# Densimetry for primary temperature metrology and a method for the in situ determination of densimeter sinker volumes* 

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#### Abstract

The feasibility of using gas-phase pressure-density-temperature ( $p-\rho-T$ ) measurements as a method of determining thermodynamic temperatures is evaluated. Densities are measured at the unknown and reference temperatures at identical pressures, and the resulting density ratios are extrapolated to zero pressure to yield the ratio of the temperatures. The method utilizes a two-sinker densimeter with a magnetic suspension coupling to separate the sinkers (in the working gas) from the balance (in atmosphere). Numerical studies investigate the optimum working gas and the effects of experimental uncertainties on the uncertainty in temperature. The method is demonstrated using an existing densimeter over the range of $234-505 \mathrm{~K}$ with argon, neon or nitrogen as the working gas. Experimental protocols and data analysis techniques are developed. New experimental $p-\rho-T$ data for argon and neon are reported. The technique is shown to be feasible, although the present densimeter has uncertainties which are too high for temperature metrology. In particular, the uncertainty in the sinker volumes is relatively large. The method is inverted to determine, in situ, the sinker volumes as a function of temperature, reducing their uncertainties, and thus the uncertainties of fluid densities measured with this apparatus.


Keywords: argon, density, gas thermometer, neon, temperature, two-sinker densimeter, volume

## 1. Introduction

Temperature is among the most important quantities in a vast array of applications. Temperature is a welldefined thermodynamic quantity, but the temperature scales in practical use, such as ITS-90, are only approximations of the thermodynamic temperature. In many applications, it is sufficient to use a temperature scale which is merely consistent and reproducible. In thermodynamics, however, it is often important that the temperature scale represents true

[^0]thermodynamic temperatures. Two of the primary sources of thermodynamic temperatures on which ITS-90 is based, namely the gas-thermometer work of Edsinger and Schooley (1989) and Guildner and Edsinger (1976), differ from each other by 30 mK at 730 K . Recent work by Strouse et al (2002) and Ripple et al (2002) provides strong evidence that ITS-90 differs from the true thermodynamic temperature by 11 mK at 505 K . Thus, temperature metrology offers room for improvement, and an additional, independent technique would be valuable.

We have recently put into operation at NIST a new apparatus for the determination of fluid density. This densimeter operates on the familiar Archimedes (buoyancy)
principle and provides an absolute determination of density. Several twists on the conventional Archimedes experimentincluding a differential weighing with two sinkers and a magnetic suspension coupling to separate the balance from the fluid being measured-reduce the uncertainties and allow operation over a wide range of temperature and pressure. Repeatabilities at the level of a few parts per million (ppm) are routinely obtained for the densities of liquids and dense gases. Such a densimeter could create a new type of gas thermometer, while avoiding many of the headaches of the traditional constant-volume gas thermometer (CVGT).

The pressure of a low-density gas is given by a simple virial expansion:

$$
\begin{equation*}
p=\frac{R T \rho\left(1+B \rho+C \rho^{2}+D \rho^{3}\right)}{M}, \tag{1.1}
\end{equation*}
$$

where $T$ is temperature, $R$ is the molar gas constant, $\rho$ is the mass density, $M$ is the molar mass, and $B, C$ and $D$ are the second, third and fourth virial coefficients on a mass basis.

With a constant-volume gas thermometer, the temperature is derived from the pressure ratio of a fixed quantity of gas measured at some unknown temperature $T_{x}$ and at a reference temperature, usually $T_{\text {ref }}=273.16 \mathrm{~K}$. With the densimeter, the pressure would be held constant for measurements at the two temperatures; the density ratio is then extrapolated to zero pressure to yield the temperature:

$$
\begin{equation*}
\frac{T_{x}}{T_{\mathrm{ref}}}=\lim (p \rightarrow 0) \frac{\rho\left(T_{\mathrm{ref}}\right)}{\rho\left(T_{x}\right)} . \tag{1.2}
\end{equation*}
$$

The densimeter operating as a 'density-ratio gas thermometer' (DRGT) offers several advantages over a CVGT. The 'reference volume' is that of the sinkers, rather than a thinwalled gas bulb and its associated plumbing; the thermal expansion corrections would be simpler. Pressure would be generated and held constant with a piston gauge. Leaks and outgassing are potential problems with CVGT, but a small leak would be of little consequence with the DRGT, and a small flow of gas may be desirable to maintain gas purity. CVGT corrects for non-ideal behaviour by either extrapolating multiple experiments at different gas densities to zero pressure or by applying corrections based on accurate knowledge of the virial coefficients. With the density approach, measurements are made at multiple pressures and extrapolated to zero pressure. This extrapolation circumvents the need for the virials (or alternatively, provides a simultaneous fit of the virial coefficients for the particular gas being used), with the benefit of reduced sensitivity to gas purity. Note that the gas constant and molar mass drop out of (1.2). While the gas constant is well known ( $u=1.7 \mathrm{ppm}$ with $k=1$, Mohr and Taylor 2005), the molar mass is dependent on sample purity and isotopic variations. With the density method the working gas need not be of unusually high purity; it need only be the same for all the tests. CVGT usually uses helium, but argon can also be used. With the DRGT, measurements with different gases would provide an additional check on the results.

The work presented here is exploratory in nature and is intended to investigate the feasibility of this approach. It combines analytical and experimental investigations. Numerical studies are used to determine the optimum working gas and the effects of measurement uncertainties. Experiments with our present densimeter over the range of $234-505 \mathrm{~K}$


Figure 1. Simulated density ratios for $T_{x}=505.078 \mathrm{~K}$ and $T_{\text {ref }}=$ 273.16 K for several fluids (+) helium; (O) neon; (ロ) nitrogen; $(\triangle)$ argon; $(\times)$ krypton and $(\diamond) \mathrm{CF}_{4}$.
with argon, neon and nitrogen as the working gases have been carried out. These demonstrate experimental protocols, examine data analysis techniques and explore systematic errors. While these experiments have demonstrated the feasibility of the method, the present densimeter has uncertainties which are too high for temperature metrologyin particular, the uncertainty in the sinker volumes is large compared to the uncertainty in temperature. The method was, thus, inverted and used with the measured data to determine, in situ, the sinker volumes as a function of temperature.

## 2. Numerical study

### 2.1. Model

The density-ratio gas thermometer involves density measurements of a working gas at two temperatures and multiple pressures. The ratios of the densities at each pressure are computed and these ratios are extrapolated to zero pressure (or density) to give the temperature ratio according to (1.2). In principle, any gas could be measured over any range of pressures. But what is the optimum working gas, optimum pressure range and best extrapolation model? The basic assumption behind the method is that any real gas will approach ideal-gas behaviour in the limit of zero pressure, but measurements must be carried out at finite pressures, so nearly ideal behaviour at higher pressures is desirable. This argues for a very simple fluid with a low critical temperature, and helium is the most obvious example. But since mass densities are measured, large relative uncertainties in the densities would result from light gases and/or measurements at low pressures. Thus one should consider operation at higher pressures and with gases which are heavier, but which still retain a simple, symmetrical structure.

Figure 1 shows simulated results for several gases measured at $T_{x}=505.078 \mathrm{~K}$ and $T_{\text {ref }}=273.16 \mathrm{~K}$ over the pressure range $0.5-6 \mathrm{MPa}$. In addition to helium, the noble gases neon, argon and krypton are considered as well as nitrogen and the quasi-spherically symmetric $\mathrm{CF}_{4}$. The densities were computed using the equations of state in the NIST REFPROP database (Lemmon et al 2002). (REFPROP


Figure 2. Simulated density ratios over the pressure range 0.56 MPa for $T_{x}=505.078 \mathrm{~K}$ and $T_{\text {ref }}=273.16 \mathrm{~K}$ for several fluids, plotted as a function of the density at $T_{\text {ref }}$ : $(+$ ) helium; ( O ) neon; ( $\square$ ) nitrogen; $(\triangle)$ argon; $(\times)$ krypton and $(\diamond) \mathrm{CF}_{4}$.
implements accurate equations of state from the literature for industrially important fluids and fluid mixtures. It provides a convenient means of calculating properties at any desired ( $T, p$ ) state point.) All of the fluids are seen to extrapolate to the same zero-pressure limit of 0.540827 ; this, of course, is the ideal-gas limit. Helium $\left(M=4.0026 \mathrm{~g} \mathrm{~mol}^{-1}\right)$ and neon ( $M=20.179 \mathrm{~g} \mathrm{~mol}^{-1}$ ) show the most ideal behaviour, as indicated by density ratios which are nearly constant over a range of pressures. The variation of the density ratio with pressure increases with molar mass for nitrogen ( $M=$ $28.013 \mathrm{~g} \mathrm{~mol}^{-1}$ ), argon $\left(M=39.948 \mathrm{~g} \mathrm{~mol}^{-1}\right)$, $\operatorname{krypton}(M=$ $\left.83.804 \mathrm{~g} \mathrm{~mol}^{-1}\right)$ and $\mathrm{CF}_{4}\left(M=88.01 \mathrm{~g} \mathrm{~mol}^{-1}\right)$. Plotting these same data versus the density at the reference temperature (figure 2) shows the same trend with molar mass but emphasizes the large difference in densities for the different gases.

Figures 1 and 2 suggest that the density ratios could be extrapolated to zero using a simple function, such as a polynomial in pressure or density:

$$
\begin{gather*}
\frac{\rho_{x}}{\rho_{\mathrm{ref}}}=a_{0}+a_{1} p+a_{2} p^{2}  \tag{2.1}\\
\frac{\rho_{x}}{\rho_{\mathrm{ref}}}=b_{0}+b_{1} \rho_{\mathrm{ref}}+b_{2} \rho_{\mathrm{ref}}^{2} \tag{2.2}
\end{gather*}
$$

But a simple polynomial discards the knowledge of how gases behave at low pressures. A more physically based model is obtained by recasting the virial equation (1.1) in terms of densities and density ratios to provide the basic working equation for the DRGT:

$$
\begin{equation*}
\frac{\rho\left(T_{x}, p\right)}{\rho\left(T_{\mathrm{ref}}, p\right)}=\frac{f\left(T_{x}, p, B_{x}, C_{x}\right)}{f\left(T_{\mathrm{ref}}, p, B_{\mathrm{ref}}, C_{\mathrm{ref}}, D_{\mathrm{ref}}\right)} \tag{2.3}
\end{equation*}
$$

where the function in the numerator of the right-hand side is the virial equation recast in terms of a cubic equation and solved explicitly for density. The denominator of (2.3) is the virial equation at the reference temperature; by including the fourth virial coefficient $D_{\text {ref }}$, an iterative solution is required to give the density as a function of the temperature and pressure.

Equations (2.1)-(2.3) were studied for the six gases shown in figure 1 at temperatures of $83.8058 \mathrm{~K}, 234.3156 \mathrm{~K}$,
$302.9146 \mathrm{~K}, 429.7458 \mathrm{~K}, 505.078 \mathrm{~K}$ and 692.677 K , with all the 'unknown' temperatures referenced to 273.16 K. These temperatures are fixed points on ITS-90. Multiple cases of temperature, fluid and maximum pressure were considered. For each case, ten 'data points' over a range of pressures were simulated up to a maximum pressure ranging from 0.5 MPa to 20 MPa . Eight pressure ranges were simulated for each maximum pressure, with $p_{\max } / p_{\min }=2-20$.

The 'data' (the density ratios and pressures or densities) were input to an orthogonal distance regression using the package ODRPACK (Boggs et al 1992). (Orthogonal distance regression allows for uncertainties in both the independent and dependent variables; in contrast, ordinary least squares assumes that all uncertainties are in the dependent variable.) The density ratio at zero pressure or density is given by the fitted parameters $a_{0}$ and $b_{0}$ for (2.1) and (2.2), respectively. For (2.3), the unknown temperature $T_{x}$ is one of the fitted parameters, along with the virial coefficients $B_{x}, C_{x}, B_{\text {ref }}, C_{\text {ref }}$ and $D_{\text {ref. }}$. In some cases, it will be more convenient to discuss the results in terms of $T_{x}$ and in others the density ratio will be more convenient. The two are equivalent and can be converted by (1.2).

In principle, all the parameters in (2.3) could be fitted simultaneously. But this approach resulted in virial parameters which lost all connection with reality-some were even of the wrong sign. In other words, (2.3) was essentially equivalent to a completely empirical model with a large number of fitted parameters. To avoid this, the $p-\rho-T$ data at $T_{\text {ref }}$ were first fitted separately to generate $B_{\mathrm{ref}}, C_{\mathrm{ref}}$ and $D_{\mathrm{ref}}$; these were then input as fixed parameters to the fit of the density-ratio data where $T_{x}, B_{x}$ and $C_{x}$ were determined. The fourth virial coefficient at $T_{x}$ was not included in (2.3) to avoid overfitting.

For densities computed directly from the equations of state, the optimum test conditions for all the fluids, models and temperatures were the smallest maximum pressure ( $p_{\max }=$ 0.5 MPa ) and largest pressure range ( $p_{\max } / p_{\min }=20$ ). At these conditions, the fitted $T_{x}$ was within 0.01 mK of the expected value with an uncertainty of less than 0.01 mK . This result is expected; the gases are most nearly ideal at the lowest pressures, and even the simple polynomial models correctly extrapolate to zero at these conditions.

### 2.2. Effects of experimental uncertainties and optimal experimental parameters

The more interesting results come from simulating experimental uncertainties. The simulated data were varied by an estimated standard deviation multiplied by a normal distribution (evaluated at a random number). Standard deviations of 2 mK in temperature, 5 Pa in pressure and $0.0002 \mathrm{~kg} \mathrm{~m}^{-3}$ in density were assumed. In addition, a systematic offset of 20 ppm in pressure was also included. These are typical uncertainties for the present densimeter. Four 'replicates' for each data point were simulated with the random number generator.

With these 'uncertainties' the optimum fluid (as indicated by the lowest uncertainty in the fitted $T_{x}$ ) was neon at 83 K , argon at 234 K (with neon a close second), krypton at 303 K , 430 K and 505 K (with argon a close second), and argon at 693 K (with krypton a close second). The optimum maximum

Table 1. Simulation results for different fluids in a density-ratio gas thermometer; listed are the standard uncertainties in the fitted value of $T_{x}$ (in mK ) at the optimum maximum pressure ( MPa ) for the experiment, and the fitting equation giving the lowest uncertainty.

| Fluid | $T_{x}$ |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 83 K | 234 K | 303 K | 430 K | 505 K | 693 K |
| (a) Results for uncertainties in the input data of $T_{\sigma}=2 \mathrm{mK}, p_{\sigma}=5 \mathrm{~Pa}, \rho_{\sigma}=0.0002 \mathrm{~kg} \mathrm{~m}^{-3}$ and a systematic offset in pressure of 20 ppm |  |  |  |  |  |  |
| Helium | 2.42 @ 12 (2.1) | 5.45@ 20 (2.2) | 7.60 @ 20 (2.2) | 12.9 @ 20 (2.2) | 16.8@ 20 (2.2) | 29.2@ 20 (2.2) |
| Neon | 0.80@ 6 (2.3) | 1.35@12(2.2) | 1.50 @ 20 (2.1) | 3.01 @ 20 (2.1) | 3.77 @ 20 (2.1) | 5.71 @ 20 (2.3) |
| Nitrogen | n.a. | 2.06 @ 6 (2.3) | 2.52 @ 8 (2.3) | 2.97 @ 12 (2.3) | 4.81 @ 12 (2.3) | 6.60 @ 12 (2.1) |
| Argon | n.a. | 1.32 @ 6 (2.3) | 1.77 @ 8 (2.3) | 2.17 @ 12 (2.3) | 2.59 @ 12 (2.3) | 3.81@16(2.2) |
| Krypton | n.a. | n.a. | 1.11 @ 6 (2.3) | 1.86@ 6 (2.3) | 2.47 @ 6(2.3) | 4.04 @ 8 (2.2) |
| $\mathrm{CF}_{4}$ | n.a. | n.a. | 3.09 @ 2 (2.2) | 3.32 @ 4 (2.1) | 4.12@ 4 (2.1) | 6.84@ 4 (2.1) |
| (b) Results for uncertainties in the input data of $T_{\sigma}=0.5 \mathrm{mK}, p_{\sigma}=1 \mathrm{~Pa}, \rho_{\sigma}=0.00005 \mathrm{~kg} \mathrm{~m}^{-3}$ and a systematic offset in pressure of 10 ppm |  |  |  |  |  |  |
| Helium | 1.02@ 6 (2.1) | 1.33 @ 20 (2.2) | 1.82 @ 20 (2.1) | 3.28 @ 20 (2.1) | 4.34 @ 20 (2.1) | 7.30 @ 20 (2.1) |
| Neon | 0.34@4(2.3) | 0.39@12(2.2) | 0.42 @ 20 (2.1) | 1.18 @ 20 (2.1) | 1.40 @ 12 (2.3) | 2.02@ 16 (2.3) |
| Nitrogen | n.a. | 0.62 @ 6 (2.2) | 0.89 @ 6 (2.3) | 0.99@ 10 (2.3) | 1.51 @ 8 (2.3) | 1.85@10(2.1) |
| Argon | n.a. | 0.34@6(2.2) | 0.63 @ 6 (2.2) | 0.93 @ 8 (2.3) | 0.92@ 10 (2.3) | 1.36@10(2.3) |
| Krypton | n.a. | n.a. | 0.35 @ 4 (2.3) | 0.56 @ 6 (2.3) | 0.66@6(2.3) | 1.12@6(2.3) |
| $\mathrm{CF}_{4}$ | n.a. | n.a. | 1.19@2(2.3) | 1.29@2(2.3) | 1.72@2(2.3) | 5.76@1(2.1) |

n.a.: not applicable, $T_{x}$ is at or below the triple point temperature of the fluid or near the critical temperature.
pressure was $6-16 \mathrm{MPa}$. The best pressure ratio was a broad optimum with $p_{\max } / p_{\text {min }}$ between 6 and 12 . In all cases, helium showed the largest uncertainties; however, the best pressure for helium may lie above the maximum of 20 MPa simulated here. $\mathrm{CF}_{4}$ was the second-worst fluid in all cases. The best model was (2.3) at five of the six temperatures. The polynomial in density was the best at 693 K , but only by a small margin. Under the optimum conditions, the standard uncertainty ranged from 0.80 mK for neon at 84 K to 3.81 mK for argon at 693 K . This standard uncertainty is the standard deviation in the $T_{x}$ parameter returned by the ODRPACK fitting routine. These results are summarized in table $1(a)$.

A second set of cases with standard deviations of 0.5 mK in temperature, 1 Pa in pressure and $0.00005 \mathrm{~kg} \mathrm{~m}^{-3}$ in density, along with a systematic offset of 10 ppm in pressure, was also considered. These were intended to simulate an improved densimeter. The optimums were very similar. Krypton was now the best fluid, and by a wider margin, at $303 \mathrm{~K}, 430 \mathrm{~K}$, 505 K and 693 K . Neon was the second best at 303 K , but, as before, argon was a good choice at all temperatures except for 84 K . The optimum $p_{\text {max }}$ was lower at $4-6 \mathrm{MPa}$. The optimum pressure represents a balance between the ability of the numerical model to represent the data and the minimization of experimental errors. The virial model is nearly exact at low pressures, but here the relative experimental uncertainties are large. With smaller experimental uncertainties the optimum moves to lower pressures. The best model was again (2.3) in most cases. The standard uncertainty ranged from 0.34 mK for neon at 84 K to 1.12 mK for krypton at 693 K . These results are summarized in table $1(b)$.

## 3. Experimental details

An experimental exploration of the density-ratio gas thermometer was carried out using a two-sinker densimeter. Measurements were carried out over the temperature range 234.316-505.078 K using argon, neon or nitrogen as the
working gas. Argon was used as the primary gas in these tests because it was identified as the best or second best fluid over this temperature range by the numerical studies described above and it is readily available in high purity. Krypton was the optimum gas at temperatures between 303 K and 505 K , but it is much more expensive than argon. Nitrogen was ranked lower, but previously measured data (McLinden and LöschWill 2006) were available which could be analysed by the present method. Limited tests using neon provided a check on the method using an additional gas.

### 3.1. Apparatus description

The key elements of the densimeter are two sinkers which are immersed in the fluid of interest inside a measuring cell (pressure vessel). The sinkers are weighed with a highprecision balance which is separated from the fluid by a magnetic suspension coupling consisting of an electromagnet (in air) hung from the balance and a permanent magnet and 'lifting fork' assembly (in the fluid) which pick up the sinkers for weighing. The sinkers are made of titanium and tantalum and have the same mass ( 60 g ) and surface area but very different volumes. The use of two sinkers greatly reduces the effects of any systematic errors in the weighings. The loading on the balance is nearly the same for the two weighings, reducing errors from any nonlinearity in the balance. More importantly, the loading on the magnetic suspension coupling is nearly the same, largely cancelling the 'force transmission error' associated with the coupling. Two-sinker densimeters are further described by Wagner and Kleinrahm (2004), and force transmission errors are discussed in this reference as well as by McLinden et al (2006).

A thermostat consisting of multiple layers of passive and actively controlled shields provides a uniform temperature. The temperature of the measuring cell was measured with a standard platinum resistance thermometer read by a resistance bridge. For this work, pressures were measured with a gasoperated piston gauge.

For the two-sinker technique, the density is given by

$$
\begin{equation*}
\rho=\frac{\left(m_{1}-W_{1}\right)-\left(m_{2}-W_{2}\right)}{\left(V_{1}-V_{2}\right)}, \tag{3.1}
\end{equation*}
$$

where $m$ and $V$ are the mass and volume of a sinker and $W$ is the balance reading; the subscripts refer to the two sinkers. The sinker volumes were determined at 293.15 K by a hydrostatic comparator technique relative to silicon density standards; the technique and procedures detailed by Bowman et al (1974) were followed closely in this determination. The sinker volumes as functions of temperature were calculated from measured values of the thermal expansion coefficients. The ( $k=1$ ) uncertainty in the sinker volumes varies from 12 ppm at 293 K to 125 ppm at 505 K . A complete description of the apparatus and a detailed uncertainty analysis are provided by McLinden and Lösch-Will (2006).

### 3.2. Test procedure

Tests were run along isotherms starting at the highest pressure. Four to six replicate density determinations were made at each $(T, p)$ state point. The sample was then vented to the next lowest pressure, and the process was repeated for six pressures per isotherm. After each isotherm, the cell was evacuated and the sinkers were again weighed; just as the zero of any instrument will drift, the calculated density in vacuum differed slightly from zero (average value of $0.005 \mathrm{~kg} \mathrm{~m}^{-3}$ ), and this 'apparatus zero' was subtracted from the measured densities.

The pressures were both measured and regulated with the piston gauge. (That is, the measuring cell and piston gauge were directly connected; there was no pressure-separating diaphragm.) For each pressure, the gas sample was vented to a pressure $1-2 \%$ higher than desired (with a valve between the measuring cell and the piston gauge closed) and allowed to come to within 0.01 K of temperature equilibrium. The piston gauge was floated and the valve connecting the piston gauge and measuring cell was opened. This caused the piston to rise, and the gas volume was adjusted using a manual pressure controller to return the piston to the reference position. (The piston gauge used the same working gas as was being measured. Nevertheless, we were careful to allow gas flow only from the cell to the piston gauge while adjusting pressures to avoid contamination of the sample.) A further $30-60 \mathrm{~min}$ of equilibration time was allowed with the piston gauge connected and floating before commencing the density determinations. This resulted in temperatures which were steady to within 1 mK . The height and the rotation speed of the piston were adjusted only between density determinations. This arrangement provided exceptionally steady pressures during a test and also made it easy to replicate pressures from one temperature to the next. Very small day-to-day variations in pressure ( on the order of 10 ppm ) arose from changes in room temperature.

### 3.3. Experimental samples

The argon was 'ULSI-grade' (Matheson Tri-Gas, Inc.) ${ }^{1}$ with a certified purity of $99.9999 \%$. A cylinder analysis by the
1 Certain trade names and products are given to adequately document the experimental equipment and procedures. This does not constitute a recommendation or endorsement of these products by the National Institute of Standards and Technology, nor does it imply that the products are necessarily the best available for the purpose.
supplier indicated impurities of nitrogen at 0.49 ppm , oxygen at 0.11 ppm and water at 0.17 ppm . Carbon dioxide, carbon monoxide and methane were not detected at the level of 0.1 ppm . The nitrogen was 'VLSI-grade' (Scott Specialty Gases). The specification for this gas is $99.9995 \%$ purity, and a cylinder analysis by the supplier indicated concentrations of oxygen, carbon dioxide, carbon monoxide, water and total hydrocarbons of $0.2-0.5 \mathrm{ppm}$. The concentration of hydrogen was 1 ppm for an overall purity of $99.9997 \%$. The supplier's batch analysis for the research grade neon (Air Products and Chemicals, Inc.) indicated helium, nitrogen, oxygen and carbon monoxide at levels of $1-5 \mathrm{ppm}$; water, total hydrocarbons and carbon dioxide were at levels of $0.2-$ 0.5 ppm for an overall purity of $99.999 \%$. Our own analysis by gas chromatography was consistent with the suppliers analysis. We also checked for the presence of argon in the nitrogen and neon samples and found none at the 1 ppm level.

### 3.4. Results

Measurements were made with argon at $234 \mathrm{~K}, 273 \mathrm{~K}$, two series at $293 \mathrm{~K}, 360 \mathrm{~K}$, two series at 430 K and 505 K ; nitrogen was measured at $293 \mathrm{~K}, 340 \mathrm{~K}, 400 \mathrm{~K}, 440 \mathrm{~K}$ and 480 K and neon was tested at 293 K and 430 K . Several of these temperatures will be recognized as fixed points on ITS-90. Measuring at these temperatures minimizes the temperature uncertainty since the PRT was calibrated at these points. The sinker volumes were determined at 293.15 K , and this will serve at the reference temperature for the present experimental results.

The measured $p-\rho-T$ data for argon and neon are given in tables 2 and 3. Densities are given based both on sinker volumes determined from the thermal expansion data (designated as $\rho_{\text {CTE }}$ in the tables) and also using the 'corrected' volumes described in section 3.7. The nitrogen measurements have been previously reported (McLinden and Lösch-Will 2006); that reference reports only the 'corrected' densities. The actual sequence of the argon and neon tests is noted in tables 2 and 3. The nitrogen tests at 293 K were carried out in June 2003; the other temperatures were measured in October 2003 in the sequence $480 \mathrm{~K}, 400 \mathrm{~K}, 340 \mathrm{~K}$ and 440 K . The uncertainties $(k=1)$ for the data are 2 mK in the temperatures and $\pm(0.0011 \%+2 \mathrm{~Pa})$ in pressure. The uncertainties in the 'corrected' densities are given by

$$
\begin{align*}
u= & \left\{[12+0.28|T-293|]^{2}+[0.62 p]^{2}\right\}^{0.5} \times \mathrm{ppm} \\
& +\left(0.0003+0.9 \times 10^{-6}|T-293|\right) \times \mathrm{kg} \mathrm{~m}^{-3}, \tag{3.2}
\end{align*}
$$

where $T$ is the temperature in K and $p$ is the pressure in MPa. The terms inside the braces arise from the uncertainties in the sinker volumes at $T_{\text {ref }}$ and also the uncertainties in the sinker volumes as functions of temperature and pressure; these result in a constant relative error in density at a given $T$ and $p$. The final terms arise from uncertainties in the weighings, which result in an absolute error in the density.

The argon data have an average absolute deviation of 34 ppm compared to the equation of state of Tegeler et al (1999). This EOS was fitted to extensive high-accuracy data for argon and has an uncertainty in gas-phase densities of 200 ppm for temperatures up to 340 K and 300 ppm for temperatures up to 520 K . Thus, the present data are consistent

Table 2. Experimental $p-\rho-T$ data for argon with a comparison to the equation of state of Tegeler et al (1999). Densities are computed using sinker volumes based on the coefficient of thermal expansion ( $\rho_{\text {СТЕ }}$ ) and using the 'corrected' sinker volumes discussed in section 3.7
( $\rho_{\text {corrected }}$ ).

| $T(\mathrm{~K})$ | $p$ (MPa) | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\text {EOS }}\right) / \rho_{\text {EOS }}(\mathrm{ppm})$ |
| :---: | :---: | :---: | :---: | :---: |
| $T=234.32 \mathrm{~K} ; 10-11$ February 2005 |  |  |  |  |
| 234.3162 | 5.994459 | 136.0194 | 136.0254 | 16.3 |
| 234.3162 | 5.994454 | 136.0196 | 136.0257 | 18.6 |
| 234.3188 | 5.994450 | 136.0178 | 136.0239 | 22.2 |
| 234.3183 | 5.994446 | 136.0173 | 136.0233 | 15.8 |
| 234.3163 | 5.994442 | 136.0188 | 136.0249 | 16.0 |
| 234.3170 | 3.996244 | 87.7084 | 87.7123 | 42.1 |
| 234.3173 | 3.996242 | 87.7075 | 87.7114 | 34.1 |
| 234.3185 | 3.996241 | 87.7067 | 87.7106 | 31.8 |
| 234.3181 | 3.996240 | 87.7065 | 87.7104 | 27.4 |
| 234.3168 | 3.996238 | 87.7067 | 87.7106 | 23.7 |
| 234.3166 | 2.997160 | 64.6734 | 64.6763 | 49.9 |
| 234.3170 | 2.997159 | 64.6725 | 64.6753 | 38.1 |
| 234.3188 | 2.997159 | 64.6718 | 64.6747 | 37.8 |
| 234.3178 | 2.997158 | 64.6717 | 64.6746 | 31.3 |
| 234.3167 | 2.997157 | 64.6720 | 64.6749 | 29.9 |
| 234.3182 | 1.998081 | 42.3852 | 42.3871 | 38.5 |
| 234.3173 | 1.998082 | 42.3852 | 42.3871 | 34.2 |
| 234.3186 | 1.998082 | 42.3851 | 42.3870 | 38.8 |
| 234.3187 | 1.998083 | 42.3851 | 42.3870 | 38.0 |
| 234.3176 | 1.998083 | 42.3852 | 42.3871 | 35.2 |
| 234.3165 | 0.999045 | 20.8360 | 20.8369 | 87.4 |
| 234.3164 | 0.999046 | 20.8356 | 20.8365 | 64.3 |
| 234.3182 | 0.999046 | 20.8353 | 20.8363 | 62.9 |
| 234.3184 | 0.999046 | 20.8351 | 20.8361 | 54.0 |
| 234.3167 | 0.999046 | 20.8352 | 20.8361 | 48.7 |
| 234.3189 | 0.499520 | 10.3296 | 10.3300 | 60.9 |
| 234.3175 | 0.499520 | 10.3297 | 10.3302 | 69.7 |
| 234.3168 | 0.499520 | 10.3295 | 10.3300 | 46.2 |
| 234.3182 | 0.499520 | 10.3294 | 10.3299 | 47.1 |
| 234.3182 | 0.499520 | 10.3295 | 10.3300 | 57.0 |
| $T=273.16$ K; 31 January-2 February 2005 |  |  |  |  |
| 273.1597 | 6.993345 | 130.4092 | 130.4122 | 6.8 |
| 273.1587 | 6.993341 | 130.4096 | 130.4125 | 5.4 |
| 273.1606 | 6.993340 | 130.4087 | 130.4117 | 7.9 |
| 273.1615 | 6.993340 | 130.4081 | 130.4111 | 7.7 |
| 273.1591 | 6.993333 | 130.4089 | 130.4118 | 3.2 |
| 273.1598 | 5.994235 | 110.9715 | 110.9740 | 10.1 |
| 273.1605 | 5.994239 | 110.9715 | 110.9740 | 12.4 |
| 273.1611 | 5.994242 | 110.9709 | 110.9734 | 9.7 |
| 273.1598 | 5.994244 | 110.9714 | 110.9739 | 7.2 |
| 273.1592 | 5.994246 | 110.9717 | 110.9742 | 7.4 |
| 273.1602 | 4.995205 | 91.7693 | 91.7714 | 29.5 |
| 273.1587 | 4.995204 | 91.7691 | 91.7712 | 21.5 |
| 273.1597 | 4.995203 | 91.7686 | 91.7707 | 20.7 |
| 273.1619 | 4.995203 | 91.7677 | 91.7698 | 20.7 |
| 273.1602 | 4.995203 | 91.7679 | 91.7700 | 14.8 |
| 273.1585 | 3.996145 | 72.8255 | 72.8272 | 52.7 |
| 273.1611 | 3.996144 | 72.8239 | 72.8255 | 41.3 |
| 273.1603 | 3.996145 | 72.8232 | 72.8248 | 27.9 |
| 273.1587 | 3.996145 | 72.8231 | 72.8247 | 19.9 |
| 273.1602 | 3.996144 | 72.8230 | 72.8246 | 25.1 |
| 273.1596 | 2.997088 | 54.1577 | 54.1590 | 24.2 |
| 273.1596 | 2.997089 | 54.1578 | 54.1590 | 24.9 |
| 273.1612 | 2.997090 | 54.1574 | 54.1587 | 24.2 |
| 273.1610 | 2.997092 | 54.1575 | 54.1588 | 25.0 |
| 273.1596 | 2.997094 | 54.1577 | 54.1590 | 22.5 |

Table 2. (Continued.)

| $T$ (K) | $p$ (MPa) | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\text {EOS }}\right) / \rho_{\text {EOS }}(\mathrm{ppm})$ |
| :---: | :---: | :---: | :---: | :---: |
| 273.1615 | 1.998056 | 35.7920 | 35.7928 | 68.5 |
| 273.1599 | 1.998055 | 35.7916 | 35.7924 | 52.0 |
| 273.1611 | 1.998055 | 35.7909 | 35.7918 | 38.5 |
| 273.1620 | 1.998055 | 35.7908 | 35.7916 | 37.5 |
| 273.1608 | 1.998056 | 35.7907 | 35.7915 | 31.3 |
| 273.1602 | 0.999032 | 17.7354 | 17.7358 | 51.6 |
| 273.1596 | 0.999032 | 17.7353 | 17.7357 | 46.4 |
| 273.1619 | 0.999032 | 17.7348 | 17.7352 | 26.1 |
| 273.1624 | 0.999032 | 17.7349 | 17.7354 | 36.8 |
| 273.1604 | 0.999032 | 17.7352 | 17.7356 | 40.8 |
| 273.1615 | 0.499515 | 8.8271 | 8.8273 | 58.0 |
| 273.1602 | 0.499515 | 8.8271 | 8.8273 | 59.1 |
| 273.1608 | 0.499515 | 8.8270 | 8.8272 | 50.0 |
| 273.1617 | 0.499515 | 8.8271 | 8.8273 | 59.0 |
| 273.1617 | 0.499515 | 8.8269 | 8.8271 | 36.1 |
| $T=293.15 \mathrm{~K}$ (series 1); 19-20 October 2004 |  |  |  |  |
| 293.1521 | 5.994107 | 101.8247 | 101.8247 | 27.7 |
| 293.1549 | 5.994101 | 101.8227 | 101.8227 | 20.9 |
| 293.1522 | 5.994096 | 101.8227 | 101.8227 | 10.6 |
| 293.1539 | 5.994091 | 101.8220 | 101.8220 | 11.8 |
| 293.1531 | 5.994085 | 101.8218 | 101.8218 | 7.3 |
| 293.1494 | 3.996026 | 67.1623 | 67.1623 | 28.6 |
| 293.1520 | 3.996024 | 67.1615 | 67.1615 | 27.4 |
| 293.1526 | 3.996025 | 67.1611 | 67.1611 | 23.9 |
| 293.1501 | 3.996022 | 67.1613 | 67.1613 | 18.0 |
| 293.1517 | 3.996021 | 67.1608 | 67.1608 | 16.8 |
| 293.1517 | 2.997004 | 50.0775 | 50.0775 | 38.3 |
| 293.1500 | 2.997004 | 50.0772 | 50.0772 | 26.3 |
| 293.1518 | 2.997003 | 50.0769 | 50.0769 | 26.8 |
| 293.1538 | 2.997003 | 50.0764 | 50.0764 | 24.0 |
| 293.1517 | 2.997003 | 50.0765 | 50.0765 | 18.1 |
| 293.1516 | 1.997958 | 33.1788 | 33.1788 | 23.1 |
| 293.1532 | 1.997957 | 33.1787 | 33.1787 | 27.9 |
| 293.1541 | 1.997958 | 33.1784 | 33.1784 | 21.1 |
| 293.1525 | 1.997958 | 33.1786 | 33.1786 | 22.3 |
| 293.1517 | 1.997958 | 33.1787 | 33.1787 | 20.1 |
| 293.1511 | 0.998986 | 16.4833 | 16.4833 | 34.5 |
| 293.1523 | 0.998986 | 16.4833 | 16.4833 | 36.3 |
| 293.1553 | 0.998985 | 16.4830 | 16.4830 | 34.4 |
| 293.1535 | 0.998985 | 16.4831 | 16.4831 | 28.5 |
| 293.1521 | 0.998985 | 16.4832 | 16.4832 | 35.3 |
| 293.1515 | 0.499494 | 8.2146 | 8.2146 | 58.7 |
| 293.1536 | 0.499494 | 8.2146 | 8.2146 | 61.1 |
| 293.1539 | 0.499494 | 8.2143 | 8.2143 | 35.5 |
| 293.1519 | 0.499494 | 8.2143 | 8.2143 | 23.4 |
| 293.1535 | 0.499494 | 8.2143 | 8.2143 | 28.9 |
| $T=293.15 \mathrm{~K}$ (series 2); 28-29 January 2005 |  |  |  |  |
| 293.1527 | 5.994233 | 101.8259 | 101.8259 | 21.1 |
| 293.1547 | 5.994230 | 101.8253 | 101.8253 | 22.9 |
| 293.1546 | 5.994226 | 101.8247 | 101.8247 | 17.5 |
| 293.1537 | 5.994221 | 101.8247 | 101.8247 | 14.6 |
| 293.1536 | 5.994216 | 101.8245 | 101.8245 | 13.7 |
| 293.1538 | 3.996110 | 67.1632 | 67.1632 | 38.3 |
| 293.1533 | 3.996106 | 67.1627 | 67.1627 | 29.9 |
| 293.1552 | 3.996103 | 67.1620 | 67.1620 | 27.3 |
| 293.1552 | 3.996100 | 67.1617 | 67.1617 | 23.8 |
| 293.1540 | 3.996098 | 67.1618 | 67.1618 | 20.4 |
| 293.1538 | 2.997057 | 50.0776 | 50.0776 | 30.3 |
| 293.1558 | 2.997057 | 50.0772 | 50.0772 | 29.7 |
| 293.1546 | 2.997057 | 50.0771 | 50.0771 | 24.1 |
| 293.1530 | 2.997057 | 50.0773 | 50.0773 | 20.4 |
| 293.1541 | 2.997058 | 50.0770 | 50.0770 | 19.3 |

Table 2. (Continued.)

| $T$ (K) | $p$ (MPa) | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\text {EOS }}\right) / \rho_{\text {EOS }}(\mathrm{ppm})$ |
| :---: | :---: | :---: | :---: | :---: |
| 293.1536 | 1.998017 | 33.1796 | 33.1796 | 25.8 |
| 293.1522 | 1.998018 | 33.1794 | 33.1794 | 14.1 |
| 293.1530 | 1.998019 | 33.1794 | 33.1794 | 16.4 |
| 293.1542 | 1.998020 | 33.1794 | 33.1794 | 18.8 |
| 293.1530 | 1.998021 | 33.1795 | 33.1795 | 18.6 |
| 293.1530 | 0.999012 | 16.4835 | 16.4835 | 26.6 |
| 293.1542 | 0.999012 | 16.4834 | 16.4834 | 27.3 |
| 293.1528 | 0.999013 | 16.4834 | 16.4834 | 21.5 |
| 293.1519 | 0.999013 | 16.4836 | 16.4836 | 27.5 |
| 293.1534 | 0.999013 | 16.4835 | 16.4835 | 25.9 |
| 293.1513 | 0.499507 | 8.2148 | 8.2148 | 59.3 |
| 293.1539 | 0.499507 | 8.2148 | 8.2148 | 68.1 |
| 293.1542 | 0.499507 | 8.2146 | 8.2146 | 43.9 |
| 293.1521 | 0.499507 | 8.2147 | 8.2147 | 42.7 |
| 293.1526 | 0.499508 | 8.2146 | 8.2146 | 31.6 |
| $T=360.00 \mathrm{~K} ; 15-17$ September 2004 |  |  |  |  |
| 360.0007 | 5.994233 | 80.5817 | 80.5738 | 1.8 |
| 360.0015 | 5.994247 | 80.5814 | 80.5735 | -1.5 |
| 360.0011 | 5.994239 | 80.5813 | 80.5735 | -1.9 |
| 360.0024 | 5.994234 | 80.5809 | 80.5730 | -2.9 |
| 360.0009 | 5.994230 | 80.5810 | 80.5731 | -4.8 |
| 360.0000 | 3.996117 | 53.6394 | 53.6341 | -3.2 |
| 360.0021 | 3.996115 | 53.6392 | 53.6339 | 0.2 |
| 360.0022 | 3.996113 | 53.6391 | 53.6339 | -1.0 |
| 360.0004 | 3.996112 | 53.6392 | 53.6340 | -3.2 |
| 360.0020 | 3.996111 | 53.6390 | 53.6338 | -2.1 |
| 360.0001 | 2.997064 | 40.1855 | 40.1816 | -0.2 |
| 360.0020 | 2.997063 | 40.1854 | 40.1814 | 2.3 |
| 360.0026 | 2.997063 | 40.1853 | 40.1814 | 2.6 |
| 360.0014 | 2.997062 | 40.1852 | 40.1813 | -2.7 |
| 360.0011 | 2.997062 | 40.1853 | 40.1813 | -2.1 |
| 360.0021 | 1.998047 | 26.7550 | 26.7524 | -8.5 |
| 360.0010 | 1.998047 | 26.7552 | 26.7526 | -6.3 |
| 359.9999 | 1.998047 | 26.7551 | 26.7525 | 11.1 |
| 360.0014 | 1.998047 | 26.7552 | 26.7526 | -3.5 |
| 360.0019 | 1.998047 | 26.7551 | 26.7525 | -6.3 |
| 360.0002 | 0.999026 | 13.3574 | 13.3561 | -1.6 |
| 360.0014 | 0.999026 | 13.3574 | 13.3561 | -1.4 |
| 360.0026 | 0.999025 | 13.3573 | 13.3560 | -5.2 |
| 360.0013 | 0.999024 | 13.3573 | 13.3560 | -8.0 |
| 360.0004 | 0.999024 | 13.3574 | 13.3561 | -2.4 |
| 360.0013 | 0.499511 | 6.6733 | 6.6726 | 26.4 |
| 360.0007 | 0.499510 | 6.6731 | 6.6725 | 2.4 |
| 360.0027 | 0.499510 | 6.6731 | 6.6725 | 8.6 |
| 360.0034 | 0.499510 | 6.6732 | 6.6725 | 18.5 |
| 360.0013 | 0.499510 | 6.6731 | 6.6725 | 5.8 |
| $T=429.75 \mathrm{~K}$ (series 1); 20-21 September 2004 |  |  |  |  |
| 429.7483 | 5.994172 | 66.6498 | 66.6389 | 48.3 |
| 429.7495 | 5.994170 | 66.6497 | 66.6388 | 50.2 |
| 429.7483 | 5.994169 | 66.6495 | 66.6386 | 44.7 |
| 429.7496 | 5.994168 | 66.6493 | 66.6384 | 44.2 |
| 429.7498 | 5.994166 | 66.6495 | 66.6386 | 47.8 |
| 429.7481 | 3.996090 | 44.5412 | 44.5339 | 50.0 |
| 429.7479 | 3.996100 | 44.5414 | 44.5341 | 51.2 |
| 429.7485 | 3.996085 | 44.5411 | 44.5339 | 49.8 |
| 429.7479 | 3.996083 | 44.5411 | 44.5338 | 47.6 |
| 429.7477 | 3.996081 | 44.5412 | 44.5339 | 49.4 |
| 429.7486 | 3.996079 | 44.5410 | 44.5337 | 48.8 |
| 429.7484 | 3.996076 | 44.5409 | 44.5336 | 46.1 |
| 429.7473 | 2.997039 | 33.4392 | 33.4338 | 39.7 |
| 429.7460 | 2.997040 | 33.4393 | 33.4338 | 37.0 |
| 429.7468 | 2.997041 | 33.4391 | 33.4336 | 32.4 |

Table 2. (Continued.)

| $T$ (K) | $p$ (MPa) | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\mathrm{EOS}}\right) / \rho_{\mathrm{EOS}}(\mathrm{ppm})$ |
| :---: | :---: | :---: | :---: | :---: |
| 429.7478 | 2.997042 | 33.4392 | 33.4337 | 39.3 |
| 429.7468 | 2.997043 | 33.4392 | 33.4337 | 35.1 |
| 429.7458 | 1.998041 | 22.3131 | 22.3094 | 58.5 |
| 429.7461 | 1.998039 | 22.3126 | 22.3090 | 38.1 |
| 429.7443 | 1.998039 | 22.3128 | 22.3091 | 42.7 |
| 429.7446 | 1.998037 | 22.3128 | 22.3091 | 42.8 |
| 429.7460 | 1.998036 | 22.3128 | 22.3092 | 48.9 |
| 429.7454 | 0.999020 | 11.1647 | 11.1629 | 46.7 |
| 429.7441 | 0.999020 | 11.1647 | 11.1629 | 44.2 |
| 429.7453 | 0.999019 | 11.1646 | 11.1627 | 32.2 |
| 429.7469 | 0.999019 | 11.1648 | 11.1630 | 56.9 |
| 429.7457 | 0.999019 | 11.1648 | 11.1630 | 58.4 |
| 429.7444 | 0.499510 | 5.5844 | 5.5834 | 78.7 |
| 429.7462 | 0.499510 | 5.5842 | 5.5833 | 57.1 |
| 429.7464 | 0.499510 | 5.5841 | 5.5832 | 38.9 |
| 429.7448 | 0.499510 | 5.5841 | 5.5832 | 34.7 |
| 429.7456 | 0.499510 | 5.5840 | 5.5831 | 19.2 |
| 429.7468 | 0.499510 | 5.5840 | 5.5831 | 21.4 |
| $T=429.75 \mathrm{~K}$ (series 2); 26-27 October 2004 |  |  |  |  |
| 429.7490 | 5.993982 | 66.6469 | 66.6360 | 38.5 |
| 429.7486 | 5.993978 | 66.6471 | 66.6362 | 40.4 |
| 429.7487 | 5.993973 | 66.6468 | 66.6359 | 37.1 |
| 429.7486 | 5.993968 | 66.6470 | 66.6361 | 40.0 |
| 429.7492 | 5.993965 | 66.6469 | 66.6360 | 40.9 |
| 429.7501 | 3.995944 | 44.5392 | 44.5319 | 44.8 |
| 429.7483 | 3.995939 | 44.5392 | 44.5320 | 43.2 |
| 429.7493 | 3.995936 | 44.5391 | 44.5318 | 42.4 |
| 429.7509 | 3.995933 | 44.5390 | 44.5317 | 46.0 |
| 429.7498 | 3.995929 | 44.5390 | 44.5317 | 43.2 |
| 429.7498 | 2.996930 | 33.4378 | 33.4324 | 40.2 |
| 429.7476 | 2.996929 | 33.4379 | 33.4324 | 38.0 |
| 429.7487 | 2.996928 | 33.4379 | 33.4324 | 40.6 |
| 429.7508 | 2.996926 | 33.4377 | 33.4322 | 40.1 |
| 429.7501 | 2.996925 | 33.4378 | 33.4323 | 40.0 |
| 429.7503 | 1.997995 | 22.3119 | 22.3082 | 38.6 |
| 429.7495 | 1.997991 | 22.3119 | 22.3082 | 38.2 |
| 429.7487 | 1.997987 | 22.3120 | 22.3084 | 43.2 |
| 429.7498 | 1.997983 | 22.3116 | 22.3080 | 31.1 |
| 429.7500 | 1.997980 | 22.3115 | 22.3079 | 29.2 |
| 429.7506 | 0.998987 | 11.1638 | 11.1620 | 12.5 |
| 429.7488 | 0.998985 | 11.1641 | 11.1622 | 29.3 |
| 429.7499 | 0.998983 | 11.1640 | 11.1622 | 28.0 |
| 429.7518 | 0.998983 | 11.1639 | 11.1621 | 28.7 |
| 429.7509 | 0.998981 | 11.1639 | 11.1621 | 24.0 |
| 429.7502 | 0.499488 | 5.5837 | 5.5828 | 27.7 |
| 429.7493 | 0.499487 | 5.5837 | 5.5828 | 25.3 |
| 429.7502 | 0.499487 | 5.5837 | 5.5828 | 27.7 |
| 429.7515 | 0.499487 | 5.5838 | 5.5829 | 50.0 |
| 429.7506 | 0.499487 | 5.5836 | 5.5827 | 9.8 |
| $T=505.08 \mathrm{~K} ; 5-7$ October 2004 |  |  |  |  |
| 505.0837 | 5.994050 | 56.3608 | 56.3561 | 221.7 |
| 505.0823 | 5.994046 | 56.3613 | 56.3565 | 227.7 |
| 505.0833 | 5.994042 | 56.3610 | 56.3563 | 225.8 |
| 505.0847 | 5.994039 | 56.3604 | 56.3557 | 219.0 |
| 505.0841 | 5.994037 | 56.3601 | 56.3554 | 213.2 |
| 505.0858 | 3.995991 | 37.7340 | 37.7308 | 198.2 |
| 505.0849 | 3.995990 | 37.7336 | 37.7304 | 186.0 |
| 505.0841 | 3.995988 | 37.7341 | 37.7310 | 199.6 |
| 505.0850 | 3.995987 | 37.7341 | 37.7309 | 200.8 |
| 505.0859 | 3.995987 | 37.7338 | 37.7306 | 194.1 |

Table 2. (Continued.)

| $T(\mathrm{~K})$ | $p(\mathrm{MPa})$ | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\text {EOS }}\right) / \rho_{\text {EOS }}(\mathrm{ppm})$ |
| :--- | :--- | :--- | :--- | :--- |
| 505.0834 | 2.996966 | 28.3556 | 28.3533 | 123.3 |
| 505.0830 | 2.996968 | 28.3555 | 28.3532 | 118.2 |
| 505.0836 | 2.996969 | 28.3556 | 28.3533 | 122.5 |
| 505.0847 | 2.996967 | 28.3555 | 28.3531 | 121.8 |
| 505.0836 | 2.996975 | 28.3555 | 28.3531 | 114.8 |
| 505.0821 | 1.997984 | 18.9398 | 18.9382 | 77.1 |
| 505.0842 | 1.997983 | 18.9400 | 18.9384 | 91.5 |
| 505.0848 | 1.997984 | 18.9398 | 18.9383 | 84.4 |
| 505.0839 | 1.997984 | 18.9398 | 18.9382 | 79.9 |
| 505.0819 | 1.997984 | 18.9397 | 18.9381 | 73.1 |
| 505.0835 | 0.998998 | 9.4874 | 9.4866 | 52.2 |
| 505.0824 | 0.998998 | 9.4873 | 9.4865 | 39.5 |
| 505.0828 | 0.998998 | 9.4874 | 9.4866 | 50.8 |
| 505.0842 | 0.998997 | 9.4874 | 9.4866 | 58.7 |
| 505.0842 | 0.998997 | 9.4875 | 9.4867 | 63.6 |

with the EOS and, indirectly, the literature data. The deviations for neon are larger ( $\mathrm{AAD}=162 \mathrm{ppm}$ ), but the EOS for neon is older, and it is based on fewer and less accurate data compared to argon.

Table 4 presents a detailed analysis of the argon results for the first measurement series at 293 K and 430 K as an example. The data at the reference temperature of 293 K and the 'test' temperature of 430 K are presented as two side-by-side groups. All of the measurements are averaged to obtain 293.1524 K and 429.7468 K for the reference and test temperatures. The replicates at each of the pressures are averaged over both temperatures to obtain average pressures. The experimental $p-\rho-T$ points are then adjusted to the average $(T, p)$ using an equation of state:

$$
\begin{equation*}
\rho_{\text {adjusted }}=\rho_{\exp } \frac{\rho_{\mathrm{EOS}}\left(T_{\mathrm{avg}}, p_{\mathrm{avg}}\right)}{\rho_{\mathrm{EOS}}\left(T_{i}, p_{i}\right)}, \tag{3.3}
\end{equation*}
$$

where $\rho_{\text {exp }}, \rho_{\text {EOS }}$ and $\rho_{\text {adjusted }}$ are the experimental density, the density calculated from the equation of state and the density adjusted to the average temperature and pressure $T_{\text {avg }}, p_{\text {avg }}$. $T_{i}$ and $p_{i}$ are the experimental temperature and pressure. The equations of Span et al (2000), Tegeler et al (1999) and Katti et al (1986) were used for nitrogen, argon and neon, respectively. The adjusted densities and the magnitudes of the adjustment (in parts per million) are given as the last two columns for each temperature in table 4. The average adjustment is 10.0 ppm and the maximum adjustment is 42.6 ppm , so the uncertainties in the equation of state will have negligible effect on the final results. For each set of ( $T_{\text {avg }}, p_{\text {avg }}$ ), the adjusted densities were averaged to calculate the density ratios.

Table 5 presents the results for all cases. Figure 3 shows the extrapolation in pressure for neon and the two replicates for argon at 429 K . Both gases extrapolate to nearly the same zero-pressure limit.

### 3.5. Uncertainty analysis

The numerical study presented above considered the effects of measurement uncertainties. A more detailed analysis of uncertainties is presented here for the actual experimental measurements. The ODRPACK software used in fitting the


Figure 3. Extrapolation to zero pressure of the experimental density ratios at 429.749 K and 293.15 K for $(\bullet)$ neon and the two replicates $(+)$ and $(\times)$ for argon.
data provides an estimate of the uncertainties in the fitted parameters, but this is based only on the actual scatter in the data and deviations from the fitted model. Significant systematic errors are possible, and these could be highly reproducible and, thus, 'invisible' to the fitting software. These include calibration uncertainties in the thermometer, piston gauge and sinker volumes. The weighings at higher temperatures are observed to have more scatter; there may be additional systematic effects on the balance, either due to temperature effects or because the measurements at $T_{\text {ref }}$ and $T_{x}$ are separated in time. The working gas is assumed to be the same for all tests, but its purity may vary with time; this is simulated by varying the molar mass. (As argued above, a pure gas is not required and 'purity' is better interpreted as 'constancy'.)

Table 6 lists these uncertainty sources, their estimated magnitude ( $k=1$ ) for the present densimeter and their effect on the calculated density ratio at zero pressure. While these error sources and their magnitudes can be reasonably estimated, calculating their effects on the final extrapolation to zero pressure would be very complex. Thus, these effects were simulated by varying the measured data by the indicated magnitude for each of the error sources. The modified data were then fitted again to obtain the difference in the final

Table 3. Experimental $p-\rho-T$ data for neon with a comparison to the equation of state of Katti et al (1986). Densities are computed using sinker volumes based on the coefficient of thermal expansion ( $\rho_{\mathrm{CTE}}$ ) and using the 'corrected' sinker volumes discussed in section 3.7
( $\rho_{\text {corrected }}$ ).

| $T$ (K) | $p$ (MPa) | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\mathrm{EOS}}\right) / \rho_{\mathrm{EOS}}(\mathrm{ppm})$ |
| :---: | :---: | :---: | :---: | :---: |
| $T=293.15 \mathrm{~K} ; 28-29$ October 2004 |  |  |  |  |
| 293.1509 | 6.993078 | 56.0230 | 56.0230 | 408.2 |
| 293.1497 | 6.993077 | 56.0235 | 56.0235 | 412.0 |
| 293.1521 | 6.993072 | 56.0233 | 56.0233 | 417.3 |
| 293.1534 | 6.993065 | 56.0229 | 56.0229 | 417.1 |
| 293.1509 | 6.993056 | 56.0233 | 56.0233 | 416.7 |
| 293.1520 | 4.994988 | 40.3958 | 40.3958 | 375.3 |
| 293.1509 | 4.994985 | 40.3956 | 40.3956 | 367.4 |
| 293.1515 | 4.994982 | 40.3956 | 40.3956 | 368.8 |
| 293.1533 | 4.994979 | 40.3954 | 40.3954 | 371.6 |
| 293.1521 | 4.994977 | 40.3956 | 40.3956 | 371.8 |
| 293.1507 | 3.995933 | 32.4687 | 32.4687 | 329.0 |
| 293.1513 | 3.995937 | 32.4686 | 32.4686 | 327.1 |
| 293.1541 | 3.995940 | 32.4687 | 32.4687 | 338.6 |
| 293.1528 | 3.995943 | 32.4687 | 32.4687 | 335.0 |
| 293.1510 | 3.995946 | 32.4688 | 32.4688 | 330.2 |
| 293.1498 | 2.996941 | 24.4668 | 24.4668 | 294.8 |
| 293.1522 | 2.996939 | 24.4666 | 24.4666 | 296.8 |
| 293.1534 | 2.996937 | 24.4664 | 24.4664 | 293.3 |
| 293.1511 | 2.996936 | 24.4667 | 24.4667 | 298.5 |
| 293.1512 | 2.996934 | 24.4669 | 24.4669 | 305.6 |
| 293.1497 | 1.997947 | 16.3880 | 16.3880 | 241.8 |
| 293.1525 | 1.997947 | 16.3881 | 16.3881 | 251.9 |
| 293.1537 | 1.997947 | 16.3879 | 16.3879 | 249.3 |
| 293.1503 | 1.997947 | 16.3881 | 16.3881 | 247.5 |
| 293.1524 | 1.997946 | 16.3881 | 16.3881 | 252.4 |
| 293.1515 | 0.998978 | 8.2326 | 8.2326 | 175.9 |
| 293.1503 | 0.998978 | 8.2327 | 8.2327 | 178.1 |
| 293.1526 | 0.998978 | 8.2327 | 8.2327 | 185.6 |
| 293.1531 | 0.998978 | 8.2328 | 8.2328 | 200.1 |
| 293.1503 | 0.998978 | 8.2328 | 8.2328 | 196.7 |
| $T=429.75 \mathrm{~K} ; 4-5$ November 2004 |  |  |  |  |
| 429.7493 | 6.992942 | 38.5084 | 38.5021 | -35.2 |
| 429.7497 | 6.992937 | 38.5084 | 38.5021 | -32.9 |
| 429.7521 | 6.992937 | 38.5081 | 38.5018 | -35.7 |
| 429.7527 | 6.992930 | 38.5082 | 38.5019 | -30.7 |
| 429.7504 | 6.992930 | 38.5082 | 38.5019 | -36.2 |
| 429.7513 | 4.994930 | 27.7061 | 27.7016 | 27.9 |
| 429.7521 | 4.994929 | 27.7061 | 27.7015 | 27.7 |
| 429.7509 | 4.994928 | 27.7060 | 27.7015 | 23.9 |
| 429.7497 | 4.994927 | 27.7061 | 27.7016 | 25.9 |
| 429.7511 | 4.994927 | 27.7060 | 27.7015 | 24.3 |
| 429.7491 | 3.995957 | 22.2460 | 22.2424 | 59.9 |
| 429.7490 | 3.995958 | 22.2459 | 22.2422 | 53.3 |
| 429.7480 | 3.995958 | 22.2458 | 22.2422 | 48.9 |
| 429.7484 | 3.995957 | 22.2459 | 22.2422 | 52.4 |
| 429.7496 | 3.995957 | 22.2459 | 22.2423 | 57.1 |
| 429.7488 | 2.996951 | 16.7457 | 16.7429 | 99.3 |
| 429.7478 | 2.996949 | 16.7456 | 16.7429 | 95.3 |
| 429.7483 | 2.996947 | 16.7453 | 16.7426 | 78.5 |
| 429.7500 | 2.996945 | 16.7453 | 16.7426 | 84.8 |
| 429.7491 | 2.996944 | 16.7451 | 16.7424 | 71.7 |
| 429.7492 | 1.997949 | 11.2045 | 11.2027 | 118.5 |
| 429.7476 | 1.997948 | 11.2047 | 11.2028 | 129.4 |
| 429.7484 | 1.997947 | 11.2045 | 11.2026 | 111.8 |
| 429.7500 | 1.997946 | 11.2044 | 11.2025 | 107.7 |
| 429.7493 | 1.997945 | 11.2042 | 11.2023 | 88.4 |

Table 3. (Continued.)

| $T(\mathrm{~K})$ | $p(\mathrm{MPa})$ | $\rho_{\text {CTE }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\rho_{\text {corrected }}\left(\mathrm{kg} \mathrm{m}^{-3}\right)$ | $\left(\rho_{\text {corrected }}-\rho_{\text {EOS }}\right) / \rho_{\text {EOS }}(\mathrm{ppm})$ |
| :--- | :--- | :--- | :--- | :--- |
| 429.7508 | 0.998974 | 5.6228 | 5.6219 | 141.6 |
| 429.7497 | 0.998973 | 5.6226 | 5.6217 | 103.8 |
| 429.7490 | 0.998973 | 5.6226 | 5.6216 | 93.3 |
| 429.7505 | 0.998973 | 5.6226 | 5.6217 | 97.8 |
| 429.7512 | 0.998972 | 5.6226 | 5.6217 | 108.7 |

Table 4. Example for data reduction: argon at $T_{\text {ref }}=293.15 \mathrm{~K}$ and $T_{x}=429.75 \mathrm{~K}$ showing experimental measurements, average temperatures and pressures, densities adjusted to average $(T, p)$ and resulting density ratios.

result. All the differences were summed (by the root sum of squares) together with the uncertainty resulting from the actual experimental scatter to arrive at the overall uncertainty.

Uncertainties in the weighings are seen to have the largest impact. This indicates where effort should be concentrated
to improve the DRGT technique. The effect of experimental scatter is next in importance. The effect of sample purity is relatively small. Any zero offset in the pressure, such as an incorrect reference pressure for the piston gauge, has a significant effect. It is also interesting to note factors which

Table 5. Fitted $T_{x}$ and experimental density ratios extrapolated to zero pressure and relative deviations from the measured temperature ratio.

| $T_{\text {ref }}$ measured <br> on ITS-90 | $T_{x}$ measured <br> on ITS-90 | $T_{x}$ fitted <br> by (2.3) | $\operatorname{Lim}(p \rightarrow 0) \rho_{\mathrm{x}} / \rho_{\text {ref }}$ |
| :--- | :--- | :--- | :--- | :--- | | $10^{6}\left[\rho_{x} / \rho_{\text {ref }}-T_{\text {ref }} / T_{x}\right] /$ |
| :--- |
| $\left(T_{\text {ref }} / T_{x}\right)$ |

have virtually no effect on the result. A constant relative error in the measured pressures has no effect since the measurements at $T_{x}$ and $T_{\text {ref }}$ are carried out at the same pressures. In other words, a piston gauge area which is in error will not affect the final extrapolation. (Of course, this assumes that the same piston gauge is used for all of the measurements.) Likewise, an error in the sinker volumes at the reference temperature has no effect. Density ratios are the important quantity, and any error in the sinker volumes at $T_{\text {ref }}$ would affect the calculated density at $T_{x}$ and $T_{\text {ref }}$ the same and thus cancel in taking the ratio.

### 3.6. Force transmission error and the effect of magnetic fluids

The magnetic suspension coupling in the densimeter is influenced by the magnetic properties of the materials of construction and by any external magnetic fields. These give rise to a force transmission error (FTE). The magnetic properties of the fluid under study also affect the measurement. The fluids used here are slightly diamagnetic with magnetic susceptibilities of $-6.1 \times 10^{-9} \mathrm{~m}^{3} \mathrm{~kg}^{-1},-4.3 \times 10^{-9} \mathrm{~m}^{3} \mathrm{~kg}^{-1}$ and $-5.4 \times 10^{-9} \mathrm{~m}^{3} \mathrm{~kg}^{-1}$ for argon, neon and nitrogen, respectively (Lide 2004). For the two-sinker densimeter used here, the differential nature of the measurement largely cancels these effects, but small errors in density-on the order of 20 ppm -are experienced. With a two-sinker densimeter, the force transmission error is nearly the same for the corresponding weighings at the two temperatures so it nearly cancels. The present two-sinker densimeter incorporates two balance calibration masses which are weighed in addition to the two sinkers for each density determination. An analysis developed by McLinden et al (2006) makes use of these additional weighings to determine and compensate for the force transmission error and the effects of magnetic fluids. With this analysis the uncertainty associated with magnetic effects is reduced to about 2 ppm in density.

### 3.7. In situ determination of sinker volumes

The final column in table 5 gives the relative difference (in ppm ) between the measured temperature ratio and the extrapolated density ratio; the same data are shown in figure 4.


Figure 4. Relative difference between the extrapolated density ratio and the measured temperature ratio for $(\mathbf{\Delta})$ argon, $(\boldsymbol{\bullet})$ neon and ( $\boldsymbol{\square}$ ) nitrogen; the solid line is a polynomial fit of the points. This difference is interpreted as either the error in the measured temperature relative to the thermodynamic temperature or the error in the sinker volumes. Error bars are for $k=1$.

The density ratio at zero pressure should be equal to the ratio of the thermodynamic temperatures. If the densimeter were operating as a gas thermometer this difference would indicate the difference between the temperature (measured on ITS-90) and the thermodynamic temperature. However, as explained above, the uncertainties in the sinker volumes for the present densimeter are much greater than those in the temperature: as much as 125 ppm in volume at 505 K versus 2 mK in temperature (equivalent to $4-8 \mathrm{ppm}$ over the temperature range of 234-505 K). Clearly, the present instrument cannot be used as a thermometer. But, by inverting the method, the sinker volumes can be calibrated in situ. The sinker volumes are adjusted by the difference between the measured temperature ratio and the extrapolated density ratio. Figure 4 depicts the correction applied to the sinker volumes compared to the volumes computed from the thermal expansion data. The experimental points are fitted to a cubic polynomial (shown by the solid line in figure 4), and the sinker volumes are multiplied by this correction for use in (3.1). The resulting sinker volumes, and thus also the measured densities, have an

Table 6. Effects of experimental uncertainties and systematic errors on the density ratios extrapolated to zero pressure.

| Source | Magnitude ( $k=1$ ) | Effect on $\lim (p \rightarrow 0)\left(\rho_{\mathrm{x}} / \rho_{\text {ref }}\right)(\mathrm{ppm})$ at $T_{x}(\mathrm{~K})$ |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Argon |  |  |  |  |  | $\begin{aligned} & \text { Neon } \\ & 430 \end{aligned}$ | Nitrogen |  |  |  |
|  |  | 234 | 273 | 360 | 430 (no. 1) | 430 (no. 2) | 505 |  | 340 | 400 | 440 | 480 |
| Fit to model | (Scatter in data) | 12.9 | 11.6 | 4.7 | 9.2 | 5.2 | 41.0 | 16.4 | 6.0 | 19.8 | 16.8 | 42.2 |
| $p$ (zero offset) | 2 Pa | 8.1 | 8.3 | 8.5 | 8.6 | 8.6 | 5.7 | 5.3 | 8.7 | 8.8 | 8.8 | 8.8 |
| $p$ (const \% error) | 10 ppm | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| Sinker V @ $T_{\text {ref }}$ | 10 ppm | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| Weighings | $+1 \mu \mathrm{~g} @ T_{\text {ref }}-1 \mu \mathrm{~g} @ T_{x}$ | 21.9 | 24.3 | 28.9 | 32.3 | 32.3 | 23.7 | 39.7 | 41.0 | 45.3 | 48.0 | 50.7 |
| $T$-dependent weighing error | $0.02 \mu \mathrm{~g}\left(T_{x}-T_{\text {ref }}\right)$ | 5.7 | 2.3 | 10.7 | 33.3 | 26.3 | 31.8 | 32.2 | 20.8 | 40.4 | 56.1 | 74.0 |
| Molar mass (sample purity) | +1 ppm@ $T_{\text {ref }}-1 \mathrm{ppm} @ T_{x}$ | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 | 2.0 |
| $T_{\text {ref }}$ | $2 \mathrm{mK} @ T_{\text {ref }}$ | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 | 6.8 |
| Root sum of squares |  | 28.1 | 29.1 | 33.1 | 48.6 | 43.4 | 57.8 | 54.4 | 47.8 | 64.9 | 76.6 | 99.8 |

estimated standard uncertainty of 60 ppm at 505 K or one-half the uncertainty of the volumes computed from the thermal expansion data.

The corrected sinker volumes are based on ITS-90 temperatures. No attempt was made to include any errors in ITS-90. The suspected error of 11 mK at 505 K (Strouse et al 2002, Ripple et al 2002) is 22 ppm , and it is lower at the other temperatures investigated here. Any error in ITS90 is significantly less than the uncertainty in the present measurements.

## 4. Discussion-feasibility of the method

This work has shown that a densimeter can, in principle, be used as a density-ratio gas thermometer. The necessary experiments have been demonstrated, and the data analysis techniques have been developed. The densimeter measurements typically required two or three days for each $T_{\text {ref }}$ or $T_{x}$, comparing very favourably with other methods. The main question remaining is whether the uncertainties can be reduced to a level comparable to other methods. Edsinger and Schooley (1989) report typical uncertainties for a CVGT of 5 mK at 505 K and 15 mK at 933 K . Uncertainties $(k=1)$ for acoustic thermometry range from 0.6 mK at 303 K to 3.0 mK at 505 K (Strouse et al 2002). Expressed as relative uncertainties, these range from 2 ppm for the acoustic thermometer at 303 K to 16 ppm for the CVGT at 933 K .

The numerical study indicated that the uncertainties in temperature for the DRGT are on the order of $1-5 \mathrm{mK}$ for reasonable values of experimental uncertainties. A more detailed analysis of the present experimental results, including consideration of likely systematic uncertainties, resulted in combined uncertainties on the order of 50 ppm in the density ratio or roughly 25 mK in temperature at 505 K . Can the uncertainties in the present densimeter be reduced by an order of magnitude in an instrument designed explicitly as a gas thermometer?

The largest uncertainty in the present densimeter is in the sinker volumes. They are so large, in fact, that the present data were analysed to determine the sinker volumes assuming the temperatures were known, rather than vice versa. A gas thermometer would require very accurate sinker volumes, and these could be determined by direct measurement of the sinker dimensions by interferometry over the temperature range of interest. Interferometric systems can measure lengths to sub-nanometre resolution or coefficients of thermal expansion (CTE) with uncertainties of $10^{-8} \mathrm{~K}^{-1}$ or less. One commercial testing company claims $10^{-8} \mathrm{~K}^{-1}$ uncertainty in CTE for temperatures from 20 K to 1300 K using a Michelson interferometer according to the ASTM test method E289 (ASTM 2004). At 505 K this would correspond to 7 ppm uncertainty in volume. An academic group (Dudik et al 2003) has developed a system operating from 30 K to 310 K with uncertainties in CTE of $2 \times 10^{-9} \mathrm{~K}^{-1}$; they thought that their system could be extended to higher temperatures as well (Halverson 2003).

More accurate weighings will be required. This implies either a more sensitive balance with sinkers of the current size or larger sinkers weighed to microgram uncertainties. The present 60 g sinkers are made of titanium
and tantalum for a volume difference of $9.74 \mathrm{~cm}^{3}$. The volume difference could be more than doubled to $22.96 \mathrm{~cm}^{3}$ with silicon and platinum sinkers of the same 60 g mass. Commercial balances can directly weigh objects up to 310 g with a precision of $1 \mu \mathrm{~g}$. Special 'kilogram comparator' balances weigh 1 kg objects with $1 \mu \mathrm{~g}$ uncertainty ( 1 part in $10^{9}$ ). However, these balances have a very limited electronic weighing range and would require tare weights to compensate for the different buoyancy forces on the sinkers. Weighings of the required sensitivity are thus feasible using larger sinkers. One of the largest uncertainties in a high-accuracy mass determination arises from uncertainties in the air density in the weighing chamber. The present densimeter has calibration masses inside the balance chamber which are weighed as part of each density determination; these masses have nearly identical volumes so that air buoyancy effects on the balance calibration cancel. The present 1 ppm uncertainty in the masses of the sinkers and calibration masses could be easily improved.

The uncertainties in temperature will need to improve to the 0.5 mK level for $T_{\text {ref }}$ near 273.16 K , and this is routinely achieved by temperature metrologists. The resistance of the PRT at $T_{x}$ will need to be measured with uncertainties corresponding to 0.5 mK in temperature. Just as important will be a very stable thermostat with small temperature gradients. A stability of 0.5 mK or better should be readily achievable with careful design of the thermostat.

The measurement of pressure presents no particular difficulties. Commercially available piston gauges would be suitable. Care will be required, however, to minimize any zero offset, for example as the result of errors in the reference pressure for the piston.

The use of a magnetic suspension coupling brings with it effects due to magnetic materials and diamagnetic fluids but these can be compensated for (McLinden et al 2006). Glos et al (2004) describe a two-sinker densimeter where the balance is 'immersed' in the test fluid. Such an approach avoids the need for a magnetic suspension coupling, but brings with it many other complications.

The improvements necessary to realize a density-ratio gas thermometer with uncertainties comparable to acoustic or constant-volume gas thermometers would require state-of-the-art temperature, mass and dimensional metrology, but all are feasible with present technologies.

## 5. Conclusions

A new technique for the determination of thermodynamic temperature based on gas density measurements has been presented and shown to be feasible. Further development of this method into a temperature standard would require a significant effort involving a collaboration among temperature, pressure, dimensional and mass metrologists, but no insurmountable obstacles are foreseen.

The method was inverted and used to determine in situ the volume of the densimeter sinkers for the case where the relative uncertainty in the sinker volumes was substantially higher than that in temperature.

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