Ref. [5] showed that there exists a constant $\kappa$ such that for all $i, j$ and for all $\omega, \delta \omega>0$ the $(i, j)$ entry of $\sigma(\omega)$ can be bounded as

$$
\begin{equation*}
\left|\sigma_{i j}([\omega, \omega+\delta \omega])\right| \leqslant e^{-\kappa \omega} \tag{26}
\end{equation*}
$$

where $f\left(\left[\omega_{1}, \omega_{2}\right]\right) \equiv \int_{\omega_{1}}^{\omega_{2}} f(\omega) d \omega$. Although the statement of Ref. [5] applies to Hamiltonians with finite-range interactions, we show in Appendix B1 that it also holds for power-law Hamiltonians for all $\alpha \geqslant 0$.

In principle, Eq. (26) already implies that the absorption rate of a finite system is exponentially small as a function of the frequency $\omega$. However, since there are $N$ sites in the system, naively applying Eq. (26) by summing over the indices $i, j$ yields a superextensive heating rate $\sim N^{2} e^{-\kappa \omega}$. Such superextensivity is nonphysical, as it would imply that a local drive instigates a divergent absorption per site in the thermodynamic limit. To address this, Ref. [5] introduced a bound complementary to Eq. (26)-based on Lieb-Robinson bounds for finite-range interactions [28]-that implies the contribution from the off-diagonal terms is also exponentially suppressed with the distance $r_{i j}$ between the sites $i, j$.

The case of power-law-interacting Hamiltonians is somewhat more involved. Due to the long-range interaction, the commutator $\left\langle\left[O_{i}(t), O_{j}\right]\right\rangle_{\beta}$ can decay more slowly as a function of $r_{i j}$ than in the finite-range case. Fortunately, we show that it still decays sufficiently quickly for us to recover the extensive, exponentially small heating rate for power-law Hamiltonians. We provide the technical proof in Appendix B 2, but a high-level argument goes as follows.

Lieb-Robinson bounds for power-law systems with $\alpha>D$ [19,20,22,23] imply that the contributions from the $(i, j)$ entries are suppressed by $1 / r_{i j}^{\alpha}$. Therefore, the total contribution to $\sigma([\omega, \omega+\delta \omega])$ from the pairs $(i, j)$ with $r_{i j}$ larger than some distance $r_{*}$ (to be chosen later) is at most

$$
\begin{equation*}
\sum_{i, j: r_{i j} \geqslant r_{*}} \frac{C}{r_{i j}^{\alpha}} \leqslant \frac{C N}{r_{*}^{\alpha-D}} \tag{27}
\end{equation*}
$$

where we use the same notation $C$ to denote different constants that are independent of $r_{i j}, t$, and $N$. The factor $N$ comes from summing over $i$ and the factor $r_{*}^{D}$ comes from summing over $j$ at least a distance $r_{*}$ from $i$.

For $r_{i j} \leqslant r_{*}$, we simply use the bound in Eq. (26) to bound their contributions:

$$
\begin{equation*}
\sum_{i, j: r_{i j} \leqslant r_{*}} C e^{-\kappa \omega} \leqslant C N r_{*}^{D} e^{-\kappa \omega} \tag{28}
\end{equation*}
$$

where $N r_{*}^{D}$ is roughly the number of pairs $(i, j)$ separated by distances less than $r_{*}$. Combining Eq. (27) with Eq. (28), we get $|\sigma([\omega, \omega+\delta \omega])| \leqslant C N r_{*}^{D}\left(e^{-\kappa \omega}+r_{*}^{-\alpha}\right)$. Finally, choosing $r_{*}=\exp (\kappa \omega / \alpha)$, we obtain a bound on the absorption rate:

$$
\begin{equation*}
|\sigma([\omega, \omega+\delta \omega])| \leqslant C N \exp \left[-\left(1-\frac{D}{\alpha}\right) \kappa \omega\right] \tag{29}
\end{equation*}
$$

which decays exponentially quickly with $\omega$ as long as $\alpha>D$. Thus, we have shown that, within the linear-response theory, the heating rate of power-law-interacting Hamiltonians obeys a bound that is qualitatively similar to that for finite-range interactions: the heating rate is extensive, but exponentially small in the driving frequency.

## V. MAGNUS-LIKE EXPANSION

We now present a more general approach to proving a bound on the heating time in a system governed by a periodically driven, power-law Hamiltonian. In particular, this approach remains correct for strongly driven systems, where linear-response theory does not apply. We generalize Ref. [6] and construct an effective time-independent Hamiltonian $H_{*}$. The leading terms of $H_{*}$ resemble the effective Hamiltonian one would get from the Magnus expansion [29-31]. Using Lieb-Robinson bounds for power-law interactions, we show that the evolution of local observables under $H_{*}$ well approximates the exact evolution up to time $t_{*}$, which is exponentially long as a function of the drive frequency. Additionally, the existence of the effective Hamiltonian $H_{*}$ also implies a prethermalization window during which the system could thermalize with respect to $H_{*}$ before eventually heating up after time $t_{*}$.

Following Ref. [6], we construct a periodic unitary transformation $Q(t)$ such that $Q(t+T)=Q(t)$ and $Q(0)=\mathbb{I}$. After moving into the frame rotated by $Q(t)$, we show that the transformed Hamiltonian is nearly time independent and the norm of the residual time-dependent part is exponentially small as a function of the frequency.

To construct the unitary $Q(t)$, we note that the state of the system in the rotated frame, $|\phi(t)\rangle=Q^{\dagger}(t)|\psi(t)\rangle$, obeys the Schrödinger equation with a transformed Hamiltonian $H^{\prime}(t)$ $(\hbar=1)$ :

$$
\begin{equation*}
i \partial_{t}|\phi(t)\rangle=\left(Q^{\dagger} H Q-i Q^{\dagger} \partial_{t} Q\right)|\phi(t)\rangle \equiv H^{\prime}(t)|\phi(t)\rangle \tag{30}
\end{equation*}
$$

We write $Q=e^{\Omega}$, where $\Omega(t)$ is a periodic, anti-Hermitian operator, i.e., $\Omega(t)=\Omega(t+T)$ and $\Omega^{\dagger}=-\Omega$. We then assume that the period $T$ is small so that we may expand $\Omega(t)=\sum_{q=1}^{\infty} \Omega_{q}$ in orders of $T$, where $\left\|\Omega_{q}\right\|=O\left(T^{q}\right)$, and we will eventually choose $\Omega_{q}$ such that the transformed Hamiltonian $H^{\prime}(t)$ is almost time-independent. In particular, we shall truncate the expansion of $\Omega(t)$ up to order $q_{\text {max }}$ and choose $\Omega_{q}$ recursively for all $q \leqslant q_{\text {max }}$ to minimize the norm of the driving term in $H^{\prime}(t)$.

We can rewrite $H^{\prime}(t)$ from Eq. (30) as

$$
\begin{equation*}
H^{\prime}(t)=e^{-\mathrm{ad}_{\Omega}}\left[H_{0}+V(t)\right]-i \frac{1-e^{-\mathrm{ad}_{\Omega}}}{\operatorname{ad}_{\Omega}} \partial_{t} \Omega \tag{31}
\end{equation*}
$$

with $\operatorname{ad}_{\Omega} A=[\Omega, A]$. From Eq. (31), we can define $H_{q}^{\prime}(t)$ for $q=0,1, \ldots$ such that $H^{\prime}=\sum_{q=0}^{\infty} H_{q}^{\prime}(t)$ is expanded in
powers of $T$ :

$$
\begin{equation*}
H_{q}^{\prime}(t)=G_{q}(t)-i \partial_{t} \Omega_{q+1}(t) \tag{32}
\end{equation*}
$$

where we define $G_{q}$ via $\Omega_{1}, \ldots, \Omega_{q}$ as follows:

$$
\begin{align*}
G_{q}(t)= & \sum_{k=1}^{q} \frac{(-1)^{k}}{k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\
i_{1}+\cdots+i_{k}=q}} \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} H(t) \\
& +i \sum_{k=1}^{q} \frac{(-1)^{k+1}}{(k+1)!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q+1 \\
i_{1}+\cdots+i_{k}+m=q+1}} \\
& \times \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} \partial_{t} \Omega_{m}, \tag{33}
\end{align*}
$$

and $G_{0}(t)=H(t)$. Now, recall that $\Omega_{q}(t)$ are operators that we can choose. From Eq. (32), we choose $\Omega_{1}(t)$ such that it cancels out the time-dependent part of $G_{0}(t)$, making $H_{0}^{\prime}$ time independent. This choice of $\Omega_{1}(t)$ also defines $G_{1}(t)$. We then choose $\Omega_{2}(t)$ to eliminate the time-dependent part of $G_{1}(t)$. In general, we choose $\Omega_{q}$ successively from $q=1$ to some $q=$ $q_{\text {max }}$ (to be specified later) so that $H_{q}^{\prime}$ are time independent for all $q<q_{\max }$. Therefore, the remaining time-dependent part of the transformed Hamiltonian $H^{\prime}(t)$ must be at least $O\left(T^{q_{\text {max }}}\right)$. Specifically, for $q<q_{\max }$, we choose the following:

$$
\begin{gather*}
\bar{H}_{q}^{\prime}=\frac{1}{T} \int_{0}^{T} G_{q}(t) d t  \tag{34}\\
\Omega_{q+1}(t)=-i \int_{0}^{t}\left(G_{q}\left(t^{\prime}\right)-\bar{H}_{q}^{\prime}\right) d t^{\prime} \tag{35}
\end{gather*}
$$

Here, Eq. (35) ensures that Eq. (32) becomes $H_{q}^{\prime}(t)=\bar{H}_{q}^{\prime}$, and, thus, that $H_{q}^{\prime}$ is time independent for all $q<q_{\text {max }}$. On the other hand, for $q \geqslant q_{\max }$, we choose $\Omega_{q+1}(t)=0$, so that $H_{q}^{\prime}(t)=$ $G_{q}(t)$. By this construction, we can rewrite the transformed Hamiltonian into the sum of a time-independent Hamiltonian $H_{*}$ and a drive $V^{\prime}(t)$ that contains higher orders in $T$ :

$$
\begin{equation*}
H^{\prime}(t)=\sum_{q=0}^{\infty} H_{q}(t)=\underbrace{\sum_{q=0}^{q_{\max }-1} \bar{H}_{q}^{\prime}}_{\equiv H_{*}}+\underbrace{\sum_{q=q_{\max }}^{\infty} G_{q}(t)}_{\equiv V^{\prime}(t)} \tag{36}
\end{equation*}
$$

As a result of the transformation, the driving term $V^{\prime}(t)$ is now $O\left(T^{q_{\text {max }}}\right)$. As discussed before, we will eventually choose the cutoff $q_{\text {max }}$ to minimize the norm of the residual drive $\left\|V^{\prime}(t)\right\|$.

To estimate the norm of $V^{\prime}(t)$, elucidating its dependence on $q_{\text {max }}$, we first need more information on the structure of the $\Omega_{q}(t)$ for all $1 \leqslant q \leqslant q_{\text {max }}$. In particular, we show that $G_{q}$ and $\Omega_{q}$ are power-law-interacting Hamiltonians. To do so, we first need to define some more notation. We denote by $\mathbb{H}_{\alpha}$ the set of power-law Hamiltonians with exponent $\alpha$ and a local energy scale $\eta=1$. In addition, we denote by $\mathbb{H}_{\alpha}^{(k)}$ the subset of $\mathbb{H}_{\alpha}$ which contains all power-law Hamiltonians the local support size of which (see Definition 1) is at most $k+1$. For a real positive constant $a$, we also denote by $a \mathbb{H}_{\alpha}$ the set of Hamiltonians $H$ such that $a^{-1} H$ is a power-law Hamiltonian with the same exponent $\alpha$.

The following lemma says that $G_{q}$ and $\Omega_{q}$ are also powerlaw Hamiltonians up to a prefactor.

Lemma 1. For all $q<q_{\max }$, we have

$$
\begin{gather*}
G_{q} \in T^{q} q!c^{q} \lambda^{q} \mathbb{H}_{\alpha}^{(q+1)},  \tag{37}\\
\partial_{t} \Omega_{q+1} \in T^{q} q!c^{q} \lambda^{q} \mathbb{H}_{\alpha}^{(q+1)},  \tag{38}\\
\Omega_{q+1} \in T^{q+1} q!c^{q} \lambda^{q} \mathbb{H}_{\alpha}^{(q+1)}, \tag{39}
\end{gather*}
$$

where $c, \lambda$ are constants to be defined later.
Observe that for any order $q$ the last two bounds, i.e., Eqs. (38) and (39), follow immediately from Eq. (37) and the definition of $\Omega_{q}$. Note that Lemma 1 holds for $G_{0}(t)=$ $H(t) \in \mathbb{H}_{\alpha}^{(1)}$. It is also straightforward to prove Lemma 1 inductively on $q$. The factor $T^{q}$ comes from the constraint in Eq. (33) that $i_{1}+\cdots+i_{k}=q$, along with the fact that each $\Omega_{i_{v}}$ is $O\left(T^{i_{v}}\right)$ for all $v=1, \ldots, k$. Similarly, the factor of $q$ ! is combinatorial and comes from the nested commutators in Eq. (33). We provide a more technical proof of Lemma 1 in Appendix C.

As a consequence of Lemma 1, we can bound the local norms of the operators:

$$
\begin{gather*}
\left\|G_{q}\right\|_{l} \leqslant T^{q} q!c^{q} \lambda^{q+1} \leqslant \lambda e \sqrt{q}\left(\frac{T q c \lambda}{e}\right)^{q}  \tag{40}\\
\left\|\Omega_{q}\right\|_{l} \leqslant T^{q}(q-1)!c^{q-1} \lambda^{q} \leqslant \frac{e}{c}\left(\frac{T q c \lambda}{e}\right)^{q} \tag{41}
\end{gather*}
$$

There are two competing factors in the bounds: $T^{q}$, which decreases with $q$, and $q!\sim q^{q}$, which increases with $q$. This suggests that the optimal choice for $q_{\max }$ —in order to minimize the local norm in Eq. (40)-should be around $e /(c T \lambda)$. In the following, we shall choose

$$
\begin{equation*}
q_{\max }=\omega_{*} \equiv \frac{e}{c T \lambda} e^{-\kappa} \tag{42}
\end{equation*}
$$

for some $\kappa>\ln 2$. Note also that $\omega_{*}$ is equal to frequency $\omega=1 / T$ up to a constant. With this choice of $q_{\max }$, Eq. (37) reduces to

$$
\begin{equation*}
G_{q} \in \lambda e \sqrt{q} e^{-\kappa q} \mathbb{H}_{\alpha}^{(q+1)}, \tag{43}
\end{equation*}
$$

for all $q<q_{\max }=\omega_{*}$. By summing over $G_{q}$ with $q<\omega_{*}$, we find that the effective time-independent Hamiltonian $H_{*}$ [see Eq. (36)] is also a power-law Hamiltonian, i.e., $H_{*} \in$ $C \mathbb{H}_{\alpha}^{\left(q_{\text {max }}\right)} \in C \mathbb{H}_{\alpha}$, up to a constant $C$ that may depend only on $\kappa$.

Similarly, we find from Eq. (39) that $\Omega_{q} \in e /(c \lambda) e^{-\kappa q} \mathbb{H}_{\alpha}^{(q)}$ for all $q \leqslant \omega_{*}$. Plugging into the definition of $G_{q}$ and noting that we choose $\Omega_{q}=0$ for all $q \geqslant q_{\max }$, we find an identity similar to Eq. (43), but for $q \geqslant \omega_{*}$ :

$$
\begin{equation*}
G_{q} \in C e^{-\kappa^{\prime} q} \mathbb{H}_{\alpha} \tag{44}
\end{equation*}
$$

where $\kappa^{\prime}>\kappa-\ln 2$ is a constant. Summing over $G_{q}$ with $q \geqslant q_{\text {max }}$ [see Eq. (36)], we again find that the residual drive $V^{\prime}(t)$ is a power-law Hamiltonian up to a prefactor that decays exponentially with $\omega_{*}$ :

$$
\begin{equation*}
V^{\prime}(t) \in C e^{-\kappa^{\prime} \omega_{*}} \mathbb{H}_{\alpha} \tag{45}
\end{equation*}
$$

where $C$ and $\kappa^{\prime}$ are some positive constants. As a result, the local norm of $V^{\prime}(t)$ decreases exponentially with $\omega_{*}$ : $\left\|V^{\prime}(t)\right\|_{l} \leqslant C \lambda e^{-\kappa^{\prime} \omega_{*}}$.

As discussed earlier, Eqs. (36) and (45) imply the existence of an effective time-independent Hamiltonian $H_{*}$ such that the difference $\left\|Q^{\dagger} H Q-H_{*}\right\|=\left\|V^{\prime}\right\|$ is exponentially small as a function of $\omega_{*} \propto 1 / T$. However, even if $\left\|V^{\prime}\right\|_{l}$ is exponentially small, $\left\|V^{\prime}\right\|$ still diverges in the thermodynamic limit. Therefore, in order to characterize the heating rate of the Hamiltonian, it is necessary to investigate the evolution of a local observable $O$ under $H(t)$. We show that the evolution is well described by the effective time-independent Hamiltonian $H_{*}$ at stroboscopic times $t=T \mathbb{Z}$. Without loss of generality, we assume the local observable $O$ is supported on a single site and $\|O\|=1$. Following a similar technique used in Abanin et al. [6], we write the difference between the approximate evolution under the effective Hamiltonian and the exact evolution (in the rotated frame):

$$
\begin{aligned}
\delta & =Q(t) U^{\dagger}(t) O U(t) Q^{\dagger}(t)-e^{i t H_{*}} O e^{-i t H_{*}} \\
& =i \int_{0}^{t} d s W^{\dagger}(s, t)\left[V^{\prime}(s), e^{i s H_{*}} O e^{-i s H_{*}}\right] W(s, t),
\end{aligned}
$$

where $U(t)=\mathcal{T} \exp \left[-i \int_{0}^{t} H\left(t^{\prime}\right) d t^{\prime}\right]$ is the time evolution generated by the full Hamiltonian $H(t)$ and $W(s, t)=$ $\mathcal{T} \exp \left[-i \int_{s}^{t} H^{\prime}\left(t^{\prime}\right) d t^{\prime}\right]$ is the evolution from time $s$ to $t$ generated by $H^{\prime}(t)$. We can then bound the norm of the difference using the triangle inequality:

$$
\begin{equation*}
\|\delta\| \leqslant \int_{0}^{t} d s\left\|\left[V^{\prime}(s), e^{i s H_{*}} O e^{-i s H_{*}}\right]\right\| . \tag{46}
\end{equation*}
$$

We can bound the right-hand side of Eq. (46) using LiebRobinson bounds for power-law interactions.

First, we provide an intuitive explanation why the norm of $\delta$ is small for small time. Recall that the operator $O$ is initially localized on a single site. At small time, it is still quasilocal and therefore significantly noncommutative with only a small number of terms of $V^{\prime}$ lying inside the "light cone" generated by the evolution under $H_{*}$. There are several Lieb-Robinson bounds for power-law interactions [19,20,23,24] [see also Eqs. (A13) and (4)]; each provides a different estimate for the shape of the light cone, resulting in a different bound for the heating time.

If the light cone is logarithmic (as bounded in Ref. [23]), the commutator norm in Eq. (46) would grow exponentially quickly with time and eventually negate the exponentially small factor $\exp \left(-\kappa^{\prime} \omega_{*}\right)$ from $\left\|V^{\prime}\right\|_{l}$. Therefore, in such cases, the system could potentially heat up only after $t_{*} \propto \omega_{*}=$ $1 / T$. On the other hand, if we use the Lieb-Robinson bounds that imply algebraic light cones (as in Refs. [19,20,24] for $\alpha>2 D$ ), the commutator norm only grows subexponentially with time, and we can expect to recover the exponentially long heating time $t_{*} \propto e^{\kappa^{\prime} \omega_{*}}$ derived for finite-range interactions [6,7].

Appendix D contains the mathematical details, but the results of this analysis are as follows. Using the bound in Gong et al. [23] [or its $k$-body generalization Eq. (A12)], which holds for $\alpha>D$ and has a logarithmic light cone $t \gtrsim \log r$, yields

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} e^{2 D v t / \alpha} \tag{47}
\end{equation*}
$$

Thus, the difference $\delta$ is only small for time $t_{*} \propto \omega_{*} \propto 1 / T$. This behavior is expected because the region inside the light
cone implied by Gong et al.'s bound expands exponentially quickly with time.

If instead we use the bound in Else et al. [19], we find

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} \xi\left(\frac{D}{1-\sigma}\right) t^{\frac{D}{1-\sigma}+1} \tag{48}
\end{equation*}
$$

where $\xi(x) \equiv \frac{1}{x} 2^{x} \Gamma(x)$ and $\Gamma$ is the Gamma function. Thus, the difference is small up to an exponentially long time $t_{*} \propto$ $e^{\kappa^{\prime} \omega_{*} \frac{1-\sigma}{++1-\sigma}}$. The result holds for $\alpha>D\left(1+\frac{1}{\sigma}\right)$, where $\sigma$ can be chosen arbitrarily close to 1 . This condition is effectively equivalent to $\alpha>2 D$ (see Appendix D for a discussion of the limit $\sigma \rightarrow 1^{-}$).

We may also use the bound in Tran et al. [20] [see Eq. (24) for its generalization to $k$-body interactions], which gives

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} t^{\frac{D(\alpha-D)}{\alpha-2 D}+1} \tag{49}
\end{equation*}
$$

Thus, the difference is small up to an exponentially long time $t_{*} \propto \exp \left(\kappa^{\prime} \omega_{*} \frac{\alpha-2 D}{\alpha(D+1)-D(D+2)}\right)$. This analysis works only when $\alpha>3 D$, but, within this regime, the exponent of the heating time using this bound is larger than obtained in Eq. (48). This is due to the tradeoff between the tail and the light cone between the bounds in Refs. [19,20]. See Appendix D for more details.

Finally, we conjecture a tight bound for power-law interactions that holds for all $\alpha>D$, and we will provide the full derivation of $\delta$ for such a bound. First, we consider the light cone of such a bound. Given the best known protocols for quantum information transfer [32], the best light cone we could hope for would be $t \gtrsim r^{\alpha-D}$ for $D+1>\alpha>D$ and linear for $\alpha>D+1$. In the following, we assume the light cone of the conjectured bound is $t \gtrsim r^{1 / \beta}$ for some constant $\beta \geqslant 1$ for all $\alpha>D$.

Next, we consider the tail of the bound, i.e., how the conjectured bound decays with the distance at a fixed time. Since it is always possible to signal between two sites using their direct interaction, which is of strength $1 / r^{\alpha}$, the tail of the bound cannot decay faster than $1 / r^{\alpha}$. We shall assume that the bound decays with the distance exactly as $1 / r^{\alpha}$.

For simplicity, we assume that the conjectured bound takes the form

$$
\begin{equation*}
\|[A(t), B]\| \leqslant C\|A\|\|B\|\left(\frac{t^{\beta}}{r}\right)^{\alpha} \tag{50}
\end{equation*}
$$

which manifestly has a light cone $t \gtrsim r^{1 / \beta}$ and decays as $1 / r^{\alpha}$ with the distance. Let $r_{*}(t)=t^{\beta}$ be the light cone boundary and consider the sum inside and outside the light cone.

For convenience, denote $V^{\prime \prime}=C^{-1} e^{\kappa^{\prime} \omega_{*}} V^{\prime}, \bar{H}^{\prime \prime}=\gamma^{-1} H_{*}$ so that $V^{\prime \prime}, \bar{H}^{\prime \prime} \in \mathbb{H}_{\alpha}$. We can rewrite the bound on $\|\delta\|$ as

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} \int_{0}^{t} d s\left\|\left[V^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \tag{51}
\end{equation*}
$$

Now write $\quad V^{\prime \prime}(s)=\sum_{r=0}^{\infty} V_{r}^{\prime \prime}(s)$, where $\quad V_{r}^{\prime \prime}(s) \equiv$ $\sum_{X: \operatorname{dist}(X, O) \in[r, r+1)} h_{X}$ denotes the terms of $V^{\prime \prime}(s)$ supported on subsets exactly a distance between $r$ and $r+1$ away from $O$. Since $V^{\prime \prime}(s)$ is a power-law Hamiltonian, it follows that $\left\|V_{r}^{\prime \prime}(t)\right\| \leqslant C r^{D-1}$. Writing the sum this way, we can now separate terms inside and outside of the light cone.

For the terms inside the light cone, we bound

$$
\begin{align*}
& \sum_{r \leqslant r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{A}^{\prime \prime}} O e^{-i s \gamma \bar{A}^{\prime \prime}}\right]\right\| \\
& \quad \leqslant 2 \sum_{r \leqslant r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\| \leqslant C r_{*}(s)^{D} \leqslant C s^{\beta D} . \tag{52}
\end{align*}
$$

For the terms outside the light cone, we use the conjectured bound:

$$
\begin{align*}
& \sum_{r>r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \quad \leqslant C \sum_{r>r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\| \frac{s^{\beta \alpha}}{r^{\alpha}}  \tag{53}\\
& \quad \leqslant C \sum_{r>r_{*}(s)} \frac{s^{\beta \alpha}}{r^{\alpha-D+1}} \leqslant C \frac{s^{\beta \alpha}}{r_{*}(s)^{\alpha-D}}=C s^{\beta D} . \tag{54}
\end{align*}
$$

Combining Eqs. (52) and (54), we get

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} t^{\beta D+1} \tag{55}
\end{equation*}
$$

which implies an exponential heating time as a function of $\omega_{*}$, i.e., $t_{*} \propto \exp \left[\kappa^{\prime} \omega_{*} /(\beta D+1)\right]$. Recall that the best values of $\beta$ for which we can hope are $\beta=1 /(\alpha-D)$ when $D+$ $1>\alpha>D$ and $\beta=1$ when $\alpha>D+1$. Note also that the exponential heating time would hold for all $\alpha>D$, matching the result given by the linear-response theory.

## VI. CONCLUSION AND OUTLOOK

Our paper generalizes the results of Refs. [5-7] for finiterange interactions to power-law interactions. Using two independent approaches, we show that periodically driven, powerlaw systems with a large enough exponent $\alpha$ can only heat up after time that is exponentially long in the drive frequency. The results only hold if $\alpha$ is larger than some critical value $\alpha_{c}$. Physically, the existence of $\alpha_{c}$ coincides with our expectation that power-law interactions with a large enough exponent $\alpha$ are effectively short range.

However, the two approaches imply different values for $\alpha_{c}$. While both the Magnus expansion in Ref. [8] and the Magnuslike expansion in this paper independently suggest $\alpha_{c}=2 D$, the linear-response theory implies $\alpha_{c}=D$. We conjecture that this gap is due to the lack of tighter Lieb-Robinson bounds for power-law interactions, especially for $\alpha$ between $D$ and $2 D$. Indeed, we demonstrated in Sec. V that a tight LiebRobinson bound for this range of $\alpha$ implies an exponentially long heating time for all $\alpha>D$, matching the result from the linear-response approach, as well as previous numerical evidence for systems with $\alpha<2 D$ [18]. Therefore, proving a tight Lieb-Robinson bound has important implications for the heating time of power-law-interacting systems.

Note added. Recently, we became aware of a related complementary work on long-range prethermal phases [33]. We also became aware of a tighter Lieb-Robinson bound for power-law interactions [34]. However, the bound has a range of validity $\alpha>2$ in one dimension and, thus, does not close the aforementioned gap.

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## APPENDIX A: GENERALIZATION OF GONG et al. [23] TO MANY-BODY INTERACTIONS

In this section, we prove Eq. (3) and thereby generalize the bound in Gong et al. [23] from two-body to $k$-body interactions, where $k$ is an arbitrary finite integer. This bound is an ingredient in the generalization of the tighter Lieb-Robinson bound in Tran et al. [20] to $k$-body interactions.

Proof. We recall that the bound in Ref. [23] is based on the Hastings and Koma series [22]:

$$
\begin{equation*}
\|[A(t), B]\| \leqslant 2\|A\|\|B\| \sum_{k=1}^{\infty} \frac{(2 t)^{k}}{k!}\left[\sum_{Z_{1}: Z_{1} \cap X \neq \varnothing} \sum_{Z_{2}: Z_{1} \cap Z_{2} \neq \varnothing} \ldots \sum_{\substack{Z_{k}: Z_{k-1} \cap Z_{k} \neq \varnothing, Z_{k} \cap Y \neq \varnothing}} \prod_{i=1}^{k}\left\|h_{Z_{i}}\right\|\right], \tag{A1}
\end{equation*}
$$

and we can bound the summation within the square brackets as

$$
\begin{align*}
\sum_{Z_{1}: Z_{1} \cap X \neq \varnothing} \sum_{Z_{2}: Z_{1} \cap Z_{2} \neq \varnothing} \ldots \sum_{\substack{Z_{k}: Z_{k-1} \cap Z_{k} \neq \varnothing, Z_{k} \cap Y \neq \varnothing}} \prod_{i=1}^{k}\left\|h_{Z_{i}}\right\| & \leqslant \sum_{i \in X} \sum_{j \in Y} \sum_{z_{1}} \sum_{z_{2}} \cdots \sum_{z_{k-1}}\left(\sum_{Z_{1} \ni i, z_{1}}\left\|h_{Z_{1}}\right\|\right) \cdots\left(\sum_{z_{k} \ni z_{k-1}, j}\left\|h_{Z_{k}}\right\|\right) \\
& \leqslant \sum_{i \in X} \sum_{j \in Y} \lambda^{k} \mathcal{J}^{k}(i, j) \tag{A2}
\end{align*}
$$

where $\mathcal{J}^{k}(i, j)$ is given by the $k$-fold convolution of the hopping terms $J_{i j} \equiv \frac{1}{r_{i j}^{\alpha}}$ [where $\left.r_{i j}=\operatorname{dist}(i, j)\right]$ for $i \neq j$ and $J_{i i}=1$ for all $i$ :

$$
\mathcal{J}^{k}(i, j) \equiv \sum_{z_{1}} \sum_{z_{2}} \cdots \sum_{z_{k-1}} J_{i z_{1}} J_{z_{1} z_{2}} \ldots J_{z_{k-1} j} .
$$

Note that Eq. (A2) comes from Definition 1: $\sum_{Z \ni i, j}\left\|h_{Z}\right\| \leqslant 1 / r_{i j}^{\alpha}=J_{i j}$ for $i \neq j$ and

$$
\begin{equation*}
\sum_{Z \ni i}\left\|h_{Z}\right\| \leqslant \sum_{j} \sum_{Z \ni i, j}\left\|h_{Z}\right\| \leqslant \lambda \tag{A3}
\end{equation*}
$$

where $\lambda=1+\sum_{j \neq i} 1 / r_{i j}^{\alpha}$ is a finite constant for all $\alpha>D$. This equation is exactly Eq. (3) in Ref. [23].
For simplicity, we consider $D=1$ in the following discussion. To put a bound on $\mathcal{J}^{k}(i, j)$, we use the same trick as in Ref. [23]. First, we consider the sum over $z_{1}$ :

$$
\begin{equation*}
\sum_{z_{1}} J_{i z_{1}} J_{z_{1} z_{2}} \leqslant 2 \sum_{z_{1}::_{i_{1}} \leqslant r_{z_{1} j}} J_{i z_{1}} J_{z_{1} z_{2}} \tag{A4}
\end{equation*}
$$

where the right-hand side sums only over $z_{1}$ being closer to $i$ than to $z_{2}$ and the factor 2 accounts for exchanging the roles of $i$ and $z_{2}$. We further separate the sum over $z_{1}$ in Eq. (A3) into two, corresponding to whether $z_{1}$ is within a unit distance from $i$ or not:

$$
\begin{equation*}
\sum_{z_{1}} J_{i z_{1}} J_{z_{1} z_{2}} \leqslant 2\left(\sum_{z_{1}: r_{i_{1} 1} \leqslant 1}+\sum_{z_{1}: r_{i_{1}} \geqslant 2}\right) J_{i z_{1}} J_{z_{1} z_{2}} . \tag{A5}
\end{equation*}
$$

Since $r_{i z_{1}} \leqslant r_{z_{1} z_{2}}$, it follows that $r_{z_{1} z_{2}} \geqslant r_{i z_{2}} / 2$. Therefore, $J_{z_{1} z_{2}} \leqslant 2^{\alpha} J_{i z_{2}}$ and we further bound the second sum in Eq. (A5) by

$$
\begin{equation*}
\sum_{z_{1}::_{i_{1}} \geqslant 2} J_{i z_{1}} J_{z_{1} z_{2}} \leqslant 2^{\alpha} J_{i z_{2}} \sum_{z_{1}: r_{i_{1}} \geqslant 2} J_{i z_{1}} \leqslant 2^{\alpha} J_{i z_{2}} 2^{1-\alpha}(\lambda-1) \leqslant 2(\lambda-1) \sum_{z_{1}: r_{i_{1}} \leqslant 1} J_{i z_{1}} J_{z_{1} z_{2}}, \tag{A6}
\end{equation*}
$$

where we bound $\sum_{z_{1}: r_{i z_{1}} \geqslant 2} J_{i z_{1}} \leqslant 2^{1-\alpha}(\lambda-1)$ and $J_{i z_{2}} \leqslant \sum_{z_{1}: r_{i_{1}} \leqslant 1} J_{i_{1} 1} J_{z_{1} z_{2}}$ similarly to Ref. [23]. Therefore, we have $\sum_{z_{1}} J_{i z_{1}} J_{z_{1} z_{2}} \leqslant 4 \lambda \sum_{z_{1}: r_{i_{1} 1} \leqslant 1} J_{i z_{1}} J_{z_{1} z_{2}}$. Repeating this analysis for $z_{2}, \ldots, z_{k}$ in Eq. (A3), we have an upper bound on $\mathcal{J}^{k}(i, j)$ :

$$
\begin{align*}
& \mathcal{J}^{k}(i, j) \leqslant(4 \lambda)^{k-1} \sum_{z_{1}: r_{z_{1}} \leqslant 1} \sum_{z_{2}: r_{z_{12} z_{2}} \leqslant 1} \ldots \quad \ldots \sum_{z_{k-1}: r_{z_{k-2} 2_{k-1}} \leqslant 1} J_{z_{1}} J_{z_{1} z_{2}} \ldots J_{z_{k-1}, j}  \tag{A7}\\
& \leqslant(12 \lambda)^{k-1} \times \begin{cases}1 /\left(r_{i j}-k+1\right)^{\alpha} & \text { if } k<\mu r_{i j}, \\
1 & \text { if } k \geqslant \mu r_{i j},\end{cases}  \tag{A8}\\
& \leqslant(12 \lambda)^{k-1} \times \begin{cases}1 /\left[(1-\mu) r_{i j}\right]^{\alpha} & \text { if } k<\mu r_{i j}, \\
1 & \text { if } k \geqslant \mu r_{i j},\end{cases} \tag{A9}
\end{align*}
$$

where $\mu \in(0,1)$ is an arbitrary constant.
To get the second to last bound, we note that the maximum value that the summand in Eq. (A7) may achieve is $1 /\left(r_{i j}-k+1\right)^{\alpha}$ when $k<\mu r_{i j}$ and 1 when $k \geqslant r_{i j}$, and the number of sites within a unit distance of any site is 3. Plugging this bound into Eqs. (A1) and (A2), we have the Lieb-Robinson bound in Ref. [23] generalized to many-body interactions:

$$
\begin{align*}
\|[A(t), B]\| & \leqslant\|A\|\|B\| \sum_{i \in X} \sum_{j \in Y}\left(\sum_{k=1}^{\left\lceil\mu r_{i j}\right\rceil-1} \frac{\left(24 \lambda^{2} t\right)^{k}}{6 \lambda k!\left[(1-\mu) r_{i j}\right]^{\alpha}}+\sum_{k=\left\lceil\mu r_{i j}\right\rceil}^{\infty} \frac{\left(24 \lambda^{2} t\right)^{k}}{6 \lambda k!}\right) \\
& \leqslant\|A\|\|B\| \sum_{i \in X} \sum_{j \in Y} C e^{v t}\left(\frac{1}{\left[(1-\mu) r_{i j}\right]^{\alpha}}+e^{-\mu r_{i j}}\right)  \tag{A10}\\
& \leqslant\|A\|\|B\||X \| Y| C e^{v t}\left(\frac{1}{[(1-\mu) r]^{\alpha}}+e^{-\mu r}\right) \tag{A11}
\end{align*}
$$

where $C=1 / 6 \lambda, v=24 \lambda^{2}$, and $r$ is, again, the distance between $X, Y$. The proof for $D>1$ follows a very similar analysis.

A feature of Eq. (A11) is that it depends on $|X|,|Y|$, which can become problematic when $A, B$ are supported on a large number of sites. In such cases, we can sum over the sites of $X, Y$ in Eq. (A10) to get more useful bounds. Without any other assumptions, summing over the sites of $Y$ gives an extra factor of $r^{D}$ :

$$
\begin{equation*}
\|[A(t), B]\| \leqslant C\|A\|\|B\||X|\left(\frac{1}{(1-\mu)^{\alpha}} \frac{e^{v t}}{r^{\alpha-D}}+e^{v t-\mu r}\right), \tag{A12}
\end{equation*}
$$

where the constant $C$ absorbs all constants that may depend on $\mu$. Note that the bound still depends on $|X|$ but not on $|Y|$.

We can go one step further and sum over the sites of $X$, but we need to assume that $X$ is convex (similarly to Ref. [20]). Then, we have a bound

$$
\begin{equation*}
\|[A(t), B]\| \leqslant C\|A\|\|B\| \phi(X)\left(\frac{1}{(1-\mu)^{\alpha}} \frac{e^{v t}}{r^{\alpha-D-1}}+e^{v t-\mu r}\right), \tag{A13}
\end{equation*}
$$

which is independent of $|X|$. Here $\phi(X)$ is the boundary area of $X$, defined as the number of sites in $X$ that are adjacent to a site outside $X$.

## APPENDIX B: ABSORPTION RATE FROM LINEAR-RESPONSE THEORY

This section provides more details on the derivation of the absorption rate within linear-response theory. In particular, we provide more mathematically rigorous proofs of Eq. (26) (Appendix B1) and Eq. (29) (Appendix B2).

## 1. Proof of Eq. (26)

In this section, we prove the statement of Eq. (26) [also Eq. (B2) below]. We recall that the system Hamiltonian $H_{0}$ is a power-law Hamiltonian, while the harmonic drive $V(t)=$ $g \cos (\omega t) O$ is a sum of local terms, $g \cos (\omega t) O_{i}$, each of which is supported on the site $i$ only, where $i$ runs over the sites of the system. In addition, we assume that the system is initially in the equilibrium state $\rho_{\beta}$ of $H_{0}$ corresponding to the temperature $1 / \beta$. To the lowest order in $g$, the energy absorption rate of the system is proportional to the dissipative (imaginary) part of the response function, $\sigma(\omega)=\sum_{i, j} \sigma_{i j}(\omega)$, where $i, j$ are the sites of the system and

$$
\begin{equation*}
\sigma_{i j}(\omega)=\frac{1}{2} \int_{-\infty}^{\infty} d t e^{i \omega t}\left\langle\left[O_{i}(t), O_{j}(0)\right]\right\rangle_{\beta}, \tag{B1}
\end{equation*}
$$

where $\langle X\rangle_{\beta} \equiv \operatorname{Tr}\left(\rho_{\beta} X\right)$ denotes the expectation value of an operator $X$ in $\rho_{\beta}$.

In Ref. [5], the authors proved that there exist constants $C, \kappa$ such that for all $\omega>0, \delta \omega>0$ and for all pairs $i, j$

$$
\begin{equation*}
\left|\sigma_{i j}([\omega, \omega+\delta \omega])\right| \leqslant C e^{-\kappa \omega} \tag{B2}
\end{equation*}
$$

The statement in Ref. [5] is for finite-range interactions, but, for completeness, we show here that it also holds for powerlaw Hamiltonians. First, we consider the diagonal terms $\sigma_{i i}(\omega)$. Let $|n\rangle$ and $E_{n}$ denote the eigenstates and eigenvalues of $H_{0}$. Similarly to Ref. [5], we rewrite $\sigma_{i i}(\omega)$ as

$$
\begin{equation*}
\sigma_{i i}(\omega)=\pi \sum_{n} p_{n}\left[\gamma_{i i}^{(n)}(\omega)-\gamma_{i i}^{(n)}(-\omega)\right] \tag{B3}
\end{equation*}
$$

where $p_{n}$ is the probability that the state is in the eigenstate $|n\rangle$, and $\gamma_{i i}^{(n)}$ denotes the contribution to $\sigma_{i i}$ from the $n$th
eigenstate:

$$
\begin{align*}
\gamma_{i i}^{(n)}(\omega) & \left.=\sum_{m}\left|\langle m| O_{i}\right| n\right\rangle\left.\right|^{2} \delta\left(E_{n}-E_{m}-\omega\right) \\
& =\sum_{m} \frac{\left.\left|\langle m| \operatorname{ad}_{H}^{k} O_{i}\right| n\right\rangle\left.\right|^{2}}{\omega^{2 k}} \delta\left(E_{n}-E_{m}-\omega\right), \tag{B4}
\end{align*}
$$

where $\operatorname{ad}_{H} O_{i}=\left[H, O_{i}\right], k$ is an integer to be chosen later, and the last equality comes from the fact that $|m\rangle,|n\rangle$ are eigenstates of $H$ and the $\delta$ function fixes the energy difference to be $\omega$.

In Ref. [5], the authors used the fact that $H$ has a finite range to upper bound the norm of $\operatorname{ad}_{H}^{k} O_{i}$ by $\lambda^{k} k$ ! for some constant $\lambda$. For power-law interactions, the proof does not apply because the Hamiltonian $H$ can contain interaction terms between arbitrarily far sites. Instead, we upper bound $\operatorname{ad}_{H}^{k} O_{i}$ by realizing that $O_{i}$ technically satisfies Definition 1 and is therefore a power-law Hamiltonian. It then follows from Lemma 3 in Appendix E that ad ${ }_{H}^{k} O_{i} \in \lambda^{k} k!\mathbb{H}_{\alpha}$, i.e., $\operatorname{ad}_{H}^{k} O_{i}$ is a power-law Hamiltonian up to a factor $\lambda^{k} k$ !, where $\lambda$ is the same constant as in Lemma 3 and $\mathbb{H}_{\alpha}$ is the set of power-law Hamiltonians with exponent $\alpha$ (see Appendix E1). Finally, we can upper bound

$$
\begin{equation*}
\left\|\operatorname{ad}_{H}^{k} O_{i}\right\| \leqslant C \lambda^{k} k! \tag{B5}
\end{equation*}
$$

by realizing that the supports of the terms in $\operatorname{ad}_{H}^{k} O_{i}$ all contain the site $i$.

Integrating Eq. (B4) over $\omega$, assuming $\delta \omega$ is small enough so that the number of energy levels in the range $[\omega, \omega+\delta \omega$ ] is finite, and using Eq. (B5), we have

$$
\begin{align*}
\left|\gamma_{i i}^{(n)}([\omega, \omega+\delta \omega])\right| & \leqslant C\left(\frac{\lambda^{k} k!}{\omega^{k}}\right)^{2} \\
& \leqslant C\left(\frac{\lambda k}{\omega}\right)^{2 k} \leqslant C e^{-\kappa \omega} \tag{B6}
\end{align*}
$$

where $\kappa=2 /(\lambda e)$ and, to get the last line, we choose $k=$ $\omega /(\lambda e)$. Plugging this bound into Eq. (B3) and summing over $n$ yields Eq. (26) for $i=j$. The bound for $i \neq j$ can be derived using the positivity of $\sigma$ [5] and the Cauchy-Schwarz inequality:

$$
\begin{equation*}
\left|\sigma_{i j}(\omega)\right| \leqslant \frac{1}{2}\left[\sigma_{i i}(\omega)+\sigma_{j j}(\omega)\right] \tag{B7}
\end{equation*}
$$

Therefore, Eq. (26) applies for all power-law Hamiltonians $H$.

## 2. Proof of Eq. (29)

We now provide a rigorous proof of Eq. (29) in the main text. Eq. (B2) indicates that the $(i, j)$ entry of $\sigma([\omega, \omega+\delta \omega])$ is exponentially suppressed. In principle, summing over all $i, j$ implies that $\sigma([\omega, \omega+\delta \omega])$ is also exponentially small as a function of $\omega$. However, since there are $N$ sites in the system, this summation results in an additional factor of $N^{2}$, making $\sigma([\omega, \omega+\delta \omega])$ superextensive. Therefore, this naive calculation breaks down in the thermodynamic limit $(N \rightarrow$ $\infty)$.

Instead, to show that $\sigma([\omega, \omega+\delta \omega])$ increases only as fast as $N$, we use Lieb-Robinson bounds to bound the off-diagonal terms $\sigma_{i j}(\omega)$. Let $r_{i j}=\operatorname{dist}(i, j)$ denote the distance between
the pair of sites $i, j$. Without loss of generality, we assume $\omega \geqslant 2 \delta \omega$. We can then bound

$$
\begin{align*}
\sigma([\omega, \omega+\delta \omega]) & =\int_{\omega}^{\omega+\delta \omega} d \omega^{\prime} \sigma\left(\omega^{\prime}\right) \\
& \leqslant c_{1} \int_{-\infty}^{\infty} d \omega^{\prime} e^{-\left(\frac{\omega^{\prime}-\omega}{\delta \omega}\right)^{2}} \sigma\left(\omega^{\prime}\right) \\
& =c_{2} \delta \omega \sum_{i, j} \int_{-\infty}^{\infty} d t e^{(-t / \delta t)^{2}} e^{-i \omega t}\left\langle\left[O_{i}(t), O_{j}\right]\right\rangle \tag{B8}
\end{align*}
$$

where $c_{1}=\frac{e}{1-e^{-8}}$ and $c_{2}=c_{1} \sqrt{\pi} / 2$, which we will combine and denote by $C$, and $\delta t=2 / \delta \omega$. The first inequality is because $\sigma(\omega)$ is positive for $\omega>0$ and $\sigma(-\omega)=-\sigma(\omega)$. The second equality comes from evaluating the integral over $\omega^{\prime}$. We then use the Lieb-Robinson bound in Ref. [23], which applies for interactions with characteristic exponent $\alpha>D$ :

$$
\begin{equation*}
\left\|\left[O_{i}(t), O_{j}(0)\right]\right\| \leqslant C e^{v t}\left(\frac{1}{r_{i j}^{\alpha}}+e^{-\mu r_{i j}}\right) \tag{B9}
\end{equation*}
$$

where $v, C, \mu$ are positive constants. While this bound was derived in Ref. [23] for two-body interactions, it also holds for more general $k$-body interactions and thus for fully general power-law Hamiltonians [see Eq. (A11)].

We now divide the sum in Eq. (B8) into two parts corresponding to $r_{i j}>r_{*}$ and $r_{i j} \leqslant r_{*}$ for some parameter $r_{*}$ we shall choose later. The sum over $i, j$ such that $r_{i j}>r_{*}$ can then be bounded by first inserting Eq. (B9) into Eq. (B8) and evaluating the integration over time. Note that the factor $e^{-t^{2} / \delta t^{2}}$ suppresses the contribution from $e^{v t}$ at large $t$. Therefore, performing the integral yields an upper bound $C\left(1 / r_{i j}^{\alpha}+\right.$ $\left.e^{-\mu r_{i j}}\right)$ for each term corresponding to the pair $(i, j)$, and the sum over $r_{i j}>r_{*}$ gives

$$
\begin{equation*}
\sum_{i, j: r_{i j}>r_{*}} C\left(\frac{1}{r_{i j}^{\alpha}}+e^{-\mu r_{i j}}\right) \leqslant C N\left(\frac{1}{r_{*}^{\alpha-D}}+e^{-\mu r_{*}}\right) \tag{B10}
\end{equation*}
$$

for $\alpha>D$, where the factor of $N$ comes from summing over $i$ and the factor of $r^{D}$ comes from summing over $j$.

On the other hand, for $r_{i j} \leqslant r_{*}$, we simply use Eq. (B2) to bound their contributions. Summing over $i, j$ such that $r_{i j} \leqslant$ $r_{*}$, we get a bound $C N r_{*}^{D} e^{-\kappa \omega}$, where the factor of $N$ again comes from summing over $i$ and the factor of $r_{*}^{D}$ comes from counting the number of sites $j$ within a distance $r_{*}$ from $i$. Combining with Eq. (B10) yields an upper bound on the total heating rate:

$$
\begin{equation*}
|\sigma([\omega, \omega+\delta \omega])| \leqslant C N r_{*}^{D}\left(e^{-\kappa \omega}+\frac{1}{r_{*}^{\alpha}}+r_{*}^{-D} e^{-\mu r_{*}}\right) \tag{B11}
\end{equation*}
$$

Choosing $r_{*}^{\alpha}=e^{\kappa \omega}$ and noting that the last term is dominated by the first two when $\omega$ is large enough, we find

$$
\begin{equation*}
|\sigma([\omega, \omega+\delta \omega])| \leqslant C N e^{-\frac{\alpha-D}{\alpha} \kappa \omega}, \tag{B12}
\end{equation*}
$$

which is exponentially small with $\omega$ as long as $\alpha>D$.

## APPENDIX C: THE EFFECTIVE HAMILTONIAN

In this section, we study the structure of the effective Hamiltonian defined in Eq. (36). Specifically, we show that the operators $G_{q}$ defined in Eq. (33) are also power-law

Hamiltonians (see also Lemma 1 in the main text for $q<q_{\text {max }}$ and Lemma 2 below for $q \geqslant q_{\max }$ ). In addition, we show that the norm $G_{q}$ for $q \geqslant q_{\text {max }}$ is exponentially small as a function of $q$ and $\omega_{*}$ (Lemma 2), implying that the norm of the residual drive $V^{\prime}$ is also exponentially small.

## 1. Structure of $\boldsymbol{G}_{\boldsymbol{q}}$ for $\boldsymbol{q}<\boldsymbol{q}_{\text {max }}$

First, we prove the statement of Lemma 1 that the operators $G_{q}$ are also power-law Hamiltonians for all $q<q_{\max }$.
Proof. We proceed by induction and assume that Lemma 1 holds for all $q$ up to $q=q_{0}-1$ for some $q_{0} \geqslant 1$. We now prove that it also holds for $q=q_{0}$. We consider the first term in the definition of $G_{q_{0}}$ [Eq. (33)]:

$$
\begin{equation*}
G_{q_{0}, 1}=\sum_{k=1}^{q_{0}} \frac{(-1)^{k}}{k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q_{0} \\ i_{1}+\cdots+i_{k}=q_{0}}} \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} H(t) \tag{C1}
\end{equation*}
$$

Using Lemma 1 (note that it applies to all $i \leqslant q_{0}$ ) and Lemma 3 in Appendix E, we have

$$
\begin{align*}
G_{q_{0}, 1} & \in \sum_{k=1}^{q_{0}} \frac{1}{k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q_{0} \\
i_{1}+\cdots+i_{k}=q_{0}}} T^{q_{0}} c^{q_{0}} \lambda^{q_{0}-k} \prod_{j=1}^{k}\left(i_{j}-1\right)!q_{0}^{k} c^{-k} \lambda^{k} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \\
& =T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} \sum_{k=1}^{q_{0}} \frac{q_{0}^{k} c^{-k}}{k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q_{0} \\
i_{1}+\cdots+i_{k}=q_{0}}} \prod_{j=1}^{k}\left(i_{j}-1\right)!\mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \\
& \subseteq T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} \sum_{k=1}^{q_{0}} \frac{q_{0}^{k} c^{-k}}{k!}\left(q_{0}-k\right)!2^{k} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \subseteq T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} q_{0}!\underbrace{q_{0}!k!}_{\sum_{k=1}^{q_{0}} \underbrace{2^{k} q_{0}^{k} c^{-k}\left(q_{0}-k\right)!}_{\leqslant c_{1}}} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \subseteq c_{1} T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} q_{0}!\mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \tag{C2}
\end{align*}
$$

where $c_{1}$ is a constant which exists because the sum over $k$ converges (see Lemma 5 in Appendix E). To get the first equation, we use Lemma 3, with $k_{\text {max }}$ upper bounded by $q_{0}$ every time. We have also used the second part of Lemma 4 in Appendix E to bound the sum over $i_{1}, \ldots, i_{k}$.

Next, we consider the second term in the definition of $G_{q_{0}}$ :

$$
\begin{equation*}
G_{q_{0}, 2}=i \sum_{k=1}^{q} \frac{(-1)^{k+1}}{(k+1)!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q+1 \\ i_{1}+\cdots+i_{k}+m=q+1}} \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} \partial_{t} \Omega_{m}(t) \tag{C3}
\end{equation*}
$$

Again, we use Lemmas 1 and 3 to show that

$$
\begin{align*}
& G_{q_{0}, 2} \in \sum_{k=1}^{q_{0}} \frac{q_{0}^{k}}{(k+1)!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q_{0}+1 \\
i_{1}+\cdots+i_{k}+m=q_{0}+1}} T^{q_{0}} c^{q_{0}-k-1} \lambda^{q_{0}-k-1} \prod_{j=1}^{k}\left(i_{j}-1\right)!(m-1)!\lambda^{k} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)}  \tag{C4}\\
& =T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} \sum_{k=1}^{q_{0}} \frac{q_{0}^{k} c^{-k}}{(k+1)!} \underbrace{\substack{\begin{subarray}{c}{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q_{0}+1 \\
i_{1}+\cdots+i_{k}+m=q_{0}+1} }}}_{\leqslant 2^{k+1}} \underbrace{\prod_{j=1}^{k}\left(i_{j}-1\right)!(m-1)!\mathbb{H}_{\alpha}^{\left(q_{0}+1\right)}}_{\leqslant\left(q_{0}+1-(k+1)\right)!=\left(q_{0}-k\right)!}  \tag{C5}\\
& \subseteq T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} 2 \sum_{k=1}^{q_{0}} \frac{2^{k} q_{0}^{k} c^{-k}}{(k+1)}\left(q_{0}-k\right)!\mathbb{H}_{\alpha}^{\left(q_{0}+1\right)}  \tag{C6}\\
& \subseteq 2 T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} q_{0}!\underbrace{\sum_{k=1}^{q_{0}} \frac{2^{k} q_{0}^{k} c^{-k}}{k!} \frac{\left(q_{0}-k\right)!}{q_{0}!}}_{\leqslant c_{1}} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \subseteq 2 c_{1} T^{q_{0}} c^{q_{0}} \lambda^{q_{0}} q_{0}!\mathbb{H}_{\alpha}^{\left(q_{0}+1\right)}, \tag{C7}
\end{align*}
$$

where we have used Lemma 4 in Appendix E to bound the sums over $i_{1}, \ldots, i_{k}, m$.

Combining Eqs. (C2) and (C7), we have

$$
\begin{equation*}
G_{q_{0}} \in 3 c_{1} T^{q_{0}} c^{q_{0}} q_{0}!\lambda^{q_{0}} \mathbb{H}_{\alpha}^{\left(q_{0}+1\right)} \tag{C8}
\end{equation*}
$$

Note that $c_{1}$ can be made arbitrarily small by choosing a larger value for $c$. Therefore, with $c$ large enough so that $3 c_{1}<1$, we have that Lemma 1 holds for $q=q_{0}$.

## 2. Structure of $\boldsymbol{G}_{\boldsymbol{q}}$ for $\boldsymbol{q} \geqslant \boldsymbol{q}_{\text {max }}$

We now prove Eq. (44), which is a similar result to Lemma 1, but for $q \geqslant q_{\max }=\omega_{*}$.

Lemma 2. For all $q \geqslant q_{\max }=\omega_{*}, G_{q} \in C e^{-\kappa^{\prime} q} \mathbb{H}_{\alpha}$, where $C$ and $\kappa^{\prime}$ are constants.

Proof. Let us first look at the first term in Eq. (33):

$$
\begin{equation*}
G_{q, 1}=\sum_{k=1}^{q} \frac{(-1)^{k}}{k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant \omega_{*} \\ i_{1}+\cdots+i_{k}=q}} \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} H(t) . \tag{C9}
\end{equation*}
$$

We also recall from Lemma 1 that for all $q \leqslant \omega_{*}$

$$
\begin{equation*}
\Omega_{q} \in T^{q}(q-1)!c^{q-1} \lambda^{q-1} \mathbb{H}_{\alpha}^{(q)} \subseteq \frac{1}{\lambda c q} T^{q} q!c^{q} \lambda^{q} \mathbb{H}_{\alpha} \tag{C10}
\end{equation*}
$$

For all $q \leqslant \omega_{*}$, we have

$$
\begin{equation*}
T^{q} q!c^{q} \lambda^{q} \leqslant(T c \lambda q)^{q} \leqslant\left(T c \lambda \omega_{*}\right)^{q} \leqslant e^{-\kappa q} \tag{C11}
\end{equation*}
$$

where we have used $\omega_{*}=e^{-\kappa} /(T c \lambda)$. Therefore, for all $q \leqslant$ $\omega_{*}$, we have

$$
\begin{equation*}
\Omega_{q} \in \frac{1}{\lambda c q} T^{q} q!c^{q} \lambda^{q} \mathbb{H}_{\alpha} \in \frac{1}{\lambda c q} e^{-\kappa q} \mathbb{H}_{\alpha} \tag{C12}
\end{equation*}
$$

Note also that $H(t) \in \mathbb{H}_{\alpha}$. Therefore, using Lemma 3, we have

$$
\begin{align*}
\operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} H(t) & \in \frac{1}{i_{1} \ldots i_{k}} \frac{q^{k}}{c^{k}} e^{-\kappa q} \lambda^{-k} \lambda^{k} \mathbb{H}_{\alpha} \\
& =\frac{1}{i_{1} \ldots i_{k}} \frac{q^{k}}{c^{k}} e^{-\kappa q} \mathbb{H}_{\alpha} \tag{C13}
\end{align*}
$$

Thus, we get for all $q$

$$
\begin{align*}
G_{q, 1} & \in \underbrace{\left(\sum_{k=1}^{q} \frac{q^{k}}{c^{k} k!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant \omega_{*} \\
i_{1}+\cdots+i_{k}=q}} \frac{1}{i_{1} \ldots i_{k}}\right)}_{\leqslant e^{q / / 2 q}} e^{-\kappa q} \mathbb{H}_{\alpha} \\
& \subseteq e^{-(\kappa-\ln 2-1 / c) q} \mathbb{H}_{\alpha} \tag{C14}
\end{align*}
$$

Note that $i_{j} \leqslant \omega_{*}$ as we only define $\Omega$ up to $\omega_{*}$. Further, the factor of $2^{q}$ comes from upper bounding $\frac{1}{i_{1} \ldots i_{k}}$ with 1 and the number of terms with $2^{q}$. Next, we consider the second term in the definition of $G_{q}$ :

$$
\begin{align*}
G_{q, 2}= & i \sum_{k=1}^{q} \frac{(-1)^{k+1}}{(k+1)!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q+1 \\
i_{1}+\cdots+i_{k}+m=q+1}} \\
& \times \operatorname{ad}_{\Omega_{i_{1}}} \ldots \operatorname{ad}_{\Omega_{i_{k}}} \partial_{t} \Omega_{m}(t) . \tag{C15}
\end{align*}
$$

Note that

$$
\begin{equation*}
\partial_{t} \Omega_{m}(t) \in T^{m-1}(m-1)!c^{m-1} \lambda^{m-1} \mathbb{H}_{\alpha} \subseteq e^{-\kappa(m-1)} \mathbb{H}_{\alpha} \tag{C16}
\end{equation*}
$$

Thus, we have

$$
\begin{align*}
G_{q, 2} & \in\left(\sum_{k=1}^{q} \frac{q^{k}}{c^{k}(k+1)!} \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k}, m \leqslant q+1 \\
i_{1}+\cdots+i_{k}+m=q+1}} e^{-\kappa q}\right) \mathbb{H}_{\alpha} \\
& \subseteq 2 e^{-(\kappa-\ln 2-1 / c) q} \mathbb{H}_{\alpha} . \tag{C17}
\end{align*}
$$

Combining Eqs. (C14) and (C17), we arrive at Lemma 2 with $\kappa^{\prime}=\kappa-\ln 2-1 / c$, which can be made to be positive by choosing $\kappa>\ln 2+1 / c$. It suffices, however, to choose $\kappa>\ln 2$, since making $c$ large enough sends $1 / c$ to zero. Equation (45) also follows.

## APPENDIX D: USING LIEB-ROBINSON BOUNDS FOR EVOLUTIONS OF LOCAL OBSERVABLES

In this section, we use the Lieb-Robinson bounds to bound the norm of $\delta$ in Eq. (46). In the main text, we argue that $\|\delta(t)\|$ would be small up to time $t_{*} \propto \omega_{*}$ if the light cone induced by the Lieb-Robinson bound is logarithmic, and $t_{*} \propto e^{\kappa^{\prime} \omega_{*}}$ if the light cone is algebraic. We provide below the mathematical details to supplement the argument.

Recall that $V^{\prime}(t) \in C e^{-\kappa^{\prime} \omega_{*}} \mathbb{H}_{\alpha}, H_{*} \in \gamma \mathbb{H}_{\alpha}(\gamma$ is a constant that depends only on $\kappa, \alpha$ ) and that we defined the normalized $V^{\prime \prime}=C^{-1} e^{\kappa^{\prime} \omega_{*}} V^{\prime}, \bar{H}^{\prime \prime}=\gamma^{-1} H_{*}$ such that

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} \int_{0}^{t} d s\left\|\left[V^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \tag{D1}
\end{equation*}
$$

We now use a Lieb-Robinson bound for power-law interactions to bound the commutator. The idea is that for a finite time $s$ the operator $O$ mostly spreads within a light cone, and only the terms of $V^{\prime \prime}(s)$ within the light cone significantly contribute to the commutator.

In contrast to the finite-range interacting Hamiltonians, a tight Lieb-Robinson bound has yet to be proven for power-law Hamiltonians with finite $\alpha>D$. In the following sections, we consider the effect of using different Lieb-Robinson bounds, namely, the bounds in Gong et al. [23], Else et al. [19], and Tran et al. [20]. The case of a hypothetical bound, which would be tight if it were proven, is treated in the main text.

## 1. Using Gong et al.'s [23] bound

First, we consider a generalization of the bound in Gong et al. [23] [see also Eq. (A12)]. The bound holds for $\alpha>D$, has a logarithmic light cone $t \gtrsim \log r$, and is extended to many-body interactions. To bound the commutator norm in Eq. (51), recall that we write $V^{\prime \prime}(s)=\sum_{r=0}^{\infty} V_{r}^{\prime \prime}(s)$, where $V_{r}^{\prime \prime}(s) \equiv \sum_{X: \operatorname{dist}(X, O) \in[r, r+1)} h_{X}$ denotes the terms of $V^{\prime \prime}(s)$ supported on subsets exactly a distance between $r$ and $r+1$ away from $O$. Furthermore, since $V^{\prime \prime}(s)$ is a power-law Hamiltonian, it follows that $\left\|V_{r}^{\prime \prime}(t)\right\| \leqslant C r^{D-1}$.

From Eq. (A12), the light cone of the bound is $r_{*}(s)=$ $e^{v s / \alpha}$. We further divide $V_{r}^{\prime \prime}(s)$ into those with $r \leqslant r_{*}(s)$ and $r>r_{*}(s)$. In the former case, we simply bound

$$
\begin{align*}
& \sum_{r \leqslant r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \quad \leqslant 2 \sum_{r \leqslant r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\| \leqslant C r_{*}(s)^{D} \leqslant C e^{D v s / \alpha} . \tag{D2}
\end{align*}
$$

For the latter case, we use Eq. (A12) to bound the commutator norm:

$$
\begin{align*}
& \sum_{r>r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \quad \leqslant C \sum_{r>r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\|\left(\frac{e^{v s}}{r^{\alpha-D}}+e^{v s-\mu r}\right)  \tag{D3}\\
& \leqslant C \sum_{r>r_{*}(s)}\left(\frac{e^{v s}}{r^{\alpha-2 D+1}}+r^{D-1} e^{v s-\mu r}\right)  \tag{D4}\\
& \leqslant C\left(\frac{e^{v s}}{r_{*}(s)^{\alpha-2 D}}+r_{*}(s)^{D-1} e^{v s-\mu r_{*}(s)}\right)  \tag{D5}\\
& \leqslant C\left(e^{2 D v s / \alpha}+e^{v s \frac{D-1}{\alpha}} e^{v s-\mu e^{v s / \alpha}}\right)  \tag{D6}\\
& \leqslant C e^{2 D v s / \alpha} \tag{D7}
\end{align*}
$$

where we use the same $C$ to denote different constants that may depend on $\mu, \alpha$. Note that while the bound in Eq. (A12) is valid for $\alpha>D$ the sum over $r$ converges only when $\alpha>2 D$.

Plugging Eqs. (D2) and (D7) into Eq. (51) and integrating over $s$, we have

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} e^{2 D v t / \alpha} \tag{D8}
\end{equation*}
$$

which is the result presented in Sec. V. Again, $\delta$ is only small for up to time $t_{*} \propto \omega_{*} \propto 1 / T$, which is expected because the region inside the light cone implied by this bound expands exponentially fast with time.

## 2. Using Else et al.'s [19] bound

Instead of using Gong et al.'s [23] bound, we now use the bound in Else et al. [19], which already holds for many-body interactions. The bound states that when $|X|=1$

$$
\begin{equation*}
\|[A(t), B]\| \leqslant C\|A\|\|B\|\left\{\exp \left(v t-r^{1-\sigma}\right)+\frac{(v t)^{1+D /(1-\sigma)}}{r^{\sigma(\alpha-D)}}\right\}, \tag{D9}
\end{equation*}
$$

where $1>\sigma>(D+1) /(\alpha-D+1)$ is a constant that we can choose. Since our aim is to prove an exponential heating time for $\alpha$ as small as possible, we need the algebraic tail exponent $\sigma(\alpha-D)$ to be as large as possible. So we will assume that we pick some $\sigma$ very close to 1 .

First, let us look at the light cone generated by Eq. (D9). The first term of the bound gives a light cone $t \gtrsim r^{1-\sigma}$, while the second term gives $t \gtrsim r^{(1-\sigma)^{\frac{\sigma(\alpha-D)}{D+1-\sigma}} \text {. Since we are choosing }}$ $\sigma$ close to $1, \frac{\sigma(\alpha-D)}{D+1-\sigma}$ will be larger than 1 when $\alpha>2 D$. The former light cone, i.e., $t \gtrsim r^{1-\sigma}$, is therefore looser and thus dominates the latter. In the rest of the calculation, we take $r_{*}(t)=t^{1 /(1-\sigma)}$ to be the light cone boundary.

Similar to Eq. (D2), we get an upper bound for the terms inside the light cone:

$$
\begin{align*}
& \sum_{r \leqslant r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \leqslant 2 \sum_{r \leqslant r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\| \\
& \leqslant C r_{*}(s)^{D} \leqslant C s^{D /(1-\sigma)} \tag{D10}
\end{align*}
$$

For the terms outside the light cone, we use Eq. (D9):

$$
\begin{align*}
& \sum_{r>r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \leqslant \sum_{r>r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\|\left(e^{v s-r^{1-\sigma}}+\frac{(v s)^{1+D /(1-\sigma)}}{r^{\sigma(\alpha-D)}}\right) \\
& \leqslant C \sum_{r>r_{*}(s)}\left(r^{D-1} e^{v s-r^{1-\sigma}}+\frac{(v s)^{1+D /(1-\sigma)}}{r^{\sigma(\alpha-D)-D+1}}\right) \\
& \leqslant C\left(\frac{1}{D} \xi\left(\frac{D}{1-\sigma}\right) e^{v s} r_{*}^{D} e^{-r_{*}^{1-\sigma}}+\frac{(v s)^{1+D /(1-\sigma)}}{r_{*}^{\sigma(\alpha-D)-D}}\right) \\
& \leqslant C\left(\xi\left(\frac{D}{1-\sigma}\right) s^{D /(1-\sigma)}+\frac{(v s)^{1+D /(1-\sigma)}}{s^{\frac{\sigma(\alpha-D)-D}{1-\sigma}}}\right) \\
& \leqslant C \xi\left(\frac{D}{1-\sigma}\right) s^{\frac{D}{1-\sigma}}, \tag{D11}
\end{align*}
$$

where $\xi(x) \equiv \frac{1}{x} 2^{x} \Gamma(x), \Gamma$ is the Gamma function, and we again absorb all constants that may depend on $D$ alone into the constant $C$. We drop the second term in the second to last inequality because for $\sigma$ arbitrarily close to 1 and $\alpha>2 D$ (see below) the second term may be upper bounded by the first. To estimate the sum over $r$, we have used Lemma 6 in Appendix E2. Plugging Eqs. (D10) and (D11) into Eq. (51) and integrating over time, we get

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} \xi\left(\frac{D}{1-\sigma}\right) t^{\frac{D}{1-\sigma}+1} \tag{D12}
\end{equation*}
$$

Thus, the difference is small up to an exponentially long time $t_{*} \propto e^{\kappa^{\prime} \omega_{*} \frac{1-\sigma}{D+1-\sigma}}$. The sum over $r$ converges if $\sigma(\alpha-D)>$ $D$, or equivalently $\alpha>D\left(1+\frac{1}{\sigma}\right)$. Since $\sigma$ can be chosen arbitrarily close to 1 , this condition is effectively equivalent to $\alpha>2 D$.

One should be careful, however; in taking the limit $\sigma$ goes to 1 (i) the heating time $t_{*} \propto e^{\kappa^{\prime} \omega_{*} \frac{1-\sigma}{D+1-\sigma}}$ is no longer exponential in $\omega_{*}$ and (ii) the prefactor $\xi\left(\frac{D}{1-\sigma}\right)$ diverges faster than exponentially in this limit. Nevertheless, the analysis is still valid for fixed values of $\sigma<1$.

## 3. Using Tran et al.'s [20] bound

In addition to Else et al. [19]'s bound, we can also use the bound in Tran et al. [20] [see also Eq. (24) for a generalization to $k$-body interactions], which also works for $\alpha>2 D$. Compared to the bound in Else et al. [19], the bound in Tran et al. [20] has a tighter light cone $r_{*}(s)=s^{(\alpha-D) /(\alpha-2 D)}$, but it decays with the distance $r$ as $r^{\alpha-2 D}$, slower than the tail $r^{\sigma(\alpha-D)}$ in Else et al. [19] when $\sigma>(\alpha-2 D) /(\alpha-D)$.

Similar to before, we further divide $V_{r}^{\prime \prime}(s)$ into those with $r \leqslant r_{*}(s)$ and $r>r_{*}(s)$. For the terms inside the light cone, we again bound

$$
\begin{align*}
& \sum_{r \leqslant r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{A}^{\prime \prime}}\right]\right\| \\
& \quad \leqslant 2 \sum_{r \leqslant r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\| \\
& \quad \leqslant C r_{*}(s)^{D} \leqslant C s^{D(\alpha-D) /(\alpha-2 D)} . \tag{D13}
\end{align*}
$$

For the terms outside the light cone, we use Eq. (24) with $\phi(X)=1$ :

$$
\begin{align*}
& \sum_{r>r_{*}(s)}\left\|\left[V_{r}^{\prime \prime}(s), e^{i s \gamma \bar{H}^{\prime \prime}} O e^{-i s \gamma \bar{H}^{\prime \prime}}\right]\right\| \\
& \leqslant \sum_{r>r_{*}(s)}\left\|V_{r}^{\prime \prime}(s)\right\|\|O\|\left(\frac{s^{\alpha-D}}{r^{\alpha-2 D}}+s r^{D-1} e^{-r / s}\right) \\
& \leqslant C \sum_{r>r_{*}(s)}\left(\frac{s^{\alpha-D}}{r^{\alpha-3 D+1}}+s r^{2 D-2} e^{-\mu r / s}\right) \\
& \leqslant C\left(\frac{s^{\alpha-D}}{r_{*}(s)^{\alpha-3 D}}+s^{2} r_{*}(s)^{2 D-2} e^{-\mu r_{*}(s) / s}\right) \\
& \leqslant C\left(s^{\frac{D(\alpha-D)}{\alpha-2 D}}+s^{2} s^{2(\alpha-D)(D-1) /(\alpha-2 D)} e^{-\mu s^{D /(\alpha-2 D)}}\right) \\
& \leqslant C s^{\frac{D(\alpha-D)}{\alpha-2 D}} \tag{D14}
\end{align*}
$$

where we have dropped the second term in the second to last inequality because it is exponentially small in $s$ and can be upper bounded by the first term. Note that we require $\alpha>3 D$ in order for the sum over $r$ to converge.

Plugging Eqs. (D13) and (D14) into Eq. (51) and integrating over time, we get

$$
\begin{equation*}
\|\delta\| \leqslant C e^{-\kappa^{\prime} \omega_{*}} t^{\frac{D(\alpha-D)}{\alpha-2 D}+1} \tag{D15}
\end{equation*}
$$

Thus, the difference is small up to an exponentially long time $t_{*} \propto e^{\kappa^{\prime} \omega_{*} \frac{\alpha-2 D}{\alpha(D+1)-D(D+2)}}$. Compared to using Else et al.'s [19] bound, this analysis works only when $\alpha>3 D$. However, within this regime, the exponent of the heating time using this bound is larger than using Else et al. This is a manifestation of the tradeoff between the tail and the light cone when switching from Else et al. to the Tran et al. [20] bound.

## APPENDIX E: MATHEMATICAL PRELIMINARIES

This section contains mathematical details omitted from the previous sections for clarity. In Appendix E1, we discuss the properties of the set of power-law Hamiltonians defined in Definition 1. In Appendix E2, we present some bounds on discrete sums.

## 1. Properties of the set $\mathbb{H}_{\boldsymbol{\alpha}}$ of power-law Hamiltonians

In this section, we explore some properties of $\mathbb{H}_{\alpha}$ that are useful for proving that the effective Hamiltonian is also power-law (see Appendix C).

TABLE I. Mutually exclusive indicator functions for Lemma 3. For example, $\xi_{1}=1$ if all of the conditions in the first row, i.e., $i, j \in$ $X$ and $i, j \notin Y$, hold and $\xi_{1}=0$ otherwise.

|  | $\in X$ | $\notin X$ | $\in Y$ | $\notin Y$ |
| :--- | :---: | :---: | :---: | :---: |
| $\xi_{1}$ | $i, j$ |  | $i, j$ |  |
| $\xi_{2}$ | $i, j$ |  | $i$ | $j$ |
| $\xi_{3}$ | $i, j$ |  | $j$ | $i$ |
| $\xi_{4}$ | $i, j$ |  |  | $i, j$ |
| $\xi_{5}$ | $i$ | $j$ | $i, j$ |  |
| $\xi_{6}$ | $i$ | $j$ | $j$ | $i$ |
| $\xi_{7}$ | $j$ | $i$ | $i, j$ |  |
| $\xi_{8}$ | $j$ | $i$ | $i$ | $j$ |
| $\xi_{9}$ |  | $i, j$ | $i, j$ |  |

We recall from the main text that $\mathbb{H}_{\alpha}$ is the set of powerlaw Hamiltonians with the exponent $\alpha$. In addition, $\mathbb{H}_{\alpha}^{(k)}$ is the subset of $\mathbb{H}_{\alpha}$ which contains all power-law Hamiltonians the local support size of which (see Definition 1) is at most $k+1$. For a real positive constant $a$, we also denote by $a \mathbb{H}_{\alpha}$ the set of Hamiltonians $H$ such that $a^{-1} H$ is a power-law Hamiltonian with the exponent $\alpha$. It is straightforward to prove the following identities:

$$
\begin{gather*}
a \mathbb{H}_{\alpha}+b \mathbb{H}_{\alpha} \subset(a+b) \mathbb{H}_{\alpha}  \tag{E1}\\
a \mathbb{H}_{\alpha} \subset b \mathbb{H}_{\alpha} \text { if } a \leqslant b . \tag{E2}
\end{gather*}
$$

The following lemma is particularly useful for the adjoint operation:

Lemma 3. For $\alpha>D$, if $H_{1} \in a \mathbb{H}_{\alpha}^{\left(k_{1}\right)}, H_{2} \in b \mathbb{H}_{\alpha}^{\left(k_{2}\right)}$ for some positive constants $a, b, k_{1}, k_{2}$, then $\operatorname{ad}_{H_{1}} H_{2} \in a b \lambda k_{\max } \mathbb{H}_{\alpha}^{\left(k_{1}+k_{2}\right)}$, where $\lambda$ is a constant to be defined later and $k_{\text {max }}=$ $\max \left\{k_{1}, k_{2}\right\}$.

Proof. Write $H_{1}=\sum_{X} a_{X}, H_{2}=\sum_{Y} b_{Y}, \operatorname{ad}_{H_{1}} H_{2}=$ $\sum_{Z} h_{Z}$, where $h_{Z}=\operatorname{ad}_{h_{X}} h_{Y}$ and $Z=X \cup Y$. By our definition of power-law Hamiltonians, we have

$$
\begin{equation*}
\sum_{X \ni i, j}\left\|a_{X}\right\| \leqslant \frac{a}{\operatorname{dist}(i, j)^{\alpha}}, \quad \sum_{Y \ni i, j}\left\|b_{Y}\right\| \leqslant \frac{b}{\operatorname{dist}(i, j)^{\alpha}} \tag{E3}
\end{equation*}
$$

When $\alpha>D$, it is also straightforward to prove that $\sum_{X \ni i}\left\|a_{X}\right\| \leqslant a \lambda_{0}$ for all $i$, where $\lambda_{0}$ is a constant that depends only on $\alpha, D$.

Note that $h_{Z} \neq 0$ only if $X \cap Y \neq \varnothing$. We seek to bound $\sum_{Z \ni i, j}\left\|h_{Z}\right\|$ which sums over $Z=X \cup Y \ni i, j$ and $X \cap Y \neq$ $\varnothing$. We discuss some useful notations. We will occasionally rewrite or label summations with restrictions using the indicator function $\xi(A)$ where $\xi(A)=1$ or 0 if $A$ is true or false, respectively. There are nine mutually exclusive cases (Table I), satisfying $i, j \in X \cup Y$ depending on whether $i, j$ are in $X, Y$, or both.

Thus, the indicator function $\xi(X \cup Y \ni i, j)$ may be written as a sum of indicator functions of mutually exclusive events listed in the table: $\xi(X \cup Y \ni i, j)=\sum_{n=1}^{9} \xi_{n}$. The overall sum that we want to bound can be written as a sum
over the nine cases,

$$
\begin{align*}
& \sum_{Z \ni i, j}\left\|h_{Z}\right\| \\
& \quad=\sum_{X \cup Y \ni i, j}\left\|\left[a_{X}, b_{Y}\right]\right\| \\
& \quad \leqslant 2 \sum_{X} \sum_{Y}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \xi(X \cup Y \ni i, j) \\
& \quad=2 \sum_{n=1}^{9} \sum_{X} \sum_{Y}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \xi_{n}, \tag{E4}
\end{align*}
$$

and we will bound each of the nine cases individually. We will often eliminate the condition that $X \cap Y \neq \varnothing$, which can only make the sum larger, and introduce an inequality by summing over all sets $X$ or $Y$. To illustrate our technique, consider first the contribution from $\xi_{5}$ :

$$
\begin{align*}
& 2 \sum_{X} \sum_{Y}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \xi_{5} \\
& \quad \leqslant 2 \sum_{X \ni i} \sum_{Y \ni i, j}\left\|a_{X}\right\|\left\|b_{Y}\right\| \\
& \quad \leqslant 2 \sum_{X \ni i}\left\|a_{X}\right\| \frac{b}{\operatorname{dist}(i, j)^{\alpha}} \leqslant \frac{2 \lambda_{0} a b}{\operatorname{dist}(i, j)^{\alpha}} \tag{E5}
\end{align*}
$$

where the first inequality comes from ignoring $j \notin X$ and the second comes from $H_{2}$ being a power-law Hamiltonian.

The bound on the term corresponding to $\xi_{7}$ follows analogously since we simply switch $i, j$. Similarly, the terms corresponding to $\xi_{2}, \xi_{3}$ switch only the roles of $X, Y$ compared to $\xi_{5}, \xi_{7}$. Meanwhile, analyzing the term corresponding to $\xi_{1}$ yields

$$
\begin{align*}
& 2 \sum_{X} \sum_{Y}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \xi_{1} \\
& \quad=2 \sum_{X \ni i, j} \sum_{Y \ni i, j}\left\|a_{X}\right\|\left\|b_{Y}\right\| \\
& \quad \leqslant \frac{2 a b}{\operatorname{dist}(i, j)^{2 \alpha}} \leqslant \frac{2 a b}{\operatorname{dist}(i, j)^{\alpha}} \tag{E6}
\end{align*}
$$

where we take into account $\operatorname{dist}(i, j) \geqslant 1$ for all $D$.
Upper bounding the term corresponding to $\xi_{6}$ is a bit trickier. Since $X \cap Y \neq \varnothing$, there exists a site $\ell \neq i, j$ such that $\ell \in X \cap Y$. Rewriting the term corresponding to $\xi_{6}$ as a sum over $\ell$, we have

$$
\begin{align*}
& \sum_{\substack{X \ni i \\
X \nexists j}} \sum_{Y \ni \nexists j}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \\
& \quad \leqslant 2 \sum_{\ell \neq i, j} \sum_{X \ni i, \ell} \sum_{Y \ni j, \ell} 2\left\|a_{X}\right\|\left\|b_{Y}\right\| \\
& \quad \leqslant 2 \sum_{\ell \neq i, j} \frac{a}{\operatorname{dist}(i, \ell)^{\alpha}} \frac{b}{\operatorname{dist}(\ell, j)^{\alpha}} \leqslant \frac{2 \lambda_{1} a b}{\operatorname{dist}(i, j)^{\alpha}}, \tag{E7}
\end{align*}
$$

where the last inequality comes from the reproducibility condition [22], applicable when $\alpha>D$, and $\lambda_{1}$ is a constant that depends only on $D, \alpha$. The term corresponding to $\xi_{8}$ contributes the same as $\xi_{6}$, as it only switches the roles of $i, j$.

Finally, we bound the terms corresponding to $\xi_{4}, \xi_{9}$. For $\xi_{4}$, we are trying to bound the sum

$$
\begin{equation*}
\sum_{X \ni i, j} \sum_{Y \ngtr i, j} 2\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) . \tag{E8}
\end{equation*}
$$

The nonempty intersection means that for there to be a nonzero contribution $\exists \ell \neq i, j$ such that $\ell \in X, Y$. Further note that by assumption the maximum extent of $X$ is $k_{1}+1$ and therefore there are at most $k_{1}-1$ sites distinct from $i, j$ where $Y$ can intersect with $X$. We bound this as follows:

$$
\begin{align*}
& 2 \sum_{X \ni i, j} \sum_{Y \ngtr i, j}\left\|a_{X}\right\|\left\|b_{Y}\right\| \xi(X \cap Y \neq \varnothing) \\
& \quad \leqslant 2 \sum_{X \ni i, j} \sum_{\substack{\ell \in X \\
\ell \neq i, j}} \sum_{Y \ni \ell}\left\|a_{X}\right\|\left\|b_{Y}\right\| \\
& \quad \leqslant 2 \sum_{X \ni i, j}\left\|a_{X}\right\| \sum_{\substack{\ell \in X \\
\ell \neq i, j}} \lambda_{0} b \leqslant \frac{2 \lambda_{0}\left(k_{1}-1\right) a b}{\operatorname{dist}(i, j)^{\alpha}} . \tag{E9}
\end{align*}
$$

We bound the term corresponding to $\xi_{9}$ similarly by switching the role of $X, Y$. Collecting everything, we have the lemma with $\lambda=2\left(6 \lambda_{0}+2 \lambda_{1}+1\right)$.

## 2. Bounds on discrete sums

In this section, we provide bounds on some discrete sums used in the main text.

Lemma 4. For all $1 \leqslant k \leqslant q$, we have the following inequalities:

$$
\begin{align*}
& \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\
i_{1}+\cdots+i_{k}=q}} \prod_{j=1}^{k} i_{j}!\leqslant \frac{q!}{(k-1)!},  \tag{E10}\\
& \sum_{\substack{0 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\
i_{1}+\cdots+i_{k}=q}} \prod_{j=1}^{k} i_{j}!\leqslant 2^{k} q!. \tag{E11}
\end{align*}
$$

Proof. We first bound

$$
\begin{equation*}
\sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\ i_{1}+\cdots+i_{k}=q}} \prod_{j=1}^{k} i_{j}!\leqslant\binom{ q-1}{k-1} \max _{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\ i_{1}+\cdots+i_{k}=q}} \prod_{j=1}^{k} i_{j}! \tag{E12}
\end{equation*}
$$

For positive integers $a \geqslant b$, we have $(a+b-1)!=a!(a+$ $b-1) \cdots(a+1) \geqslant a!b$ ! with equality if either $a, b=1$. This implies that the maximal product occurs for some $i_{j}=q-$ $k+1$ and $i_{k \neq j}=1$ (we omit the simple proof by induction), yielding

$$
\begin{align*}
& \binom{q-1}{k-1}_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q \\
i_{1}+\cdots+i_{k}=q}} \prod_{j=1}^{k} i_{j}! \\
& \\
& \leqslant \frac{(q-1)!}{(k-1)!(q-k)!}(q-k+1)!  \tag{E13}\\
& \\
& \leqslant \frac{(q-1)!}{(k-1)!}(q-k+1) \leqslant \frac{q!}{(k-1)!}
\end{align*}
$$

as $k \neq 0$ by the summation restrictions. Equation (E11) is essentially the same as Eq. (E10) with some indices allowed to be zero. For example, if $i_{1}=0$ while the other $i$ are nonzero, it is just Eq. (E10) with $k \rightarrow k-1$. This part of the sum is then crudely upper bounded by $q$ !, while summing over all possible choices of zero indices leads to a factor $2^{k}$.

Corollary 1. For all $1 \leqslant k \leqslant q$, we have

$$
\begin{equation*}
\sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q_{0} \\ i_{1}+\cdots+i_{k}=q_{0}}} \prod_{j=1}^{k}\left(i_{j}-1\right)!\leqslant 2^{k}\left(q_{0}-k\right)!. \tag{E14}
\end{equation*}
$$

Proof. Define $p_{j}=i_{j}-1$ such that $0 \leqslant p_{j} \leqslant q_{0}-1$ and $p_{1}+\cdots+p_{k}=q_{0}-k$. This second condition implies that we may simplify the first condition to $0 \leqslant p_{j} \leqslant q_{0}-k$. Therefore,

$$
\begin{align*}
& \sum_{\substack{1 \leqslant i_{1}, \ldots, i_{k} \leqslant q_{0} \\
i_{1}+\cdots+i_{k}=q_{0}}} \prod_{j=1}^{k}\left(i_{j}-1\right)! \\
& =\sum_{\substack{ \\
0 \leqslant p_{1}, \ldots, p_{k} \leqslant q_{0}-k \\
p_{1}+\cdots+p_{k}=q_{0}-k}}^{\substack{j=1}} p_{j}!\leqslant 2^{k}\left(q_{0}-k\right)!, \tag{E15}
\end{align*}
$$

where the last inequality is from Eq. (E11).
Lemma 5. For all $1 \leqslant k \leqslant q$, we have

$$
\begin{equation*}
\sum_{k=1}^{q_{0}} \frac{2^{k} q_{0}^{k} c^{-k}\left(q_{0}-k\right)!}{q_{0}!k!} \leqslant \frac{e}{\sqrt{2 \pi}}\left(e^{2 e / c}-1\right) \tag{E16}
\end{equation*}
$$

Proof. Using Stirling's approximation, $\sqrt{2 \pi} n^{n+\frac{1}{2}} e^{-n} \leqslant$ $n!\leqslant e n^{n+\frac{1}{2}} e^{-n}$ for $q_{0}$ ! and $\left(q_{0}-k\right)!$, we can bound

$$
\begin{align*}
& \sum_{k=1}^{q_{0}} \frac{2^{k} q_{0}^{k} c^{-k}}{k!} \frac{\left(q_{0}-k\right)!}{q_{0}!} \\
& \quad \leqslant \sum_{k=1}^{q_{0}} \frac{2^{k} q_{0}^{k} c^{-k}}{k!} \frac{e}{\sqrt{2 \pi}} \frac{\left(q_{0}-k\right)^{q_{0}-k}}{q_{0}^{q_{0}}} \frac{\sqrt{q_{0}-k}}{\sqrt{q_{0}}} \frac{e^{-\left(q_{0}-k\right)}}{e^{-q_{0}}} \\
& \\
& \leqslant \frac{e}{\sqrt{2 \pi}} \sum_{k=1}^{q_{0}} \frac{2^{k} e^{k} c^{-k}}{k!} \underbrace{\frac{\left(q_{0}-k\right)^{q_{0}-k}}{q_{0}^{q_{0}-k}}}_{\leqslant 1} \underbrace{\frac{\sqrt{q_{0}-k}}{\sqrt{q_{0}}}}_{\leqslant 1}  \tag{E17}\\
&
\end{align*}
$$

We note that the bound approaches zero as $c \rightarrow \infty$.
Lemma 6. For $D \in \mathbb{N}_{>0}, r_{*}>1,0<\eta<1$,

$$
\begin{equation*}
\sum_{r>r_{*}} r^{D-1} e^{-r^{\eta}} \leqslant \frac{2}{\eta} 2^{D / \eta} \Gamma(D / \eta) r_{*}^{D} e^{-r_{*}^{\eta}}, \tag{E18}
\end{equation*}
$$

where $\Gamma$ is the Gamma function.

Proof. Let $f(r)=r^{D-1} e^{-r^{\eta}}$. Our strategy is to upper bound $\sum_{r>r_{*}} f(r)$ by an integral. For $r \in(0, \infty), f$ has a maximum at $r=r_{0}=(D-1)^{1 / \eta} \eta^{-1 / \eta}$. Let $r_{0}^{-}=\left\lfloor r_{0}\right\rfloor$ and $r_{0}^{+}=r_{0}^{-}+1>r_{0}$. Then, the function $f(r)$ is increasing for $r \in\left(r_{*}, r_{0}^{-}\right)$and decreasing for $r \geqslant r_{0}^{+}$. Therefore, we can upper bound

$$
\begin{align*}
\sum_{r>r_{*}} f(r) & \leqslant \int_{r_{*}}^{r_{0}^{-}} f(r) d r+\int_{r_{0}^{+}}^{\infty} f(r) d r+f\left(r_{0}^{-}\right)+f\left(r_{0}^{+}\right) \\
& \leqslant \int_{r_{*}}^{r_{0}^{-}} f(r) d r+\int_{r_{0}^{+}}^{\infty} f(r) d r+2 \int_{r_{0}^{-}}^{r_{0}^{+}} f(r) d r \\
& \leqslant 2 \int_{r_{*}}^{\infty} f(r) d r \tag{E19}
\end{align*}
$$

where we use the fact that $f(r)$ is concave between $r_{0}^{-}$and $r_{0}^{+}$ to bound the first line by the second line. Next, to bound the integral, we make a change of variable to $x=r^{\eta}$ so that

$$
\begin{align*}
2 \int_{r_{*}}^{\infty} f(r) d r & =2 \int_{r_{*}}^{\infty} r^{D-1} e^{-r^{\eta}} d r \\
& =\frac{2}{\eta} \int_{x_{*}}^{\infty} x^{\frac{D-\eta}{\eta}} e^{-x} d x \\
& \leqslant \frac{2}{\eta} \int_{x_{*}}^{\infty} x^{\beta} e^{-x} d x \leqslant \frac{2}{\eta} 2^{\beta} \beta!x_{*}^{\beta} e^{-x_{*}} \\
& =\frac{2}{\eta} 2^{\beta} \beta!r_{*}^{\eta \beta} e^{-r_{*}^{\eta}} \\
& \leqslant \frac{2}{\eta} 2^{D / \eta} \Gamma(D / \eta) r_{*}^{D} e^{-r_{*}^{\eta}} \tag{E20}
\end{align*}
$$

where $x_{*}=r_{*}^{\eta}, \beta=\lceil(D-\eta) / \eta\rceil \leqslant D / \eta$ is an integer, and $\Gamma$ is the Gamma function. Note that we have also used a bound for the integral

$$
\begin{equation*}
\int_{x_{*}}^{\infty} x^{\beta} e^{-x} d x \leqslant 2^{\beta} \beta!x_{*}^{\beta} e^{-x_{*}} \tag{E21}
\end{equation*}
$$

which can be proven inductively on $\beta$ for all $\beta \geqslant 0$ and $x_{*} \geqslant 2$. Indeed, the inequality is trivial for $\beta=0$. Suppose the inequality holds for $\beta-1$; using integration by parts, we have

$$
\begin{align*}
\int_{x_{*}}^{\infty} x^{\beta} e^{-x} d x & =x_{*}^{\beta} e^{-x_{*}}+\beta \int_{x_{*}}^{\infty} x^{\beta-1} e^{-x} d x \\
& \leqslant x_{*}^{\beta} e^{-x_{*}}+\beta 2^{\beta-1}(\beta-1)!x_{*}^{\beta-1} e^{-x_{*}} \\
& \leqslant 2^{\beta-1}\left(\frac{1}{2^{\beta-1} \beta!}+\frac{1}{x_{*}}\right) \beta!x_{*}^{\beta} e^{-x_{*}} \\
& \leqslant 2^{\beta} \beta!x_{*}^{\beta} e^{-x_{*}} \tag{E22}
\end{align*}
$$

where the terms inside the bracket in the second to last line are always less than or equal to 2 for all $x_{*} \geqslant 1$ (corresponding to $r_{*}>1$ ).
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